# Heavy Bosons Other Than Higgs Bosons, Searches for

We list here various limits on charged and neutral heavy vector bosons (other than W's and Z's), heavy scalar bosons (other than Higgs bosons), vector or scalar leptoquarks, and axigluons.

### THE W' SEARCHES

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Any electrically charged gauge boson outside of the Standard Model is generically denoted W'. A W' always couples to two different flavors of fermions, similar to the W boson. In particular, if a W' couples quarks to leptons it is a leptoquark gauge boson.

The most attractive candidate for W' is the  $W_R$  gauge boson associated with the left-right symmetric models [1]. These models seek to provide a spontaneous origin for parity violation in weak interactions. Here the gauge group is extended to  $\mathrm{SU}(3)_C\times\mathrm{SU}(2)_L\times\mathrm{SU}(2)_R\times\mathrm{U}(1)_{B-L}$  with the Standard Model hypercharge identified as  $Y = T_{3R} + (B-L)/2$ ,  $T_{3R}$  being the third component of  $SU(2)_R$ . The fermions transform under the gauge group in a left-right symmetric fashion:  $q_L(3, 2, 1, 1/3) +$  $q_R(3,1,2,1/3)$  for quarks and  $\ell_L(1,2,1,-1) + \ell_R(1,1,2,-1)$ for leptons. Note that the model requires the introduction of right-handed neutrinos, which can facilitate the see-saw mechanism for explaining the smallness of the ordinary neutrino masses. A Higgs bidoublet  $\Phi(1,2,2,0)$  is usually employed to generate quark and lepton masses and to participate in the electroweak symmetry breaking. Under left-right (or parity) symmetry,  $q_L \leftrightarrow q_R$ ,  $\ell_L \leftrightarrow \ell_R$ ,  $W_L \leftrightarrow W_R$  and  $\Phi \leftrightarrow \Phi^{\dagger}$ .

After spontaneous symmetry breaking, the two W bosons of the model,  $W_L$  and  $W_R$ , will mix. The physical mass eigenstates are denoted as

$$W_1 = \cos \zeta W_L + \sin \zeta W_R, \qquad W_2 = -\sin \zeta W_L + \cos \zeta W_R \quad (1)$$

with  $W_1$  identified as the observed W boson. The most general Lagrangian that describes the interactions of the  $W_{1,2}$  with the quarks can be written as [2]

$$\mathcal{L} = -\frac{1}{\sqrt{2}} \overline{u} \gamma_{\mu} \left[ \left( g_L \cos \zeta \, V^L P_L - g_R e^{i\omega} \sin \zeta \, V^R P_R \right) W_1^{\mu} + \left( g_L \sin \zeta \, V^L P_L + g_R e^{i\omega} \cos \zeta \, V^R P_R \right) W_2^{\mu} \right] d + h.c.(2)$$

where  $g_{L,R}$  are the SU(2)<sub>L,R</sub> gauge couplings,  $P_{L,R} = (1 \mp \gamma_5)/2$ and  $V^{L,R}$  are the left- and right-handed CKM matrices in the quark sector. The phase  $\omega$  reflects a possible complex mixing parameter in the  $W_L$ – $W_R$  mass-squared matrix. Note that there is CP violation in the model arising from the right-handed currents even with only two generations. The Lagrangian for leptons is identical to that for quarks, with the replacements  $u \rightarrow \nu$ ,  $d \rightarrow e$  and the identification of  $V^{L,R}$  with the CKM matrices in the leptonic sector.

If parity invariance is imposed on the Lagrangian, then  $g_L = g_R$ . Furthermore, the Yukawa coupling matrices that arise from coupling to the Higgs bidoublet  $\Phi$  will be Hermitian. If in addition the vacuum expectation values of  $\Phi$  are assumed to be real, the quark and lepton mass matrices will also be Hermitian, leading to the relation  $V^L = V^R$ . Such models are called manifest left-right symmetric models and are approximately realized with a minimal Higgs sector [3]. If instead parity and CP are both imposed on the Lagrangian, then the Yukawa coupling matrices will be real symmetric and, after spontaneous

CP violation, the mass matrices will be complex symmetric. In this case, which is known in the literature as *pseudo-manifest* left-right symmetry,  $V^L = (V^R)^*$ .

*Indirect constraints:* In minimal version of manifest or pseudo-manifest left-right symmetric models with  $\omega = 0$  or  $\pi$ , there are only two free parameters,  $\zeta$  and  $M_{W_2}$ , and they can be constrained from low energy processes. In the large  $M_{W_2}$  limit, stringent bounds on the angle  $\zeta$  arise from three processes. (i) Nonleptonic K decays: The decays  $K \to 3\pi$  and  $K \rightarrow 2\pi$  are sensitive to small admixtures of right-handed currents. Assuming the validity of PCAC relations in the Standard Model it has been argued in Ref. 4 that the success in the  $K \to 3\pi$  prediction will be spoiled unless  $|\zeta| \leq 4 \times 10^{-3}$ . (ii)  $b \to s\gamma$ : The amplitude for this process has an enhancement factor  $m_t/m_b$  relative to the Standard Model and thus can be used to constrain  $\zeta$  yielding the limit  $-0.01 \leq \zeta \leq 0.003$  [5]. (iii) Universality in weak decays: If the right-handed neutrinos are heavy, the right-handed admixture in the charged current will contribute to  $\beta$  decay and K decay, but not to the  $\mu$ decay. This will modify the extracted values of  $V_{ud}^L$  and  $V_{us}^L$ . Demanding that the difference not upset the three generation unitarity of the CKM matrix, a bound  $|\zeta| \leq 10^{-3}$  has been derived [6].

If the  $\nu_R$  are heavy, leptonic and semileptonic processes do not constrain  $\zeta$  since the emission of  $\nu_R$  will not be kinematically allowed. However, if the  $\nu_R$  is light enough to be emitted in  $\mu$  decay and  $\beta$  decay, stringent limits on  $\zeta$  do arise. For example,  $|\zeta| \leq 0.039$  can be obtained from polarized  $\mu$  decay [7] in the large  $M_{W_2}$  limit of the manifest left-right model. Alternatively, in the  $\zeta = 0$  limit, there is a constraint  $M_{W_2} \geq 484$  GeV from direct  $W_2$  exchange. For the constraint on the case in

which  $M_{W_2}$  is not taken to be heavy, see Ref. 2. There are also cosmological and astrophysical constraints on  $M_{W_2}$  and  $\zeta$  in scenarios with a light  $\nu_R$ . During nucleosynthesis the process  $e^+e^- \rightarrow \nu_R \overline{\nu}_R$ , proceeding via  $W_2$  exchange, will keep the  $\nu_R$  in equilibrium leading to an overproduction of <sup>4</sup>He unless  $M_{W_2}$  is greater than about 1 TeV [8]. Likewise the  $\nu_{eR}$ produced via  $e_R^- p \rightarrow n\nu_R$  inside a supernova must not drain too much of its energy, leading to limits  $M_{W_2} > 16$  TeV and  $|\zeta| \leq 3 \times 10^{-5}$  [9]. Note that models with light  $\nu_R$  do not have a see-saw mechanism for explaining the smallness of the neutrino masses, though other mechanisms may arise in variant models [10].

The mass of  $W_2$  is severely constrained (independent of the value of  $\zeta$ ) from  $K_L$ – $K_S$  mass-splitting. The box diagram with exchange of one  $W_L$  and one  $W_R$  has an anomalous enhancement and yields the bound  $M_{W_2} \geq 1.6$  TeV [11] for the case of manifest or pseudo-manifest left-right symmetry. If the  $\nu_R$  have Majorana masses, another constraint arises from neutrinoless double  $\beta$  decay. Combining the experimental limit from <sup>76</sup>Ge decay with arguments of vacuum stability, a limit of  $M_{W_2} \geq 1.1$  TeV has been obtained [12].

**Direct search limits:** Limits on  $M_{W_2}$  from direct searches depend on the available decay channels of  $W_2$ . If  $\nu_R$  is heavier than  $W_2$ , the decay  $W_2^+ \to \ell_R^+ \nu_R$  will be forbidden kinematically. Assuming that  $\zeta$  is small, the dominant decay of  $W_2$ will be into dijets. UA2 [13] has excluded a  $W_2$  in the mass range of 100 to 251 GeV in this channel. DØ excludes the mass range of 340 to 680 GeV [14], while CDF excludes the mass range of 300 to 420 GeV for such a  $W_2$  [15]. If  $\nu_R$  is lighter than  $W_2$ , the decay  $W_2^+ \to e_R^+ \nu_R$  is allowed. The  $\nu_R$ can then decay into  $e_R W_R^*$ , leading to an eejj signature. DØ

has a limit of  $M_{W_2} > 720$  GeV if  $m_{\nu_R} \ll M_{W_2}$ ; the bound weakens, for example, to 650 GeV for  $m_{\nu_R} = M_{W_2}/2$  [16]. CDF finds  $M_{W_2} > 652$  GeV if  $\nu_R$  is stable and much lighter than  $W_2$  [17]. All of these limits assume manifest or pseudo-manifest left-right symmetry. See [16] for some variations in the limits if the assumption of left-right symmetry is relaxed.

Alternative models: W' gauge bosons can also arise in other models. We shall briefly mention some such popular models, but for details we refer the reader to the original literature. The *alternate* left-right model [18] is based on the same gauge group as the left-right model, but arises in the following way: In  $E_6$  unification, there is an option to identify the righthanded down quarks as  $SU(2)_R$  singlets or doublets. If they are  $SU(2)_R$  doublets, one recovers the conventional left-right model; if they are singlets it leads to the alternate left-right model. A similar ambiguity exists in the assignment of lefthanded leptons; the alternate left-right model assigns them to a (1, 2, 2, 0) multiplet. As a consequence, the ordinary neutrino remains exactly massless in the model. One important difference from the usual left-right model is that the limit from the  $K_L-K_S$ mass difference is no longer applicable, since the  $d_R$  do not couple to the  $W_R$ . There is also no limit from polarized  $\mu$  decay, since the  $SU(2)_R$  partner of  $e_R$  can receive a large Majorana mass. Other W' models include the un-unified Standard Model of Ref. 19 where there are two different SU(2) gauge groups, one each for the quarks and leptons; models with separate SU(2) gauge factors for each generation [20]; and the SU(3)<sub>C</sub> ×  $SU(3)_L \times U(1)$  model of Ref. 21.

Leptoquark gauge bosons: The  $SU(3)_C \times U(1)_{B-L}$  part of the gauge symmetry discussed above can be embedded into a simple  $SU(4)_C$  gauge group [22]. The model then will contain

leptoquark gauge boson as well, with couplings of the type  $\{(\overline{e}_L\gamma_\mu d_L + \overline{\nu}_L\gamma_\mu u_L)W'^\mu + (L \to R)\}$ . The best limit on such leptoquark W' comes from nonobservation of  $K_L \to \mu e$ , which requires  $M_{W'} \geq 1400$  TeV; for the corresponding limits on less conventional leptoquark flavor structures, see Ref. 23. Thus such a W' is inaccessible to direct searches with present machines which are sensitive to vector leptoquark masses of order 300 GeV only.

## References

- J.C. Pati and A. Salam, Phys. Rev. D10, 275 (1974);
   R.N. Mohapatra and J.C. Pati, Phys. Rev. D11, 566 (1975); *ibid.* Phys. Rev. D11, 2558 (1975);
   G. Senjanovic and R.N. Mohapatra, Phys. Rev. D12, 1502 (1975).
- P. Langacker and S. Uma Sankar, Phys. Rev. D40, 1569 (1989).
- A. Masiero, R.N. Mohapatra, and R. Peccei, Nucl. Phys. B192, 66 (1981);
   J. Basecq, et al., Nucl. Phys. B272, 145 (1986).
- J. Donoghue and B. Holstein, Phys. Lett. **113B**, 383 (1982).
- 5. K.S. Babu, K. Fujikawa, and A. Yamada, Phys. Lett. B333, 196 (1994);
  P. Cho and M. Misiak, Phys. Rev. D49, 5894 (1994);
  T.G. Rizzo, Phys. Rev. D50, 3303 (1994).
- 6. L. Wolfenstein, Phys. Rev. **D29**, 2130 (1984).
- 7. P. Herczeg, Phys. Rev. **D34**, 3449 (1986).
- G. Steigman, K.A. Olive, and D. Schramm, Nucl. Phys. B180, 497 (1981).
- 9. R. Barbieri and R.N. Mohapatra, Phys. Rev. D39, 1229 (1989);
  G. Raffelt and D. Seckel, Phys. Rev. Lett. 60, 1793 (1988).
- D. Chang and R.N. Mohapatra, Phys. Rev. Lett. 58, 1600 (1987);

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K.S. Babu and X.G. He, Mod. Phys. Lett. A4, 61 (1989).

- G. Beall, M. Bender, and A. Soni, Phys. Rev. Lett. 48, 848 (1982).
- 12. R.N. Mohapatra, Phys. Rev. **D34**, 909 (1986).
- J. Alitti, et al. (UA2 Collaboration), Nucl. Phys. B400, 3 (1993).
- B. Abbott, et al. (DØ Collaboration), International Europhysics Conference on High Energy Physics, August 19-26, 1997, Jerusalem, Israel.
- F. Abe, *et al.* (CDF Collaboration), Phys. Rev. **D55**, R5263 (1997).
- S. Abachi, *et al.* (DØ Collaboration), Phys. Rev. Lett. **76**, 3271 (1996).
- 17. F. Abe, *et al.* (CDF Collaboration), Phys. Rev. Lett. **74**, 2900 (1995).
- 18. E. Ma, Phys. Rev. D36, 274 (1987);
  K.S. Babu, X-G. He and E. Ma, Phys. Rev. D36, 878 (1987).
- H. Georgi and E. Jenkins, Phys. Rev. Lett. 62, 2789 (1989);
  Nucl. Phys. B331, 541 (1990).
- 20. X. Li and E. Ma, Phys. Rev. Lett. 47, 1788 (1981);
  R.S. Chivukula, E.H. Simmons, and J. Terning, Phys. Lett. B331, 383 (1994);
  D.J. Muller and S. Nandi, Phys. Lett. B383, 345 (1996).
- 21. F. Pisano, V. Pleitez, Phys. Rev. D46, 410 (1992);
  P. Frampton, Phys. Rev. Lett. 69, 2889 (1992).
- 22. J.C. Pati and A. Salam, Phys. Rev. **D10**, 275 (1974).
- 23. A. Kuznetsov and N. Mikheev, Phys. Lett. B329, 295 (1994);
  G. Valencia and S. Willenbrock, Phys. Rev. D50, 6843 (1994).

# MASS LIMITS for W' (Heavy Charged Vector Boson Other Than W) in Hadron Collider Experiments

Couplings of W' to quarks and leptons are taken to be identical with those of W. The following limits are obtained from  $p\overline{p} \rightarrow W'X$  with W' decaying to the mode indicated in the comments. New decay channels (e.g.,  $W' \rightarrow WZ$ ) are assumed to be suppressed. UA1 and UA2 experiments assume that the  $t\overline{b}$  channel is not open.

VALUE (GeV)	CL%	DOCUMENT ID		TECN	COMMENT
>786	95	<sup>1</sup> AFFOLDER	011	CDF	$W'  ightarrow e  u,  \mu  u$
$\bullet$ $\bullet$ $\bullet$ We do not use the	following	data for averages	, fits,	, limits, e	etc. • • •
>660	95	<sup>2</sup> ABE	00	CDF	$W' \rightarrow \mu \nu$
none 300-420	95	<sup>3</sup> ABE	<b>97</b> G	CDF	$W'  ightarrow q \overline{q}$
>720	95	<sup>4</sup> ABACHI	<b>96</b> C	D0	W'  ightarrow e  u
>610	95	<sup>5</sup> ABACHI	95e	D0	W'  ightarrow e  u, $ au  u$
>652	95	<sup>6</sup> ABE	95M	CDF	W'  ightarrow e  u
>251	90	<sup>7</sup> ALITTI	93	UA2	$W' \rightarrow q \overline{q}$
none 260-600	95		93	RVUE	$W' \rightarrow q \overline{q}$
>220	90	<sup>9</sup> ALBAJAR	89	UA1	W'  ightarrow e  u
>209	90	<sup>10</sup> ANSARI	<b>87</b> D	UA2	W'  ightarrow e  u

<sup>1</sup>AFFOLDER 011 combine a new bound on  $W' \rightarrow e\nu$  of 754 GeV with the bound of ABE 00 on  $W' \rightarrow \mu\nu$  to obtain quoted bound.

<sup>2</sup>ABE 00 assume that the neutrino from W' decay is stable and has a mass significantly less than  $m_{W'}$ .

<sup>3</sup>ABE 97G search for new particle decaying to dijets.

<sup>4</sup> For bounds on  $W_R$  with nonzero right-handed mass, see Fig. 5 from ABACHI 96C.

<sup>5</sup>ABACHI 95E assume that the decay  $W' \rightarrow WZ$  is suppressed and that the neutrino from W' decay is stable and has a mass significantly less  $m_{W'}$ .

<sup>6</sup>ABE 95M assume that the decay  $W' \rightarrow WZ$  is suppressed and the (right-handed) neutrino is light, noninteracting, and stable. If  $m_{\nu}$ =60 GeV, for example, the effect on \_ the mass limit is neglibible.

<sup>7</sup> ALITTI 93 search for resonances in the two-jet invariant mass. The limit assumes  $\Gamma(W')/m_{W'} = \Gamma(W)/m_W$  and  $B(W' \rightarrow jj) = 2/3$ . This corresponds to  $W_R$  with  $m_{\nu_R} > m_{W_R}$  (no leptonic decay) and  $W_R \rightarrow t \overline{b}$  allowed. See their Fig. 4 for limits in the  $m_{W'}$ -B( $q \overline{q}$ ) plane.

<sup>8</sup> RIZZO 93 analyses CDF limit on possible two-jet resonances. The limit is sensitive to the inclusion of the assumed K factor.

<sup>9</sup> ALBAJAR 89 cross section limit at 630 GeV is  $\sigma(W')$  B( $e\nu$ ) < 4.1 pb (90% CL).

<sup>10</sup> See Fig. 5 of ANSARI 87D for the excluded region in the  $m_{W'}^{-}[(g_{W'q})^2 B(W' \rightarrow Q_{W'q})^2]$ 

 $e\overline{\nu}$ ] plane. Note that the quantity  $(g_{W'q})^2 B(W' \rightarrow e\overline{\nu})$  is normalized to unity for the standard W couplings.

#### W<sub>R</sub> (Right-Handed W Boson) MASS LIMITS

Assuming a light right-handed neutrino, except for BEALL 82, LANGACKER 89B, and COLANGELO 91.  $g_R = g_L$  assumed. [Limits in the section MASS LIMITS for W' below are also valid for  $W_R$  if  $m_{\nu_R} \ll m_{W_R}$ .] Some limits assume manifest left-right symmetry, *i.e.*, the equality of left- and right Cabibbo-Kobayashi-Maskawa matrices. For a comprehensive review, see LANGACKER 89B. Limits on the  $W_L$ - $W_R$  mixing angle  $\zeta$  are found in the next section. Values in brackets are from cosmological and astrophysical considerations and assume a light right-handed neutrino.

VALUE (GeV)	CL%	DOCUMENT ID		TECN	COMMENT	
> 715	90	<sup>11</sup> CZAKON	99	RVUE	Electroweak	
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• • • We do not use the following data for averages, fits, limits, etc. • • •

		B aata ter areia	600,,		
> 310	90	<sup>12</sup> THOMAS	01 CN	TR $\beta^+$	decay
> 137	95	<sup>13</sup> ACKERSTAFF			
>1400	68	<sup>14</sup> BARENBOIM	98 RV	UE Eleo	ctroweak, Z-Z′ mixing
> 549	68	<sup>15</sup> BARENBOIM	97 RV	UE $\mu$ d	ecay
> 220	95	<sup>16</sup> STAHL	97 RV	UE $\tau$ d	ecay
> 220	90	<sup>17</sup> ALLET	96 CN	TR $\beta^+$	decay
> 281	90	<sup>18</sup> KUZNETSOV	95 CN	TR Pol	arized neutron decay
> 282	90	<sup>19</sup> KUZNETSOV	94B CN	TR Pol	arized neutron decay
> 439	90	<sup>20</sup> BHATTACH	93 RV	UE Z-Z	Z' mixing
> 250	90	<sup>21</sup> SEVERIJNS			
		<sup>22</sup> IMAZATO			
> 475	90	<sup>23</sup> POLAK	928 RV	UE μd	ecay
> 240	90	<sup>24</sup> AQUINO	91 RV	UE Nei	itron decay
> 496	90		01 RV	ILE No.	itron and muon decay
> 700		<sup>25</sup> COLANGELO	91 TH	IEO <sup>m</sup> K	$\kappa_1^0 - m \kappa_s^0$
> 477	90	<sup>26</sup> POLAK	91 RV	UE $\mu$ d	ecay
[none 540–23000]		<sup>27</sup> BARBIERI	89B AS	TR SN	1987A; light $\nu_R$
> 300	90	<sup>28</sup> LANGACKER	89B RV	UE Ger	ieral
> 160	90	<sup>29</sup> BALKE	88 CN	TR $\mu$ –	$\rightarrow e \nu \overline{\nu}$
> 406	90	<sup>30</sup> JODIDIO	86 EL	EC Any	ιζ
> 482	90	<sup>30</sup> JODIDIO	86 EL	EC $\zeta =$	: 0
> 800		MOHAPATRA	86 RV	UE SU	$(2)_L \times SU(2)_R \times U(1)$
> 400	95	<sup>31</sup> STOKER	85 EL	EC Any	ις
> 475	95	<sup>31</sup> STOKER			
		<sup>32</sup> BERGSMA	83 CH	RM $\nu_{\mu}$	$e \rightarrow \mu \nu_e$
> 380	90	<sup>33</sup> CARR		EC $\mu^{+}$	
>1600		<sup>34</sup> BEALL	82 TH		$r_{0}^{0} - m_{K_{S}^{0}}$
[> 4000]					
[> 4000]		JIEIGWAN	19 00		cleosynthesis; light $ u_R$

 $^{11}$ CZAKON 99 perform a simultaneous fit to charged and neutral sectors.

<sup>12</sup> THOMAS 01 limit is from measurement of  $\beta^+$  polarization in decay of polarized <sup>12</sup>N. The listed limit assumes no mixing.

 $^{13}$  ACKERSTAFF 99D limit is from au decay parameters. Limit increase to 145 GeV for zero mixing.

<sup>14</sup> BARENBOIM 98 assumes minimal left-right model with Higgs of SU(2)<sub>R</sub> in SU(2)<sub>L</sub> doublet. For Higgs in SU(2)<sub>L</sub> triplet,  $m_{W_R} > 1100$  GeV. Bound calculated from effect of corresponding  $Z_{LR}$  on electroweak data through  $Z - Z_{LR}$  mixing.

- $^{15}$  The quoted limit is from  $\mu$  decay parameters. BARENBOIM 97 also evaluate limit from  $K_I - K_S$  mass difference.
- <sup>16</sup>STAHL 97 limit is from fit to  $\tau$ -decay parameters.
- <sup>17</sup>ALLET 96 measured polarization-asymmetry correlaton in  ${}^{12}N\beta^+$  decay. The listed limit assumes zero L-R mixing.

<sup>18</sup> KUZNETSOV 95 limit is from measurements of the asymmetry  $\langle \vec{p}_{\nu} \cdot \sigma_n \rangle$  in the  $\beta$  decay of polarized neutrons. Zero mixing assumed. See also KUZNETSOV 94B.

<sup>19</sup> KUZNETSOV 94B limit is from measurements of the asymmetry  $\langle \vec{p}_{\nu} \cdot \sigma_{n} \rangle$  in the  $\beta$  decay of polarized neutrons. Zero mixing assumed.

 $^{20}$  BHATTACHARYYA 93 uses Z-Z' mixing limit from LEP '90 data, assuming a specific Higgs sector of  $SU(2)_L \times SU(2)_R \times U(1)$  gauge model. The limit is for  $m_t = 200$  GeV and slightly improves for smaller  $m_t$ .

- $^{21}\,{\rm SEVERIJNS}$  93 measured polarization-asymmetry correlation in  $^{107}{\rm In}\,\beta^+$  decay. The listed limit assumes zero L-R mixing. Value quoted here is from SEVERIJNS 94 erratum.
- <sup>22</sup> IMAZATO 92 measure positron asymmetry in  $K^+ \rightarrow \mu^+ \nu_\mu$  decay and obtain

 $\xi P_{\mu} > 0.990$  (90%CL). If  $W_R$  couples to  $u\overline{s}$  with full weak strength ( $V_{us}^R = 1$ ), the result corresponds to  $m_{W_R} > 653$  GeV. See their Fig. 4 for  $m_{W_R}$  limits for general  $|V_{us}^R|^2 = 1 - |V_{ud}^R|^2$ .

- <sup>23</sup> POLAK 92B limit is from fit to muon decay parameters and is essentially determined by JODIDIO 86 data assuming  $\zeta$ =0. Supersedes POLAK 91.
- <sup>24</sup> AQUINO 91 limits obtained from neutron lifetime and asymmetries together with unitarity of the CKM matrix. Manifest left-right symmetry assumed. Stronger of the two limits also includes muon decay results.
- <sup>25</sup> COLANGELO 91 limit uses hadronic matrix elements evaluated by QCD sum rule and is less restrictive than BEALL 82 limit which uses vacuum saturation approximation. Manifest left-right symmetry assumed.
- <sup>26</sup> POLAK 91 limit is from fit to muon decay parameters and is essentially determined by JODIDIO 86 data assuming  $\zeta$ =0. Superseded by POLAK 92B.
- $^{27}$  BARBIERI 89B limit holds for  $m_{
  u_R} \leq$  10 MeV.
- <sup>28</sup> LANGACKER 89B limit is for any  $\nu_R$  mass (either Dirac or Majorana) and for a general class of right-handed quark mixing matrices.
- <sup>29</sup> BALKE 88 limit is for  $m_{\nu_{eR}} = 0$  and  $m_{\nu_{\mu R}} \leq 50$  MeV. Limits come from precise measurements of the muon decay asymmetry as a function of the positron energy.
- <sup>30</sup> JODIDIO 86 is the same TRIUMF experiment as STOKER 85 (and CARR 83); however, it uses a different technique. The results given here are combined results of the two techniques. The technique here involves precise measurement of the end-point  $e^+$ spectrum in the decay of the highly polarized  $\mu^+$ .
- <sup>31</sup> STOKER 85 is same TRIUMF experiment as CARR 83. Here they measure the decay  $e^+$  spectrum asymmetry above 46 MeV/c using a muon-spin-rotation technique. Assumed a light right-handed neutrino. Quoted limits are from combining with CARR 83.
- <sup>32</sup> BERGSMA 83 set limit  $m_{W_2}/m_{W_1}$  >1.9 at CL = 90%.
- <sup>33</sup> CARR 83 is TRIUMF experiment with a highly polarized  $\mu^+$  beam. Looked for deviation from V-A at the high momentum end of the decay  $e^+$  energy spectrum. Limit from previous world-average muon polarization parameter is  $m_{W_R}$  >240 GeV. Assumes a light right-handed neutrino.
- <sup>34</sup> BEALL 82 limit is obtained assuming that  $W_R$  contribution to  $K_L^0 K_S^0$  mass difference is smaller than the standard one, neglecting the top quark contributions. Manifest left-right symmetry assumed.

#### Limit on $W_L$ - $W_R$ Mixing Angle $\zeta$

		$W_1 = W_L \cos\zeta - W_R$ to cosmological and a			$\nu_R$ assumed unless noted. considerations.
VALUE	<u>CL%</u>	DOCUMENT ID		TECN	COMMENT
• • • We do not use the	e followir	ng data for averages	, fits	, limits,	etc. • • •
< 0.12	95	<sup>35</sup> ACKERSTAFF	<b>99</b> D	OPAL	au decay
< 0.013	90	<sup>36</sup> CZAKON			Electroweak
< 0.0333		<sup>37</sup> BARENBOIM	97	RVUE	$\mu$ decay
< 0.04	90	<sup>38</sup> MISHRA	92	CCFR	u N scattering
-0.0006 to $0.0028$	90	<sup>39</sup> AQUINO	91	RVUE	
[none 0.00001-0.02]		<sup>40</sup> BARBIERI	<b>89</b> B	ASTR	SN 1987A
< 0.040	90	<sup>41</sup> JODIDIO	86	ELEC	$\mu$ decay
-0.056 to 0.040	90	<sup>41</sup> JODIDIO	86	ELEC	$\mu$ decay
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 $^{35}$  ACKERSTAFF 99D limit is from au decay parameters.

- <sup>36</sup>CZAKON 99 perform a simultaneous fit to charged and neutral sectors.
- <sup>37</sup> The quoted limit is from  $\mu$  decay parameters. BARENBOIM 97 also evaluate limit from  $K_L$ - $K_S$  mass difference.
- <sup>38</sup> MISHRA 92 limit is from the absence of extra large-*x*, large- $y \overline{\nu}_{\mu} N \rightarrow \overline{\nu}_{\mu} X$  events at Tevatron, assuming left-handed  $\nu$  and right-handed  $\overline{\nu}$  in the neutrino beam. The result gives  $\zeta^2(1-2m_{W_1}^2/m_{W_2}^2) < 0.0015$ . The limit is independent of  $\nu_R$  mass.
- $^{39}$  AQUINO 91 limits obtained from neutron lifetime and asymmetries together with unitarity of the CKM matrix. Manifest left-right asymmetry is assumed.  $^{40}$  BARBIERI 89B limit holds for  $m_{\nu_R} \leq 10$  MeV.
- <sup>41</sup> First JODIDIO 86 result assumes  $m_{W_R} = \infty$ , second is for unconstrained  $m_{W_R}$ .

#### THE Z' SEARCHES

Revised March 2002 by K.S. Babu (Oklahoma State University) and C. Kolda (Notre Dame University).

New massive and electrically neutral gauge bosons are a common feature of physics beyond the Standard Model. They are present in most extensions of the Standard Model gauge group, including models in which the Standard Model is embedded into a unifying group. They can also arise in certain classes of theories with extra dimensions. Whatever the source, such a gauge boson is called a Z'. While current theories suggest that there may be a multitude of such states at or just below the Planck scale, there exist many models in which the Z' sits at or near the weak scale. Models with extra neutral gauge bosons often contain charged gauge bosons as well; these are discussed in the review of W' physics.

The Lagrangian describing a single Z' and its interactions with the fields of the Standard Model is [1,2,3]:

$$\mathcal{L}_{Z'} = -\frac{1}{4} F'_{\mu\nu} F'^{\mu\nu} - \frac{\sin \chi}{2} F'_{\mu\nu} F^{\mu\nu} + M_{Z'}^2 Z'_{\mu} Z'^{\mu} + \delta M^2 Z'_{\mu} Z^{\mu} - \frac{e}{2c_W s_W} \sum_i \overline{\psi}_i \gamma^{\mu} (f_V^i - f_A^i \gamma^5) \psi_i Z'_{\mu}$$
(1)

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where  $c_W$ ,  $s_W$  are the cosine and sine of the weak angle,  $F_{\mu\nu}$ ,  $F'_{\mu\nu}$ are the field strength tensors for the hypercharge and the Z'gauge bosons respectively,  $\psi_i$  are the matter fields with Z'vector and axial charges  $f_V^i$  and  $f_A^i$ , and  $Z_\mu$  is the electroweak Z-boson. (The overall Z' coupling strength has been normalized to that of the usual Z.) The mass terms are assumed to come from spontaneous symmetry breaking via scalar expectation values; the  $\delta M^2$  term is generated by Higgs bosons that are charged under both the Standard Model and the extra gauge symmetry, and can have either sign. The above Lagrangian is general to all abelian and non-abelian extensions; however, for the non-abelian case,  $F'_{\mu\nu}$  is not gauge invariant and so the kinetic mixing parameter  $\chi = 0$ . Most analyses take  $\chi = 0$ , even for the abelian case, and so we do likewise here; see Ref. 3 for a discussion of observables with  $\chi \neq 0$ .

Strictly speaking, the Z' defined in the Lagrangian above is not a mass eigenstate since it can mix with the usual Z boson. The mixing angle is given by

$$\xi \simeq \frac{\delta M^2}{M_Z^2 - M_{Z'}^2} \,. \tag{2}$$

This mixing can alter a large number of the Z-pole observables, including the T-parameter which receives a contribution

$$\alpha T_{\text{new}} = \xi^2 \left( \frac{M_{Z'}^2}{M_Z^2} - 1 \right) \tag{3}$$

to leading order in small  $\xi$ . (For  $\chi \neq 0$ , both S and T receive additional contributions [4,3].) However, the oblique parameters do not encode all the effects generated by Z - Z' mixing; the mixing also alters the couplings of the Z itself, shifting its vector and axial couplings to  $T_3^i - 2Q^i s_W^2 + \xi f_V^i$  and  $T_3^i + \xi f_A^i$ respectively.

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If the Z' charges are generation-dependent, tree-level flavorchanging neutral currents will generically arise. There exist severe constraints in the first two generations coming from precision measurements such as the  $K_L - K_S$  mass splitting and  $B(\mu \to 3e)$ ; constraints on a Z' which couples differently only to the third generation are somewhat weaker. If the Z' interactions commute with the Standard Model gauge group, then per generation, there are only five independent  $Z'\bar{\psi}\psi$ couplings; one can choose them to be  $f_V^u, f_A^u, f_V^d, f_V^e, f_A^e$ . All other couplings can be determined in terms of these, e.g.,  $f_V^\nu = (f_V^e + f_A^e)/2$ .

**Experimental Constraints:** There are four primary sets of constraints on the existence of a Z' which will be considered here: precision measurements of neutral current processes at low energies, Z-pole constraints on Z-Z' mixing, indirect constraints from precision electroweak measurements off the Z-pole, and direct search constraints from production at very high energies. In principle, one should expect other new states to appear at the same scale as the Z', including its symmetry-breaking sector and any additional fermions necessary for anomaly cancellation. Because these states are highly model-dependent, searches for these states, or for Z' decays into them, are not included in the Listings.

Low-energy Constraints: After the gauge symmetry of the Z' and the electroweak symmetry are both broken, the Z of the Standard Model can mix with the Z', with mixing angle  $\xi$  defined above. As already discussed, this Z - Z' mixing implies a shift in the usual oblique parameters. Current bounds on T (and S) translate into stringent constraints on the mixing angle,  $\xi$ , requiring  $\xi \ll 1$ ; similar constraints on  $\xi$  arise from

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the LEP Z-pole data. Thus, we will only consider the small- $\xi$  limit henceforth.

Whether or not the new gauge interactions are parity violating, stringent constraints can arise from atomic parity violation (APV) and polarized electron-nucleon scattering experiments [5]. At low energies, the effective neutral current Lagrangian is conventionally written:

$$\mathcal{L}_{\rm NC} = \frac{G_F}{\sqrt{2}} \sum_{q=u,d} \left\{ C_{1q} (\bar{e}\gamma_\mu \gamma^5 e) (\bar{q}\gamma^\mu q) + C_{2q} (\bar{e}\gamma_\mu e) (\bar{q}\gamma^\mu \gamma^5 q) \right\}.$$
(4)

APV experiments are sensitive only to  $C_{1u}$  and  $C_{1d}$  through the "weak charge"  $Q_W = -2 \left[ C_{1u} (2Z + N) + C_{1d} (Z + 2N) \right]$ , where

$$C_{1q} = 2(1+\alpha T)(g_A^e + \xi f_A^e)(g_V^q + \xi f_V^q) + 2r(f_A^e f_V^q)$$
(5)

with  $r = M_Z^2/M_{Z'}^2$ . (Terms  $\mathcal{O}(r\xi)$  are dropped.) The *r*-dependent terms arise from Z' exchange and can interfere constructively or destructively with the Z contribution. In the limit  $\xi = r = 0$ , this reduces to the Standard Model expression. Polarized electron scattering is sensitive to both the  $C_{1q}$  and  $C_{2q}$  couplings, again as discussed in the Standard Model review. The  $C_{2q}$  can be derived from the expression for  $C_{1q}$  with the complete interchange  $V \leftrightarrow A$ .

Stringent limits also arise from neutrino-hadron scattering. One usually expresses experimental results in terms of the effective 4-fermion operators  $(\bar{\nu}\gamma_{\mu}\nu)(\bar{q}_{L,R}\gamma^{\mu}q_{L,R})$  with coefficients  $(2\sqrt{2}G_F)\epsilon_{L,R}(q)$ . (Again, see the Standard Model review.) In the presence of the Z and Z', the  $\epsilon_{L,R}(q)$  are given by:

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$$\epsilon_{L,R}(q) = \frac{1 + \alpha T}{2} \left\{ (g_V^q \pm g_A^q) [1 + \xi (f_V^\nu \pm f_A^\nu)] + \xi (f_V^q \pm f_A^q) \right\} \\ + \frac{r}{2} (f_V^q \pm f_A^q) (f_V^\nu \pm f_A^\nu) .$$
(6)

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Again, the r-dependent terms arise from Z'-exchange.

**Z-pole Constraints:** Electroweak measurements made at LEP and SLC while sitting on the Z-resonance are generally sensitive to Z' physics only through the mixing with the Z, unless the Z and Z' are very nearly degenerate. Constraints on the allowed mixing angle and Z' couplings arise by fitting all data simultaneously to the *ansatz* of Z - Z' mixing. A number of such fits are included in the Listings. If the listed analysis uses data only from the Z resonance, it is marked with a comment "Z parameters" while it is commented as "Electroweak" if low-energy data is also included in the fits. Both types of fits place simultaneous limits on the Z' mass and on  $\xi$ .

High-energy Indirect Constraints: At  $\sqrt{s} < M_{Z'}$ , but off the Z-pole, strong constraints on new Z' physics arise by comparing measurements of asymmetries and leptonic and hadronic cross-sections with their Standard Model predictions. These processes are sensitive not only to Z - Z' mixing, but also to direct Z' exchange primarily through  $\gamma - Z'$  and Z - Z' interference; therefore, information on the Z' couplings and mass can be extracted that is not accessible via Z - Z' mixing alone.

Far below the Z' mass scale, experiments at a given  $\sqrt{s}$  are only sensitive to the scaled Z' couplings  $\sqrt{s}f_{V,A}^i/M_{Z'}$ . However, the Z' mass and overall magnitude of the couplings can be separately extracted if measurements are made at more than one energy. As  $\sqrt{s}$  approaches  $M_{Z'}$  the Z' exchange can no longer be approximated by a contact interaction and the mass and couplings can be simultaneously extracted.

Z' studies done before LEP relied heavily on this approach; see, for example, Ref. 6. LEP has also done similar work using data collected above the Z-peak; see, for example, Ref. 7.

For indirect Z' searches at future facilities, see, for example, Refs. 8,9. At a hadron collider the possibility of measuring leptonic forward-backward asymmetries has been suggested [10] and used [11] in searches for a Z' below its threshold.

**Direct Search Constraints:** Finally, high-energy experiments have searched for on-shell Z' production and decay. Searches can be classified by the initial state off of which the Z' is produced, and the final state into which the Z' decays; exotic decays of a Z' are not included in the listings. Experiments to date have been sensitive to Z' production via their coupling to quarks ( $p\bar{p}$  colliders), to electrons ( $e^+e^-$ ), or to both (ep).

For a heavy Z'  $(M_{Z'} \gg M_Z)$ , the best limits come from  $p\bar{p}$  machines via Drell-Yan production and subsequent decay to charged leptons. For  $M_{Z'} > 600 \text{ GeV}$ , CDF [12] quotes limits on  $\sigma(p\bar{p} \rightarrow Z'X) \cdot B(Z' \rightarrow \ell^+ \ell^-) < 0.04 \text{ pb}$  at 95% C.L. for  $\ell = e + \mu$  combined; DØ [13] quotes  $\sigma \cdot B < 0.06 \text{ pb}$  for  $\ell = e$  and  $M_{Z'} > 500 \text{ GeV}$ . For smaller masses, the bounds can be found in the original literature. For studies of the search capabilities of future facilities, see, for example, Ref. 8.

If the Z' has suppressed, or no, couplings to leptons (*i.e.*, it is leptophobic), then experimental sensitivities are much weaker. Searches for a Z' via hadronic decays at CDF [14] are unable to rule out a Z' with quark couplings identical to those of the Z in any mass region. UA2 [15] does find  $\sigma \cdot B(Z' \to jj) < 11.7$  pb at 90% C.L. for  $M_{Z'} > 200$  GeV, with more complicated bounds in the range 130 GeV  $< M_{Z'} < 200$  GeV.

For a light Z'  $(M_{Z'} < M_Z)$ , direct searches in  $e^+e^-$  colliders have ruled out any Z', unless it has extremely weak couplings to leptons. For a combined analysis of the various pre-LEP experiments see Ref. 6.

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**Canonical Models:** One of the prime motivations for an additional Z' has come from string theory, in which certain compactifications lead naturally to an  $E_6$  gauge group, or one of its subgroups.  $E_6$  contains two U(1) factors beyond the Standard Model, a basis for which is formed by the two groups U(1)<sub> $\chi$ </sub> and U(1)<sub> $\psi$ </sub>, defined via the decompositions  $E_6 \rightarrow$  SO(10) × U(1)<sub> $\psi$ </sub> and SO(10)  $\rightarrow$  SU(5) × U(1)<sub> $\chi$ </sub>; one special case often encountered is U(1)<sub> $\eta$ </sub>, where  $Q_{\eta} = \sqrt{\frac{3}{8}}Q_{\chi} - \sqrt{\frac{5}{8}}Q_{\psi}$ . The charges of the SM fermions under these U(1)'s can be found in Table 1, and a discussion of their experimental signatures can be found in Ref. 16. A separate listing appears for each of the canonical models, with direct and indirect constraints combined.

Table 1: Charges of Standard Model fermions
in canonical $Z'$ models.

	Y	$T_{3R}$	B-L	$\sqrt{24}Q_{\chi}$	$\sqrt{\frac{72}{5}}Q_{\psi}$	$Q_\eta$
$ u_L, e_L$	$-\frac{1}{2}$	0	-1	+3	+1	$+\frac{1}{6}$
$ u_R$	0	$+\frac{1}{2}$	-1	+5	-1	$+\frac{5}{6}$
$e_R$	-1	$-\frac{1}{2}$	-1	+1	-1	$+\frac{1}{3}$
$u_L, d_L$	$+\frac{1}{6}$	0	$+\frac{1}{3}$	-1	+1	$-\frac{1}{3}$
$u_R$	$+\frac{2}{3}$	$+\frac{1}{2}$	$+\frac{1}{3}$	+1	-1	$+\frac{1}{3}$
$d_R$	$-\frac{1}{3}$	$-\frac{1}{2}$	$+\frac{1}{3}$	-3	-1	$-\frac{1}{6}$

It is also common to express experimental bounds in terms of a toy Z', usually denoted  $Z'_{\rm SM}$ . This  $Z'_{\rm SM}$ , of arbitrary mass, couples to the SM fermions identically to the usual Z. Almost all analyses of Z' physics have worked with one of these canonical models and have assumed zero kinetic mixing at the weak scale.

**Extra Dimensions:** A new motivation for Z' searches comes from recent work on extensions of the Standard Model into extra dimensions. (See the "Review of Extra Dimensions" for many details not included here.) In some classes of these models, the gauge bosons of the Standard Model can inhabit these new directions [17]. When compactified down to the usual (3+1) dimensions, the extra degrees of freedom that were present in the higher-dimensional theory (associated with propagation in the extra dimensions) appear as a tower of massive gauge bosons, called Kaluza-Klein (KK) states. The simplest case is the compactification of a (4+d)-dimensional space on a d-torus  $(T^d)$  of uniform radius R in all d directions. Then a tower of massive gauge bosons are present with masses

$$M_{V_{\vec{n}}}^2 = M_{V_{\vec{0}}}^2 + \frac{\vec{n} \cdot \vec{n}}{R^2} \,, \tag{7}$$

where V represents any of the gauge fields of the Standard Model and  $\vec{n}$  is a *d*-vector whose components are semi-positive integers; the vector  $\vec{n} = (0, 0, ...0)$  corresponds to the "zeromode" gauge boson, which is nothing more than the usual gauge boson of the Standard Model, with mass  $M_{V_{\vec{0}}} = M_V$ . Compactifications on either non-factorizable or asymmetric manifolds can significantly alter the KK mass formula, but a tower of states will nonetheless persist. All bounds cited in the Listings

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assume the maximally symmetric spectrum given above for simplicity.

The KK mass formula, coupled with the absence of any observational evidence for W' or Z' states below the weak scale, implies that the extra dimensions in which gauge bosons can propagate must have inverse radii greater than at least a few hundred GeV. If any extra dimensions are larger than this, gravity alone may propagate in them.

Though the gauge principle guarantees that the usual Standard Model gauge fields couple with universal strength (or gauge coupling) to all charged matter, the coupling of KK bosons to ordinary matter is highly model-dependent. In the simplest case, all Standard Model fields are localized at the same point in the d-dimensional subspace; in the parlance of the field, they all live on the same 3-brane. Then the couplings of KK bosons are identical to those of the usual gauge fields, but enhanced:  $g_{KK} = \sqrt{2} g$ . However, in many models, particularly those which naturally suppress proton decay [18], it is common to find ordinary fermions living on different, parallel branes in the extra dimensions. In such cases, different fermions experience very different coupling strengths for the KK states; the effective coupling varies fermion by fermion, and also KK mode by KK mode. In the particular case that fermions of different generations with identical quantum numbers are placed on different branes, large flavor-changing neutral currents can occur unless the mass scale of the KK states is very heavy:  $R^{-1} \gtrsim 1000 \,\mathrm{TeV}$  [19]. In the Listings, all bounds assume that Standard Model fermions live on a single 3-brane. (The case of the Higgs field is again complicated; see the footnotes on the individual listings.)

In some sense, searches for KK bosons are no different than searches for any other Z' or W'; in fact, bounds on

the artificially defined  $Z'_{\rm SM}$  are almost precisely bounds on the first KK mode of the  $Z^0$ , modulo the  $\sqrt{2}$  enhancement in the coupling strength. To date, no experiment has examined direct production of KK  $Z^0$  bosons, but an approximate bound of 820 GeV [20] can be inferred from the CDF bound on  $Z'_{SM}$  [12].

Indirect bounds have a very different behavior for KK gauge bosons than for canonical Z' bosons; a number of indirect bounds are given in the Listings. Indirect bounds arise from virtual boson exchange and require a summation over the entire tower of KK states. For d > 1, this summation diverges, a remnant of the non-renormalizability of the underlying (4 + d)dimensional field theory. In a fully consistent theory, such as a string theory, the summation would be regularized and finite. However, this procedure cannot be uniquely defined within the confines of our present knowledge, and so most authors choose to terminate the sum with an explicit cut-off,  $\Lambda_{KK}$ , set equal to the "Planck scale" of the D-dimensional theory,  $M_D$  [21]. Reasonable arguments exist that this cut-off could be very different and could vary by process, and so these bounds should be regarded merely as indicative [22].

## References

- 1. B. Holdom, Phys. Lett. **166B**, 196 (1986).
- F. del Aguila, Acta Phys. Polon. **B25**, 1317 (1994);
   F. del Aguila, M. Cvetič, and P. Langacker, Phys. Rev. **D52**, 37 (1995).
- K.S. Babu, C. Kolda, and J. March-Russell, Phys. Rev. D54, 4635 (1996); *ibid.*, D57, 6788 (1998).
- 4. B. Holdom, Phys. Lett. **B259**, 329 (1991).
- 5. J. Kim *et al.*, Rev. Mod. Phys. **53**, 211 (1981);
  U. Amaldi *et al.*, Phys. Rev. **D36**, 1385 (1987);
  W. Marciano and J. Rosner, Phys. Rev. Lett. **65**, 2963 (1990) (*Erratum:* **68**, 898 (1992));

K. Mahanthappa and P. Mohapatra, Phys. Rev. D43, 3093 (1991) (*Erratum:* D44, 1616 (1991));
P. Langacker and M. Luo, Phys. Rev. D45, 278 (1992);
P. Langacker, M. Luo, and A. Mann, Rev. Mod. Phys. 64, 87 (1992).

- 6. L. Durkin and P. Langacker, Phys. Lett. **166B**, 436 (1986).
- 7. P. Abreu *et al.*, (DELPHI Collaboration), Eur. Phys. J. C11, 383 (1999);
  R. Barate *et al.*, (ALEPH Collaboration) Eur. Phys. J. C12, 183 (1999).
- 8. M. Cvetiĉ and S. Godfrey, in *Electroweak Symmetry Break*ing and New Physics at the TeV Scale, eds. T. Barklow et al., (World Scientific 1996), p. 383 [hep-ph/9504216].
- 9. T. Rizzo, Phys. Rev. **D55**, 5483 (1997).
- 10. J. L. Rosner, Phys. Rev. **D54**, 1078 (1996).
- T. Affolder *et al.*, (CDF Collaboration), Phys. Rev. Lett. 87, 131802 (2001).
- 12. F. Abe *et al.*, (CDF Collaboration), Phys. Rev. Lett. **79**, 2191 (1997).
- V. Abazov *et al.*, (D0 Collaboration), Phys. Rev. Lett. 87, 061802 (2001).
- F. Abe *et al.*, (CDF Collaboration), Phys. Rev. D55, 5263R (1997) and Phys. Rev. Lett. 82, 2038 (1999).
- J. Alitti *et al.*, (UA2 Collaboration), Nucl. Phys. **B400**, 3 (1993).
- 16. J. Hewett and T. Rizzo, Phys. Rept. 183, 193 (1989).
- 17. I. Antoniadis, Phys. Lett. B246, 377 (1990);
  I. Antoniadis, K. Benakli, and M. Quiros, Phys. Lett. B331, 313 (1994);
  K. Dienes, E. Dudas, and T. Gherghetta, Phys. Lett. B436, 55 (1998);
  A. Pomarol and M. Quiros, Phys. Lett. B438, 255 (1998).
- 18. N. Arkani-Hamed and M. Schmaltz, Phys. Rev. **D61**, 033005 (2000).
- 19. A. Delgado, A. Pomarol, and M. Quiros, JHEP **0001**, 030 (2000).

- 20. M. Masip and A. Pomarol, Phys. Rev. **D60**, 096005 (1999).
- 21. G. Giudice, R. Rattazzi, and J. Wells, Nucl. Phys. B544, 3 (1999);
  T. Han, J. Lykken, and R. Zhang, Phys. Rev. D59, 105006 (1999);
  J. Hewett, Phys. Rev. Lett. 82, 4765 (1999).
- 22. See for example: M. Bando *et al.*, Phys. Rev. Lett. 83, 3601 (1999);
  T. Rizzo and J. Wells, Phys. Rev. D61, 016007 (2000);

S. Cullen, M. Perelstein, and M. Peskin, Phys. Rev. **D62**, 055012 (2000).

#### MASS LIMITS for Z' (Heavy Neutral Vector Boson Other Than Z)

### Limits for $Z'_{SM}$

 $Z'_{SM}$  is assumed to have couplings with quarks and leptons which are identical to those of Z, and decays only to known fermions.

VALUE (GeV)	<u>CL%</u>	DOCUMENT ID	TECN	COMMENT
>1500	95	<sup>42</sup> CHEUNG	01b RVUE	Electroweak
> 690	95	<sup>43</sup> ABE	97s CDF	$p \overline{p}; Z'_{SM} \rightarrow e^+ e^-,$
				$\mu^+\mu^-$
• • • We do not ι	ise the followi	ng data for average	s, fits, limits	s, etc. ● ● ●
> 670	95	<sup>44</sup> ABAZOV	01B D0	$p \overline{p}, Z'_{SM} \rightarrow e^+ e^-$
> 710	95	<sup>45</sup> ABREU	00s DLPH	$I e^+e^-$
> 898	95	<sup>46</sup> BARATE	001 ALEP	$e^+e^-$
> 809	95	<sup>47</sup> ERLER	99 RVUE	
> 490	95	ABACHI	96d D0	$p\overline{p}; Z'_{SM} \rightarrow e^+e^-$
> 398	95	<sup>48</sup> VILAIN	94b CHM2	$\begin{array}{ccc} & & & \\ 2 & \nu_{\mu} e \rightarrow \nu_{\mu} e \text{ and} \\ & & \overline{\nu}_{\mu} e \rightarrow \overline{\nu}_{\mu} e \end{array}$
				$\overline{\overline{\nu}}_{\mu} e  ightarrow \overline{\overline{\nu}}_{\mu} e$
> 237	90	<sup>49</sup> ALITTI	93 UA2	$p\overline{p}; Z'_{SM} \rightarrow q\overline{q}$
none 260–600	95	<sup>50</sup> RIZZO	93 RVUE	$p \overline{p}; Z'_{SM} \rightarrow q \overline{q}$ $p \overline{p}; Z'_{SM} \rightarrow q \overline{q}$
> 426	90	<sup>51</sup> ABE	90F VNS	e <sup>+</sup> e <sup>-</sup>
		£		the second state of the state o

<sup>42</sup> CHEUNG 01B limit is derived from bounds on contact interactions in a global electroweak analysis.

<sup>43</sup>ABE 97S find  $\sigma(Z') \times B(e^+e^-, \mu^+\mu^-) <$  40 fb for  $m_{Z'} >$  600 GeV at  $\sqrt{s}$ = 1.8 TeV.

<sup>44</sup> ABAZOV 01B search for resonances in  $p\overline{p} \rightarrow e^+e^-$  at  $\sqrt{s}=1.8$  TeV. They find  $\sigma \cdot B(Z' \rightarrow ee) < 0.06$  pb for  $M_{Z'} > 500$  GeV.

<sup>45</sup> ABREU 00S uses LEP data at  $\sqrt{s}$ =90 to 189 GeV.

<sup>46</sup> BARATE 00I search for deviations in cross section and asymmetries in  $e^+e^- \rightarrow$  fermions at  $\sqrt{s}=90$  to 183 GeV. Assume  $\theta=0$ . Bounds in the mass-mixing plane are shown in their Figure 18.

 $^{47}$  ERLER 99 give 90%CL limit on the Z-Z' mixing  $-0.0041 < \theta < 0.0003.$   $\rho_0{=}1$  is assumed.

<sup>48</sup> VILAIN 94B assume  $m_t = 150$  GeV.

<sup>49</sup> ALITTI 93 search for resonances in the two-jet invariant mass. The limit assumes  $B(Z' \rightarrow q\overline{q})=0.7$ . See their Fig. 5 for limits in the  $m_{Z'}-B(q\overline{q})$  plane.

<sup>50</sup> RIZZO 93 analyses CDF limit on possible two-jet resonances.

<sup>51</sup> ABE 90F use data for R,  $R_{\ell\ell}$ , and  $A_{\ell\ell}$ . They fix  $m_W = 80.49 \pm 0.43 \pm 0.24$  GeV and  $m_Z = 91.13 \pm 0.03$  GeV.

#### Limits for $Z_{LR}$

 $Z_{LR}$  is the extra neutral boson in left-right symmetric models.  $g_L = g_R$  is assumed unless noted. Values in parentheses assume stronger constraint on the Higgs sector, usually motivated by specific left-right symmetric models (see the Note on the W'). Values in brackets are from cosmological and astrophysical considerations and assume a light right-handed neutrino. Direct search bounds assume decays to Standard Model fermions only, unless noted.

VALUE (GeV)	<u>CL%</u>	DOCUMENT ID		TECN	COMMENT				
>860	95	<sup>52</sup> CHEUNG	<b>01</b> B	RVUE	Electroweak				
>630	95	<sup>53</sup> ABE	<b>97</b> S	CDF	$p\overline{p}; Z'_{IR} \rightarrow e^+e^-,$				
					$\mu^+\mu^-$				
• • • We do not use the following data for averages, fits, limits, etc. • •									
>380	95	<sup>54</sup> ABREU	00s	DLPH	e <sup>+</sup> e <sup>-</sup>				
>436	95	<sup>55</sup> BARATE	001	ALEP	e <sup>+</sup> e <sup>-</sup>				
>550	95	<sup>56</sup> CHAY	00	RVUE	Electroweak				
		<sup>57</sup> ERLER	00		Cs				
		<sup>58</sup> CASALBUONI	99	RVUE	Cs				
(> 1205)	90	<sup>59</sup> CZAKON	99	RVUE	Electroweak				
>564	95	<sup>60</sup> ERLER	99	RVUE	Electroweak				
(> 1673)	95	<sup>61</sup> ERLER	99	RVUE	Electroweak				
(> 1700)	68	<sup>62</sup> BARENBOIM	98	RVUE	Electroweak				
>244	95	<sup>63</sup> CONRAD	98	RVUE	$ u_{\mu}$ N scattering				
>253	95	<sup>64</sup> VILAIN			$\overline{ u_{\mu}^{r}} e  ightarrow \  u_{\mu} e$ and $\overline{ u}_{\mu} e  ightarrow$				
					$\overline{\nu}_{\mu}e^{\mu}$				
none 200–600	95	<sup>65</sup> RIZZO	93	RVUE	$p\overline{p}; \overline{Z}_{IR} \rightarrow q\overline{q}$				
[> 2000]		WALKER	91		Nucleosynthesis; light $\nu_R$				
none 200–500		<sup>66</sup> GRIFOLS	90	ASTR	SN 1987A; light $\nu_R$				
none 350–2400		<sup>67</sup> BARBIERI			SN 1987A; light $\nu_R$				
52 01 51 10 0 01 11									

<sup>52</sup> CHEUNG 01B limit is derived from bounds on contact interactions in a global electroweak analysis.

 $^{53}$  ABE 97S find  $\sigma(Z') \times B(e^+e^-, \mu^+\mu^-) < 40$  fb for  $m_{Z'} > 600$  GeV at  $\sqrt{s} = 1.8$  TeV.

<sup>54</sup> ABREU 00S give 95%CL limit on Z-Z' mixing  $|\theta| < 0.0018$ . See their Fig. 6 for the limit contour in the mass-mixing plane.  $\sqrt{s}=90$  to 189 GeV.

<sup>55</sup> BARATE 00I search for deviations in cross section and asymmetries in  $e^+e^- \rightarrow$  fermions at  $\sqrt{s}=90$  to 183 GeV. Assume  $\theta=0$ . Bounds in the mass-mixing plane are shown in their Figure 18.

<sup>56</sup> CHAY 00 also find  $-0.0003 < \theta < 0.0019$ . For  $g_R$  free,  $m_{Z'} > 430$  GeV.

<sup>57</sup> ERLER 00 discuss the possibility that a discrepancy between the observed and predicted values of  $Q_W(Cs)$  is due to the exchange of Z'. The data are better described in a certain class of the Z' models including  $Z_{LR}$  and  $Z_{\gamma}$ .

- <sup>58</sup> CASALBUONI 99 discuss the discrepancy between the observed and predicted values of  $Q_W(Cs)$ . It is shown that the data are better described in a class of models including the  $Z_{LR}$  model.
- <sup>59</sup> CZAKON 99 perform a simultaneous fit to charged and neutral sectors. Assumes manifest left-right symmetric model. Finds  $|\theta| < 0.0042$ .
- $^{60}$  ERLER 99 give 90%CL limit on the Z-Z' mixing  $-0.0009 < \theta < 0.0017.$
- $^{61}$  ERLER 99 assumes 2 Higgs doublets, tranforming as 10 of SO(10), embedded in  $E_6$ .
- <sup>62</sup> BARENBOIM 98 also gives 68% CL limits on the Z-Z' mixing  $-0.0005 < \theta < 0.0033$ . Assumes Higgs sector of minimal left-right model.
- $^{63}$  CONRAD 98 limit is from measurements at CCFR, assuming no Z-Z' mixing.
- $^{64}$  VILAIN 94B assume  $m_t = 150$  GeV and  $\theta{=}0$ . See Fig. 2 for limit contours in the mass-mixing plane.
- <sup>65</sup> RIZZO 93 analyses CDF limit on possible two-jet resonances.
- $^{66}\,{\rm GRIFOLS}$  90 limit holds for  $m_{\nu_R}\lesssim 1$  MeV. A specific Higgs sector is assumed. See also GRIFOLS 90D, RIZZO 91.
- $^{67}\,{\rm BARBIERI}$  89B limit holds for  $m_{\nu_R} \leq$  10 MeV. Bounds depend on assumed supernova core temperature.

### Limits for $Z_{\chi}$

 $Z_{\chi}$  is the extra neutral boson in SO(10)  $\rightarrow$  SU(5)  $\times$  U(1)<sub> $\chi$ </sub>.  $g_{\chi} = e/\cos\theta_W$  is assumed unless otherwise stated. We list limits with the assumption  $\rho = 1$  but with no further constraints on the Higgs sector. Values in parentheses assume stronger constraint on the Higgs sector motivated by superstring models. Values in brackets are from cosmological and astrophysical considerations and assume a light right-handed neutrino

Neutrino. VALUE (GeV)	CL%	DOCUMENT ID		TECN	COMMENT
>680	<u>95</u>	<sup>68</sup> CHEUNG	010		Electroweak
>595	95	<sup>69</sup> АВЕ	97S	CDF	$p \overline{p}; ~Z'_{\chi}  ightarrow ~e^+ e^-,~\mu^+ \mu^-$
$\bullet \bullet \bullet$ We do not use	the follo	wing data for avera	iges,	fits, limi	ts, etc. ● ● ●
>440	95	<sup>70</sup> ABREU	<b>00</b> S	DLPH	e <sup>+</sup> e <sup>-</sup>
>533	95	<sup>71</sup> BARATE	001	ALEP	e <sup>+</sup> e <sup>-</sup>
>554	95	<sup>72</sup> CHO	00	RVUE	Electroweak
		<sup>73</sup> ERLER	00	RVUE	Cs
		<sup>74</sup> ROSNER	00	RVUE	Cs
>545	95	<sup>75</sup> ERLER	99	RVUE	Electroweak
(> 1368)	95	<sup>76</sup> ERLER	99	RVUE	Electroweak
>215	95	<sup>77</sup> CONRAD	98	RVUE	$ u_{\mu} N$ scattering
>190	95	<sup>78</sup> ARIMA	97	VNS	Bhabha scattering
>262	95	<sup>79</sup> VILAIN	<b>94</b> B	CHM2	$ u_{\mu} e  ightarrow   u_{\mu} e$ and $\overline{ u}_{\mu} e  ightarrow$
					$\overline{\nu}_{\mu}e$
[>1470]		<sup>80</sup> FARAGGI	91	COSM	Nucleosynthesis; light $\nu_R$
>231	90	<sup>81</sup> ABE	90F	VNS	e <sup>+</sup> e <sup>-</sup>
[> 1140]		<sup>82</sup> GONZALEZ-G	<b>90</b> D	COSM	Nucleosynthesis; light $ u_R$
[> 2100]		<sup>83</sup> GRIFOLS			SN 1987A; light $\nu_R$
<sup>68</sup> CHEUNG 01B lim analysis.	it is deriv	red from bounds on	conta	act intera	actions in a global electroweak

<sup>69</sup> ABE 97s find  $\sigma(Z') \times B(e^+e^-, \mu^+\mu^-) < 40$  fb for  $m_{Z'} > 600$  GeV at  $\sqrt{s} = 1.8$  TeV.

<sup>70</sup>ABREU 00S give 95%CL limit on Z-Z' mixing  $|\theta| < 0.0017$ . See their Fig. 6 for the limit contour in the mass-mixing plane.  $\sqrt{s}$ =90 to 189 GeV.

- <sup>71</sup> BARATE 001 search for deviations in cross section and asymmetries in  $e^+e^- \rightarrow$  fermions at  $\sqrt{s}$ =90 to 183 GeV. Assume  $\theta$ =0. Bounds in the mass-mixing plane are shown in their Figure 18.
- <sup>72</sup> CHO 00 use various electroweak data to constrain Z' models assuming  $m_H$ =100 GeV. See Fig. 3 for limits in the mass-mixing plane.
- <sup>73</sup> ERLER 00 discuss the possibility that a discrepancy between the observed and predicted values of  $Q_W(Cs)$  is due to the exchange of Z'. The data are better described in a certain class of the Z' models including  $Z_{LR}$  and  $Z_{\gamma}$ .
- <sup>74</sup> ROSNER 00 discusses the possiblitiy that a discrepancy between the observed and predicted values of  $Q_W(Cs)$  is due to the exchange of Z'. The data are better described in a certain class of the Z' models including  $Z_{\gamma}$ .

<sup>75</sup> ERLER 99 give 90%CL limit on the Z-Z' mixing  $-0.0020 < \theta < 0.0015$ .

<sup>76</sup> ERLER 99 assumes 2 Higgs doublets, tranforming as 10 of SO(10), embedded in  $E_6$ .

- <sup>77</sup> CONRAD 98 limit is from measurements at CCFR, assuming no Z-Z' mixing.
- <sup>78</sup> Z-Z' mixing is assumed to be zero.  $\sqrt{s}$  = 57.77 GeV.
- $^{79}$  VILAIN 94B assume  $m_t$  = 150 GeV and  $heta{=}0$ . See Fig. 2 for limit contours in the mass-mixing plane.
- $^{80}$  FARAGGI 91 limit assumes the nucleosynthesis bound on the effective number of neutrinos  $\Delta \textit{N}_{\nu}~<~0.5$  and is valid for  $m_{\nu_{R}}~<1$  MeV.
- <sup>81</sup> ABE 90F use data for R,  $R_{\ell\ell}$ , and  $A_{\ell\ell}$ . ABE 90F fix  $m_W$  = 80.49 ± 0.43 ± 0.24 GeV and  $m_Z$  = 91.13 ± 0.03 GeV.
- $^{82}$  Assumes the nucleosynthesis bound on the effective number of light neutrinos ( $\delta N_{
  u}~<~1)$ and that  $\nu_R$  is light (  $\lesssim$  1 MeV).  $^{83}\,{\rm GRIFOLS}$  90 limit holds for  $m_{\nu_R}\,\lesssim$  1 MeV. See also GRIFOLS 90D, RIZZO 91.

#### Limits for $Z_{\psi}$

 $Z_{\psi}$  is the extra neutral boson in  $E_6 \to SO(10) imes U(1)_{\psi}$ .  $g_{\psi} = e/\cos\theta_W$  is assumed unless otherwise stated. We list limits with the assumption  $\rho = 1$  but with no further constraints on the Higgs sector. Values in brackets are from cosmological and astrophysical considerations and assume a light right-handed neutrino.

VALUE (GeV)	CL%	DOCUMENT ID		TECN	COMMENT
>350	95	<sup>84</sup> ABREU	00S	DLPH	e <sup>+</sup> e <sup>-</sup>
>590	95	<sup>85</sup> ABE	<b>97</b> S	CDF	$p  \overline{p};  Z_{\psi}^{\prime}  ightarrow  e^+  e^-,  \mu^+  \mu^-$
• • • We do not use	the follow	wing data for avera	ges,	fits, limi	ts, etc. ● ● ●
>294	95	<sup>86</sup> BARATE	001	ALEP	e <sup>+</sup> e <sup>-</sup>
>137	95	<sup>87</sup> CHO	00	RVUE	Electroweak
>146	95	<sup>88</sup> ERLER	99	RVUE	Electroweak
> 54	95	<sup>89</sup> CONRAD	98	RVUE	$ u_{\mu} N$ scattering
>135	95	<sup>90</sup> VILAIN			$\stackrel{\sim}{ u_{\mu}} e  ightarrow \  u_{\mu} e  ext{ and } \overline{ u}_{\mu} e  ightarrow$
					$\overline{\nu}_{\mu}e^{-i\omega}$
>105	90	<sup>91</sup> ABE			e <sup>+</sup> e <sup>-</sup>
[> 160]		<sup>92</sup> GONZALEZ-G.	.90D	COSM	Nucleosynthesis; light $ u_R$
[> 2000]		<sup>93</sup> GRIFOLS			SN 1987A; light $\nu_R$

<sup>84</sup>ABREU 00S give 95%CL limit on Z-Z' mixing  $|\theta| < 0.0018$ . See their Fig. 6 for the limit contour in the mass-mixing plane.  $\sqrt{s}=90$  to 189 GeV.

<sup>85</sup> ABE 97S find  $\sigma(Z') \times B(e^+e^-, \mu^+\mu^-) < 40$  fb for  $m_{\tau'} > 600$  GeV at  $\sqrt{s} = 1.8$  TeV.

- <sup>86</sup> BARATE 00I search for deviations in cross section and asymmetries in  $e^+e^- \rightarrow$  fermions at  $\sqrt{s}=90$  to 183 GeV. Assume  $\theta=0$ . Bounds in the mass-mixing plane are shown in their Figure 18.
- <sup>87</sup> CHO 00 use various electroweak data to constrain Z' models assuming  $m_H$ =100 GeV. See Fig. 3 for limits in the mass-mixing plane.

 $^{88}$  ERLER 99 give 90%CL limit on the Z-Z' mixing  $-0.0013 < \theta < 0.0024.$ 

- $^{89}$  CONRAD 98 limit is from measurements at CCFR, assuming no Z-Z' mixing.
- $^{90}$  VILAIN 94B assume  $m_t = 150$  GeV and  $\theta{=}0$ . See Fig. 2 for limit contours in the mass-mixing plane.
- <sup>91</sup>ABE 90F use data for R,  $R_{\ell\ell}$ , and  $A_{\ell\ell}$ . ABE 90F fix  $m_W = 80.49 \pm 0.43 \pm 0.24$  GeV and  $m_Z = 91.13 \pm 0.03$  GeV.
- <sup>92</sup> Assumes the nucleosynthesis bound on the effective number of light neutrinos ( $\delta N_{\nu} < 1$ ) and that  $\nu_R$  is light ( $\leq 1$  MeV).

 $^{93}\,{\rm GRIFOLS}$  90D limit holds for  $m_{\nu_R}\,\lesssim\,1$  MeV. See also RIZZO 91.

#### Limits for $Z_n$

 $Z_{\eta}$  is the extra neutral boson in E<sub>6</sub> models, corresponding to  $Q_{\eta} = \sqrt{3/8} Q_{\chi} - \sqrt{5/8} Q_{\psi}$ .  $g_{\eta} = e/\cos\theta_W$  is assumed unless otherwise stated. We list limits with the assumption  $\rho = 1$  but with no further constraints on the Higgs sector. Values in parentheses assume stronger constraint on the Higgs sector motivated by superstring models. Values in brackets are from cosmological and astrophysical considerations and assume a light right-handed neutrino.

VALUE (GeV)	<u>CL%</u>	DOCUMENT ID		TECN	COMMENT			
>619	95	<sup>94</sup> CHO	00	RVUE	Electroweak			
>620	95	<sup>95</sup> ABE	<b>97</b> S	CDF	p $\overline{p};~Z_{\eta}^{\prime} ightarrow~e^{+}e^{-},~\mu^{+}\mu^{-}$			
$\bullet$ $\bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet$ $\bullet$								
>310	95	<sup>96</sup> ABREU	00s	DLPH	e <sup>+</sup> e <sup>-</sup>			
>329	95	<sup>97</sup> BARATE	001	ALEP	e <sup>+</sup> e <sup>-</sup>			
>365	95	<sup>98</sup> ERLER	99	RVUE	Electroweak			
> 87	95	<sup>99</sup> CONRAD	98	RVUE	$ u_{\mu} N$ scattering			
>100	95	<sup>100</sup> VILAIN			$\overline{ u}_{\mu}^{r}e ightarrow  u_{\mu}e$ and $\overline{ u}_{\mu}e ightarrow$			
					$\overline{\nu}_{\mu}$ e			
>125	90	<sup>101</sup> ABE		VNS				
[> 820]		<sup>102</sup> GONZALEZ-G	<b>90</b> D	COSM	Nucleosynthesis; light $ u_R$			
[> 3300]		<sup>103</sup> GRIFOLS	90	ASTR	SN 1987A; light $\nu_R$			
[> 1040]		<sup>102</sup> LOPEZ	90	COSM	Nucleosynthesis; light $\nu_R$			
<sup>94</sup> CHO 00 use vario	ous elect	roweak data to cons	strair	$Z' \mod$	dels assuming <i>m<sub>H</sub>=</i> 100 GeV.			
See Fig. 3 for limi	ts in the	e mass-mixing plane						
$^{95}$ ABE 97S find $\sigma(Z$	$\mathbf{Z}')  imes B(\epsilon)$	$e^+e^-$ , $\mu^+\mu^-$ ) $<$ 40 $^+$	fb fo	r <i>m <sub>7</sub>, &gt;</i>	$\cdot$ 600 GeV at $\sqrt{s}$ = 1.8 TeV.			
					024. See their Fig. 6 for the			

limit contour in the mass-mixing plane.  $\sqrt{s}=90$  to 189 GeV.

<sup>97</sup> BARATE 00I search for deviations in cross section and asymmetries in  $e^+e^- \rightarrow$  fermions at  $\sqrt{s}=90$  to 183 GeV. Assume  $\theta=0$ . Bounds in the mass-mixing plane are shown in their Figure 18.

 $^{98}$  ERLER 99 give 90%CL limit on the Z-Z' mixing  $-0.0062 < \theta < 0.0011$ .

 $^{99}$ CONRAD 98 limit is from measurements at CCFR, assuming no Z-Z' mixing.

- $^{100}$  VILAIN 94B assume  $m_t = 150$  GeV and  $\theta = 0$ . See Fig. 2 for limit contours in the mass-mixing plane.
- $^{101}$  ABE 90F use data for R,  $R_{\ell\ell}$ , and  $A_{\ell\ell}$ . ABE 90F fix  $m_W=80.49\pm0.43\pm0.24$  GeV and  $m_Z=91.13\pm0.03$  GeV.
- <sup>102</sup> These authors claim that the nucleosynthesis bound on the effective number of light neutrinos ( $\delta N_{\nu} < 1$ ) constrains Z' masses if  $\nu_R$  is light ( $\lesssim 1$  MeV).
- $^{103}\,{\rm GRIFOLS}$  90 limit holds for  $m_{\nu_R}\,\lesssim\,1$  MeV. See also GRIFOLS 90D, RIZZO 91.

#### Limits for other Z'

VALUE (GeV)	DOCUMENT IL	)	TECN	COMMENT			
$\bullet \bullet \bullet$ We do not use the follow	ving data for averag	ges, fits	, limits,	etc. • • •			
	<sup>104</sup> CHO 105 CHO 106 ABE	98	RVUE	$E_6$ -motivated $E_6$ -motivated $Z' \rightarrow \overline{q} q$			
<sup>104</sup> CHO 00 use various electroweak data to constrain Z' models assuming $m_H$ =100 GeV. See Fig. 2 for limits in general $E_6$ -motivated models.							
<sup>105</sup> CHO 98 study constraints on four-Fermi contact interactions obtained from low-energy electroweak experiments, assuming no $Z-Z'$ mixing.							

<sup>106</sup> Search for Z' decaying to dijets at  $\sqrt{s}$ =1.8 TeV. For Z' with electromagnetic strength coupling, no bound is obtained.

#### Indirect Constraints on Kaluza-Klein Gauge Bosons

Bounds on a Kaluza-Klein excitation of the Z boson or photon in d=1 extra dimension. These bounds can also be interpreted as a lower bound on 1/R, the size of the extra dimension. Unless otherwise stated, bounds assume all fermions live on a single brane and all gauge fields occupy the 4+d-dimensional bulk.

VALU	JE (TeV)	CL%	DOCUMENT ID		TECN	COMMENT
••	• We do not use the	e follow	ing data for averages	s, fits	s, limits,	etc. • • •
>	3.3	95	<sup>107</sup> CORNET	00	RVUE	evqq′
>50	000		<sup>108</sup> DELGADO	00	RVUE	€K
>	2.6	95	<sup>109</sup> DELGADO	00	RVUE	Electroweak
>	3.3	95	<sup>110</sup> RIZZO	00	RVUE	Electroweak
>	2.9	95	<sup>111</sup> MARCIANO	99	RVUE	Electroweak
>	2.5	95	<sup>112</sup> MASIP	99	RVUE	Electroweak
>	1.6	90	<sup>113</sup> NATH	99	RVUE	Electroweak
>	3.4	95	<sup>114</sup> STRUMIA	99	RVUE	Electroweak

<sup>107</sup> Bound is derived from limits on  $e\nu q q'$  contact interaction, using data from HERA and the Tevatron.

<sup>108</sup> Bound holds only if first two generations of quarks lives on separate branes. If quark mixing is not complex, then bound lowers to 400 TeV from  $\Delta m_K$ .

 $^{109}$  See Figs. 1 and 2 of DELGADO 00 for several model variations. Special boundary conditions can be found which permit KK states down to 950 GeV and that agree with the measurement of  $Q_W({\rm Cs})$ . Quoted bound assumes all Higgs bosons confined to brane; placing one Higgs doublet in the bulk lowers bound to 2.3 TeV.

<sup>110</sup> Bound is derived from global electroweak analysis assuming the Higgs field is trapped on the matter brane. If the Higgs propagates in the bulk, the bound increases to 3.8 TeV.

<sup>111</sup>Bound is derived from global electroweak analysis but considering only presence of the KK W bosons.

 $^{112}$ Global electroweak analysis used to obtain bound independent of position of Higgs on brane or in bulk.

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- <sup>113</sup> Bounds from effect of KK states on  $G_F$ ,  $\alpha$ ,  $M_W$ , and  $M_Z$ . Hard cutoff at string scale determined using gauge coupling unification. Limits for d=2,3,4 rise to 3.5, 5.7, and 7.8 TeV.
- <sup>114</sup> Bound obtained for Higgs confined to the matter brane with  $m_H$ =500 GeV. For Higgs in the bulk, the bound increases to 3.5 TeV.

## LEPTOQUARK QUANTUM NUMBERS

Revised September 2001 by M. Tanabashi (Tohoku University).

Leptoquarks are particles carrying both baryon number (B)and lepton number (L). They are expected to exist in various extensions of the Standard Model (SM). The possible quantum numbers of leptoquark states can be restricted by assuming that their direct interactions with the ordinary SM fermions are dimensionless and invariant under the SM gauge group. Table 1 shows the list of all possible quantum numbers with this assumption [1]. The columns of  $SU(3)_C$ ,  $SU(2)_W$ , and  $U(1)_Y$ in Table 1 indicate the QCD representation, the weak isospin representation, and the weak hypercharge, respectively. The spin of a leptoquark state is taken to be 1 (vector leptoquark) or 0 (scalar leptoquark).

If we do not require leptoquark states to couple directly with SM fermions, different assignments of quantum numbers become possible.

The Pati-Salam model [2] is an example predicting the existence of a leptoquark state. In this model a vector leptoquark appears at the scale where the Pati-Salam SU(4) "color" gauge group breaks into the familiar QCD SU(3)<sub>C</sub> group (or  $SU(3)_C \times U(1)_{B-L}$ ). The Pati-Salam leptoquark is a weak isosinglet and its hypercharge is 2/3. The coupling strength of the Pati-Salam leptoquark is given by the QCD coupling at the Pati-Salam symmetry breaking scale.

Bounds on leptoquark states are obtained both directly and indirectly. Direct limits are from their production cross sections at colliders, while indirect limits are calculated from the bounds

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Spin	3B + L	$\mathrm{SU}(3)_c$	$\mathrm{SU}(2)_W$	$\mathrm{U}(1)_Y$	Allowed coupling
0	-2	$\bar{3}$	1	1/3	$\bar{q}_L^c \ell_L$ or $\bar{u}_R^c e_R$
0	-2	$\bar{3}$	1	4/3	$ar{d}_R^c e_R$
0	-2	$\overline{3}$	3	1/3	$ar{q}_L^c\ell_L$
1	-2	$\bar{3}$	2	5/6	$\bar{q}_L^c \gamma^\mu e_R$ or $\bar{d}_R^c \gamma^\mu \ell_L$
1	-2	$\bar{3}$	2	-1/6	$ar{u}_R^c \gamma^\mu \ell_L$
0	0	3	2	7/6	$\bar{q}_L e_R$ or $\bar{u}_R \ell_L$
0	0	3	2	1/6	$ar{d}_R\ell_L$
1	0	3	1	2/3	$\bar{q}_L \gamma^\mu \ell_L$ or $\bar{d}_R \gamma^\mu e_R$
1	0	3	1	5/3	$ar{u}_R \gamma^\mu e_R$
1	0	3	3	2/3	$ar q_L \gamma^\mu \ell_L$

**Table 1:** Possible leptoquarks and their quantum numbers.

on the leptoquark induced four-fermion interactions which are obtained from low energy experiments.

The pair production cross sections of leptoquarks are evaluated from their interactions with gauge bosons. The gauge couplings of a scalar leptoquark are determined uniquely according to its quantum numbers in Table 1. The magneticdipole-type and the electric-quadrupole-type interactions of a vector leptoquark are, however, not determined even if we fix its gauge quantum numbers as listed in the table [3]. We need extra assumptions about these interactions to evaluate the pair production cross section for a vector leptoquark.

If a leptoquark couples to fermions of more than a single generation in the mass eigenbasis of the SM fermions, it can induce four-fermion interactions causing flavor-changing-neutralcurrents and lepton-family-number violations. Non-chiral leptoquarks, which couple simultaneously to both left- and righthanded quarks, cause four-fermion interactions affecting the

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 $(\pi \to e\nu)/(\pi \to \mu\nu)$  ratio [4]. Indirect limits provide stringent constraints on these leptoquarks. Since the Pati-Salam leptoquark has non-chiral coupling with both e and  $\mu$ , indirect limits from the bounds on  $K_L \to \mu e$  lead to severe bounds on the Pati-Salam leptoquark mass. For detailed bounds obtained in this way, see the Boson Particle Listings for "Indirect Limits for Leptoquarks" and its references.

It is therefore often assumed that a leptoquark state couples only to a single generation in a chiral interaction, where indirect limits become much weaker. This assumption gives strong constraints on concrete models of leptoquarks, however. Leptoquark states which couple only to left- or right-handed quarks are called chiral leptoquarks. Leptoquark states which couple only to the first (second, third) generation are referred as the first (second, third) generation leptoquarks in this section.

### Reference

- W. Buchmüller, R. Rückl, and D. Wyler, Phys. Lett. B191, 442 (1987).
- 2. J.C. Pati and A. Salam, Phys. Rev. D10, 275 (1974).
- J. Blümlein, E. Boos, and A. Kryukov, Z. Phys. C76, 137 (1997).
- 4. O. Shanker, Nucl. Phys. **B204**, 375 (1982).

#### MASS LIMITS for Leptoquarks from Pair Production

These limits rely only on the color or electroweak charge of the leptoquark.

VALUE (GeV)	CL% EVTS	DOCUMENT ID	TECN	COMMENT
>200	95	<sup>115</sup> АВВОТТ	00C D0	Second generation
>148	95	<sup>116</sup> AFFOLDER	00k CDF	Third generation
>202	95	<sup>117</sup> ABE	98s CDF	Second generation
>242	95	<sup>118</sup> GROSS-PILCH	.98	First generation
$\bullet \bullet \bullet$ We do not	use the follow	ing data for averages	s, fits, limits,	etc. • • •
> 98	95	<sup>119</sup> ABAZOV	02 D0	All generatrions
>225	95	<sup>120</sup> ABAZOV	01D D0	First generation
> 85.8	95	<sup>121</sup> ABBIENDI	00м OPAL	First generation
> 85.5	95	<sup>121</sup> ABBIENDI	00м OPAL	Second generation
> 82.7	95	<sup>121</sup> ABBIENDI	00M OPAL	Third generation

>123 >160 >225 > 94 > 99 >213	95 95 95 95 95 95	<ul> <li>122 AFFOLDER</li> <li>123 ABBOTT</li> <li>124 ABBOTT</li> <li>125 ABBOTT</li> <li>126 ABE</li> <li>127 ABE</li> </ul>	00к CDF 99J D0 98E D0 98J D0 97F CDF 97x CDF	Second generation Second generation First generation Third generation Third generation First generation
> 45.5	95	<sup>128,129</sup> ABREU	93J DLPH	First + second genera-
> 44.4 > 44.5 > 45 none 8.9-22.6 none 10.2-23.2 none 5-20.8 none 7-20.5	95 95 95 95 95 95 95	130 ADRIANI 130 ADRIANI 130 DECAMP 131 KIM 131 KIM 132 BARTEL 2 133 BEHREND	93M L3 93M L3 92 ALEP 90 AMY 90 AMY 87B JADE 86B CELL	tion First generation Second generation Third generation First generation Second generation

<sup>115</sup> ABBOTT 00C search for scalar leptoquarks using  $\mu\mu jj$ ,  $\mu\nu jj$ , and  $\nu\nu jj$  events in  $p\overline{p}$  collisions at  $E_{\rm cm}$ =1.8 TeV. The limit above assumes B( $\mu q$ )=1. For B( $\mu q$ )=0.5 and 0, the bound becomes 180 and 79 GeV respectively. Bounds for vector leptoquarks are also given.

<sup>116</sup> AFFOLDER 00K search for scalar leptoquark using  $\nu\nu bb$  events in  $p\overline{p}$  collsions at  $E_{\rm cm}$ =1.8 TeV. The quoted limit assumes B( $\nu b$ )=1. Bounds for vector leptoquarks are also given.

<sup>117</sup> ABE 985 search for scalar leptoquarks using  $\mu \mu j j$  events in  $p \overline{p}$  collisions at  $E_{cm} =$  1.8 TeV. The limit is for B( $\mu q$ )= 1. For B( $\mu q$ )=B( $\nu q$ )=0.5, the limit is > 160 GeV.

<sup>118</sup> GROSS-PILCHER 98 is the combined limit of the CDF and DØ Collaborations as determined by a joint CDF/DØ working group and reported in this FNAL Technical Memo. Original data published in ABE 97X and ABBOTT 98E.

<sup>119</sup> ABAZOV 02 search for scalar leptoquarks using  $\nu \nu j j$  events in  $\overline{p}p$  collisions at  $E_{cm}$ =1.8 TeV. The bound holds for all leptoquark generations. Vector leptoquarks are likewise constrained to lie above 200 GeV.

<sup>120</sup> ABAZOV 01D search for scalar leptoquarks using  $e\nu jj$ , eejj, and  $\nu\nu jj$  events in  $p\overline{p}$  collisions at  $E_{\rm cm}$ =1.8 TeV. The limit above assumes B(eq)=1. For B(eq)=0.5 and 0, the bound becomes 204 and 79 GeV, respectively. Bounds for vector leptoquarks are also given. Supersedes ABBOTT 98E.

<sup>121</sup> ABBIENDI 00M search for scalar/vector leptoquarks in  $e^+e^-$  collisions at  $\sqrt{s}$ =183 GeV. The quoted limits are for charge -4/3 isospin 0 scalar-leptoquarks with B( $\ell q$ )=1. See their Table 8 and Figs. 6–9 for other cases.

<sup>122</sup> AFFOLDER 00K search for scalar leptoquark using  $\nu\nu cc$  events in  $p\overline{p}$  collsions at  $E_{\rm cm}$ =1.8 TeV. The quoted limit assumes B( $\nu c$ )=1. Bounds for vector leptoquarks are also given.

<sup>123</sup> ABBOTT 99J search for leptoquarks using  $\mu \nu j j$  events in  $p \overline{p}$  collisions at  $E_{\rm cm} = 1.8$ TeV. The quoted limit is for a scalar leptoquark with  $B(\mu q) = B(\nu q) = 0.5$ . Limits on vector leptoquarks range from 240 to 290 GeV.

<sup>124</sup> ABBOTT 98E search for scalar leptoquarks using  $e\nu jj$ , eejj, and  $\nu\nu jj$  events in  $p\overline{p}$  collisions at  $E_{\rm cm}$ =1.8 TeV. The limit above assumes B(eq)=1. For B(eq)=0.5 and 0, the bound becomes 204 and 79 GeV, respectively.

<sup>125</sup> ABBOTT 98J search for charge -1/3 third generation scalar and vector leptoquarks in  $p\overline{p}$  collisions at  $E_{cm} = 1.8$  TeV. The quoted limit is for scalar leptoquark with B( $\nu b$ )=1.

- <sup>126</sup>ABE 97F search for third generation scalar and vector leptoquarks in  $p\overline{p}$  collisions at  $E_{\rm cm} = 1.8$  TeV. The quoted limit is for scalar leptoquark with B( $\tau b$ ) = 1.
- <sup>127</sup> ABE 97X search for scalar leptoquarks using eejj events in  $p\overline{p}$  collisions at  $E_{cm}$ =1.8 TeV. The limit is for B(eq)=1.

<sup>128</sup>Limit is for charge -1/3 isospin-0 leptoquark with  $B(\ell q) = 2/3$ .

<sup>129</sup> First and second generation leptoquarks are assumed to be degenerate. The limit is slightly lower for each generation.

- <sup>130</sup> Limits are for charge -1/3, isospin-0 scalar leptoquarks decaying to  $\ell^- q$  or  $\nu q$  with any branching ratio. See paper for limits for other charge-isospin assignments of leptoquarks.
- <sup>131</sup> KIM 90 assume pair production of charge 2/3 scalar-leptoquark via photon exchange. The decay of the first (second) generation leptoquark is assumed to be any mixture of  $de^+$  and  $u\overline{\nu}$  ( $s\mu^+$  and  $c\overline{\nu}$ ). See paper for limits for specific branching ratios.
- <sup>132</sup> BARTEL 87B limit is valid when a pair of charge 2/3 spinless leptoquarks X is produced with point coupling, and when they decay under the constraint  $B(X \rightarrow c\overline{\nu}_{\mu}) + B(X \rightarrow s\mu^+) = 1$ .
- <sup>133</sup> BEHREND 86B assumed that a charge 2/3 spinless leptoquark,  $\chi$ , decays either into  $s\mu^+$  or  $c\overline{\nu}$ : B( $\chi \rightarrow s\mu^+$ ) + B( $\chi \rightarrow c\overline{\nu}$ ) = 1.

#### MASS LIMITS for Leptoquarks from Single Production

These limits depend on the *q*- $\ell$ -leptoquark coupling  $g_{LQ}$ . It is often assumed that  $g_{LQ}^2/4\pi = 1/137$ . Limits shown are for a scalar, weak isoscalar, charge -1/3 lepto-

	quark.					
VALUE	(GeV)	CL%	DOCUMENT ID		TECN	COMMENT
>290		95	<sup>134</sup> ADLOFF	<b>01</b> C	H1	First generation
> 73		95	<sup>135</sup> ABREU	93J	DLPH	Second generation
• • •	We do not use the	followi	ng data for averages	, fits,	limits,	etc. ● ● ●
>197		95	<sup>136</sup> ABBIENDI	<b>0</b> 2B	OPAL	First generation
>204		95	<sup>137</sup> BREITWEG	01	ZEUS	First generation
			<sup>138</sup> BREITWEG	00e	ZEUS	First generation
>161		95	<sup>139</sup> ABREU	<b>99</b> G	DLPH	First generation
>200		95	<sup>140</sup> ADLOFF	99	H1	First generation
			<sup>141</sup> DERRICK	97	ZEUS	Lepton-flavor violation
>168		95	<sup>142</sup> DERRICK	93	ZEUS	First generation

<sup>134</sup> For limits on states with different quantum numbers and the limits in the mass-coupling plane, see their Fig. 3.

- <sup>135</sup>Limit from single production in Z decay. The limit is for a leptoquark coupling of electromagnetic strength and assumes  $B(\ell q) = 2/3$ . The limit is 77 GeV if first and second leptoquarks are degenerate.
- <sup>136</sup> For limits on states with different quantum numbers and the limits in the mass-coupling plane, see their Fig. 4 and Fig. 5.
- <sup>137</sup>See their Fig. 14 for limits in the mass-coupling plane.
- <sup>138</sup> BREITWEG 00E search for F=0 leptoquarks in  $e^+p$  collisions. For limits in masscoupling plane, see their Fig. 11.
- <sup>139</sup> ABREU 99G limit obtained from process  $e\gamma \rightarrow LQ+q$ . For limits on vector and scalar states with different quantum numbers and the limits in the coupling-mass plane, see their Fig. 4 and Table 2.
- <sup>140</sup> For limits on states with different quantum numbers and the limits in the mass-coupling plane, see their Fig. 13 and Fig. 14. ADLOFF 99 also search for leptoquarks with leptonflavor violating couplings. ADLOFF 99 supersedes AID 96B.
- <sup>141</sup> DERRICK 97 search for various leptoquarks with lepton-flavor violating couplings. See their Figs. 5–8 and Table 1 for detailed limits.
- <sup>142</sup> DERRICK 93 search for single leptoquark production in ep collisions with the decay eq and  $\nu q$ . The limit is for leptoquark coupling of electromagnetic strength and assumes  $B(eq) = B(\nu q) = 1/2$ . The limit for B(eq) = 1 is 176 GeV. For limits on states with different quantum numbers, see their Table 3.

Ind	irect Limits for Le	ptoqua	arks			
VALU	UE (TeV)	CL%	DOCUMENT ID		TECN	COMMENT
• •	• We do not use the	followi	ing data for averages	s, fits	, limits,	etc. • • •
>	1.7	95	<sup>143</sup> CHEUNG	<b>01</b> B	RVUE	First generation
>	0.39	95	<sup>144</sup> ACCIARRI	<b>00</b> P	L3	$e^+e^- \rightarrow q q$
>	1.4	95	<sup>145</sup> ADLOFF	00	H1	First generation
>	0.2	95	<sup>146</sup> BARATE	001	ALEP	e <sup>+</sup> e <sup>-</sup>
			<sup>147</sup> BARGER	00	RVUE	Cs
			<sup>148</sup> GABRIELLI	00	RVUE	Lepton flavor violation
>	0.74	95	<sup>149</sup> ZARNECKI		RVUE	$S_{f 1}$ leptoquark
			<sup>150</sup> ABBIENDI		OPAL	
>	19.3	95	<sup>151</sup> ABE	98v	CDF	$B_{m{s}}  o e^{\pm} \mu^{\mp}$ , Pati-Salam type
			<sup>152</sup> ACCIARRI	98J		$e^+e^- \rightarrow q \overline{q}$
			<sup>153</sup> ACKERSTAFF	98v	OPAL	$e^+e^- \rightarrow q \overline{q},$
>	0.76	95	<sup>154</sup> DEANDREA	97	RVUE	$e^+e^-  ightarrow b\overline{b}$ $\widetilde{R}_2$ leptoquark
			<sup>155</sup> DERRICK	97	ZEUS	Lepton-flavor violation
			<sup>156</sup> GROSSMAN	97	RVUE	$B \rightarrow \tau^+ \tau^- (X)$
			<sup>157</sup> JADACH	97	RVUE	$e^+e^- \rightarrow q\overline{q}$
>12	200		<sup>158</sup> KUZNETSOV	<b>95</b> B	RVUE	Pati-Salam type
			<sup>159</sup> MIZUKOSHI	95	RVUE	Third generation scalar leptoquark
>	0.3	95	<sup>160</sup> BHATTACH	94	RVUE	Spin-0 leptoquark cou- pled to e <sub>R</sub> t <sub>I</sub>
			<sup>161</sup> DAVIDSON	94	RVUE	
>	18		<sup>162</sup> KUZNETSOV	94	RVUE	Pati-Salam type
>	0.43	95	<sup>163</sup> LEURER	94	RVUE	First generation spin-1 leptoquark
>	0.44	95	<sup>163</sup> LEURER	<b>94</b> B	RVUE	First generation spin-0 leptoquark
			<sup>164</sup> MAHANTA	94	RVUE	P and $T$ violation
>	1		<sup>165</sup> SHANKER	82	RVUE	Nonchiral spin-0 lepto- quark
> 3	125		<sup>165</sup> SHANKER	82	RVUE	Nonchiral spin-1 lepto- quark

 $^{143}\,\text{CHEUNG 01B}$  quoted limit is for a scalar, weak isoscalar, charge -1/3 leptoquark with a coupling of electromagnetic strength. The limit is derived from bounds on contact interactions in a global electroweak analysis. For the limits of leptoquarks with different quantum numbers, see Table 5.

<sup>144</sup> ACCIARRI 00P limit is for the weak isoscalar spin-0 leptoquark with the coupling of electromagnetic strength. For the limits of leptoquarks with different quantum numbers, see their Table 4.

<sup>145</sup> ADLOFF 00 limit is for the weak isotriplet spin-0 leptoquark at strong coupling,  $\lambda = \sqrt{4\pi}$ . For the limits of leptoquarks with different quantum numbers, see their Table 2. ADLOFF 00 limits are from the  $Q^2$  spectrum measurement of  $e^+ p \rightarrow e^+ X$ .

<sup>146</sup> BARATE 00I search for deviations in cross section and jet-charge asymmetry in  $e^+e^- \rightarrow \overline{q} q$  due to *t*-channel exchange of a leptoquark at  $\sqrt{s}$ =130 to 183 GeV. Limits for other scalar and vector leptoquarks are also given in their Table 22.

<sup>147</sup> BARGER 00 explain the deviation of atomic parity violation in cesium atoms from prediction is explained by scalar leptoquark exchange.

 $^{148}$  GABRIELLI 00 calculate various process with lepton flavor violation in leptoquark models.

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- <sup>149</sup> ZARNECKI 00 limit is derived from data of HERA, LEP, and Tevatron and from various low-energy data including atomic parity violation. Leptoquark coupling with electromagnetic strength is assumed.
- <sup>150</sup> ABBIENDI 99 limits are from  $e^+e^- \rightarrow q \overline{q}$  cross section at 130–136, 161–172, 183 GeV. See their Fig. 8 and Fig. 9 for limits in mass-coupling plane.
- <sup>151</sup> ABE 98V quoted limit is from  $B(B_s \rightarrow e^{\pm}\mu^{\mp}) < 8.2 \times 10^{-6}$ . ABE 98V also obtain a similar limit on  $M_{LQ} > 20.4$  TeV from  $B(B_d \rightarrow e^{\pm}\mu^{\mp}) < 4.5 \times 10^{-6}$ . Both bounds assume the non-canonical association of the *b* quark with electrons or muons under SU(4).
- <sup>152</sup> ACCIARRI 98J limit is from  $e^+e^- \rightarrow q \bar{q}$  cross section at  $\sqrt{s}$ = 130–172 GeV which can be affected by the *t* and *u*-channel exchanges of leptoquarks. See their Fig. 4 and Fig. 5 for limits in the mass-coupling plane.
- <sup>153</sup> ACKERSTAFF 98V limits are from  $e^+e^- \rightarrow q \overline{q}$  and  $e^+e^- \rightarrow b \overline{b}$  cross sections at  $\sqrt{s} = 130-172$  GeV, which can be affected by the *t* and *u*-channel exchanges of leptoquarks. See their Fig. 21 and Fig. 22 for limits of leptoquarks in mass-coupling plane.
- <sup>154</sup> DEANDREA 97 limit is for  $\tilde{R}_2$  leptoquark obtained from atomic parity violation (APV). The coupling of leptoquark is assumed to be electromagnetic strength. See Table 2 for limits of the four-fermion interactions induced by various scalar leptoquark exchange. DEANDREA 97 combines APV limit and limits from Tevatron and HERA. See Fig. 1–4 for combined limits of leptoquark in mass-coupling plane.
- $^{155}$  DERRICK 97 search for lepton-flavor violation in *ep* collision. See their Tables 2–5 for limits on lepton-flavor violating four-fermion interactions induced by various leptoquarks.
- <sup>156</sup> GROSSMAN 97 estimate the upper bounds on the branching fraction  $B \rightarrow \tau^+ \tau^-(X)$  from the absence of the *B* decay with large missing energy. These bounds can be used to constrain leptoquark induced four-fermion interactions.
- <sup>157</sup> JADACH 97 limit is from  $e^+e^- \rightarrow q \overline{q}$  cross section at  $\sqrt{s}=172.3$  GeV which can be affected by the *t* and *u*-channel exchanges of leptoquarks. See their Fig. 1 for limits on vector leptoquarks in mass-coupling plane.
- <sup>158</sup> KUZNETSOV 95B use  $\pi$ , K, B,  $\tau$  decays and  $\mu e$  conversion and give a list of bounds on the leptoquark mass and the fermion mixing matrix in the Pati-Salam model. The quoted limit is from  $K_L \rightarrow \mu e$  decay assuming zero mixing.
- <sup>159</sup> MIZUKOSHI 95 calculate the one-loop radiative correction to the Z-physics parameters in various scalar leptoquark models. See their Fig. 4 for the exclusion plot of third generation leptoquark models in mass-coupling plane.
- <sup>160</sup>BHATTACHARYYA 94 limit is from one-loop radiative correction to the leptonic decay width of the Z.  $m_H$ =250 GeV,  $\alpha_s(m_Z)$ =0.12,  $m_t$ =180 GeV, and the electroweak strength of leptoquark coupling are assumed. For leptoquark coupled to  $\overline{e}_L t_R$ ,  $\overline{\mu} t$ , and  $\overline{\tau} t$ , see Fig. 2 in BHATTACHARYYA 94B erratum and Fig. 3.
- <sup>161</sup> DAVIDSON 94 gives an extensive list of the bounds on leptoquark-induced four-fermion insteractions from  $\pi$ , K, D, B,  $\mu$ ,  $\tau$  decays and meson mixings, *etc.* See Table 15 of DAVIDSON 94 for detail.
- <sup>162</sup> KUZNETSOV 94 gives mixing independent bound of the Pati-Salam leptoquark from the cosmological limit on  $\pi^0 \rightarrow \overline{\nu}\nu$ .
- $^{163}$  LEURER 94, LEURER 94B limits are obtained from atomic parity violation and apply to any chiral leptoquark which couples to the first generation with electromagnetic strength. For a nonchiral leptoquark, universality in  $\pi_{\ell 2}$  decay provides a much more stringent bound.
- <sup>164</sup> MAHANTA 94 gives bounds of P- and T-violating scalar-leptoquark couplings from atomic and molecular experiments.
- <sup>165</sup> From  $(\pi \rightarrow e\nu)/(\pi \rightarrow \mu\nu)$  ratio. SHANKER 82 assumes the leptoquark induced four-fermion coupling  $4g^2/M^2$  ( $\overline{\nu}_{eL} \ u_R$ ) ( $\overline{d}_L e_R$ )with g=0.004 for spin-0 leptoquark and  $g^2/M^2$  ( $\overline{\nu}_{eL} \ \gamma_{\mu} u_L$ ) ( $\overline{d}_R \ \gamma^{\mu} e_R$ ) with  $g\simeq 0.6$  for spin-1 leptoquark.

#### MASS LIMITS for Diquarks

VALUE (GeV)	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the	following	data for averages	s, fits, limits	, etc. ● ● ●
none 290–420		<sup>66</sup> ABE	97G CDF	E <sub>6</sub> diquark
none 15-31.7	95 1	<sup>67</sup> ABREU	940 DLPH	SŬSY <i>E</i> 6 diquark
166				-

<sup>166</sup> ABE 97G search for new particle decaying to dijets.

<sup>167</sup>ABREU 940 limit is from  $e^+e^- \rightarrow \overline{cs}cs$ . Range extends up to 43 GeV if diquarks are degenerate in mass.

#### MASS LIMITS for $g_A$ (axigluon)

Axigluons are massive color-octet gauge bosons in chiral color models and have axialvector coupling to quarks with the same coupling strength as gluons.

VALUE (GeV)	CL%	DOCUMENT ID	0	TECN	COMMENT
• • • We do not use the	followi	ing data for averages	, fits	, limits,	etc. • • •
>365	95	168 DONCHESKI			
none 200–980	95	<sup>169</sup> ABE			$p \overline{p} \rightarrow g_A X, X \rightarrow 2 \text{ jets}$
none 200–870	95	<sup>170</sup> ABE	95N	CDF	$p \overline{p} \rightarrow g_A X, g_A \rightarrow q \overline{q}$
none 240-640	95	<sup>171</sup> ABE	<b>93</b> G	CDF	$p \overline{p} \rightarrow g_A X, g_A \rightarrow 2 jets$
> 50	95	<sup>172</sup> CUYPERS	91	RVUE	$\sigma(e^+e^- \rightarrow hadrons)$
none 120–210	95	<sup>173</sup> ABE	90H	CDF	$p\overline{p} \rightarrow g_A X, g_A \rightarrow 2 \text{ jets}$
> 29		<sup>174</sup> ROBINETT	89	THEO	Partial-wave unitarity
none 150–310	95	<sup>175</sup> ALBAJAR	<b>88</b> B	UA1	$p\overline{p}  o g_{\mathcal{A}} X$ , $g_{\mathcal{A}}  o$ 2jets
> 20		BERGSTROM	88	RVUE	$p \overline{p}  ightarrow \Upsilon X$ via $g_A g$
> 9		<sup>176</sup> CUYPERS	88		
> 25		<sup>177</sup> DONCHESKI	<b>88</b> B	RVUE	$\gamma$ decay

 $^{168}$  DONCHESKI 98 compare  $\alpha_{s}$  derived from low-energy data and that from  $\Gamma(Z$   $\rightarrow$ hadrons)/ $\Gamma(Z \rightarrow \text{ leptons})$ .

 $^{169}$  ABE 97G search for new particle decaying to dijets.

 $^{170}\mathrm{ABE}$  95N assume axigluons decaying to quarks in the Standard Model only.

 $^{171}$  ABE 93G assume  $\Gamma(g_A) = N lpha_s m_{g_A}/6$  with N=10.

 $^{172}$  CUYPERS 91 compare  $lpha_{
m s}$  measured in arY decay and that from R at PEP/PETRA energies.

<sup>173</sup> ABE 90H assumes  $\Gamma(g_A) = N\alpha_s m_{g_A}/6$  with  $N = 5 (\Gamma(g_A) = 0.09 m_{g_A})$ . For N = 10, the excluded region is reduced to 120-150 GeV.

- <sup>174</sup> ROBINETT 89 result demands partial-wave unitarity of J = 0  $t\overline{t} \rightarrow t\overline{t}$  scattering amplitude and derives a limit  $m_{g_A}$  > 0.5  $m_t$ . Assumes  $m_t$  > 56 GeV.
- $^{175}$  ALBAJAR 88B result is from the nonobservation of a peak in two-jet invariant mass distribution.  $\Gamma(g_A) < 0.4 m_{g_A}$  assumed. See also BAGGER 88.
- <sup>176</sup> CUYPERS 88 requires  $\Gamma(\Upsilon \rightarrow gg_A) < \Gamma(\Upsilon \rightarrow ggg)$ . A similar result is obtained by DONCHESKI 88.
- <sup>177</sup> DONCHESKI 88B requires  $\Gamma(\Upsilon \rightarrow g q \overline{q})/\Gamma(\Upsilon \rightarrow g g g) < 0.25$ , where the former decay proceeds via axigluon exchange. A more conservative estimate of < 0.5 leads to  $m_{g_A} > 21$  GeV.

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#### $X^0$ (Heavy Boson) Searches in Z Decays

Searches for radiative transition of Z to a lighter spin-0 state  $X^0$  decaying to hadrons, a lepton pair, a photon pair, or invisible particles as shown in the comments. The limits are for the product of branching ratios.

VALUE		DOCUMENT ID	TECN	COMMENT
• • • We do not use	e the follow	ing data for average	s, fits, limits,	etc. • • •
		<sup>178</sup> BARATE	98∪ ALEP	$X^0 \rightarrow \ell \overline{\ell}, q \overline{q}, g g, \gamma \gamma,$
		<sup>179</sup> ACCIARRI	97Q L3	$X_{cle(s)}^{\nu\overline{ u}}$ invisible parti-
		<sup>180</sup> ACTON		$X^0 \rightarrow \gamma \gamma$
		<sup>181</sup> ABREU	92d DLPH	$X^0  ightarrow$ hadrons
		<sup>182</sup> ADRIANI	92F L3	$X^0 \rightarrow$ hadrons
		<sup>183</sup> ACTON	91 OPAL	$X^0 \rightarrow$ anything
$< 1.1 \times 10^{-4}$	95	<sup>184</sup> ACTON	91b OPAL	$X^0 \rightarrow e^+ e^-$
$< 9 \times 10^{-5}$	95	<sup>184</sup> ACTON	91b OPAL	$X^0 \rightarrow \mu^+ \mu^-$
$< 1.1 \times 10^{-4}$	95	<sup>184</sup> ACTON	91b OPAL	$X^0 \rightarrow \tau^+ \tau^-$
$< 2.8 \times 10^{-4}$	95	<sup>185</sup> ADEVA	91d L3	
$< 2.3 \times 10^{-4}$	95	<sup>185</sup> ADEVA	91d L3	$X^0 \rightarrow \mu^+ \mu^-$
$< 4.7 \times 10^{-4}$	95	<sup>186</sup> ADEVA	91d L3	$X^0 \rightarrow$ hadrons
$< 8 \times 10^{-4}$	95	<sup>187</sup> AKRAWY	90J OPAL	$X^0 \rightarrow$ hadrons
170		0.	· · · · –	

<sup>178</sup>BARATE 98U obtain limits on  $B(Z \rightarrow \gamma X^0)B(X^0 \rightarrow \ell \overline{\ell}, q \overline{q}, g g, \gamma \gamma, \nu \overline{\nu})$ . See their Fig. 17.

- <sup>179</sup>See Fig. 4 of ACCIARRI 97Q for the upper limit on B( $Z \rightarrow \gamma X^0$ ;  $E_{\gamma} > E_{min}$ ) as a function of  $E_{\min}$ .
- <sup>180</sup> ACTON 93E give  $\sigma(e^+e^- \rightarrow X^0\gamma) \cdot B(X^0 \rightarrow \gamma\gamma) < 0.4 \text{ pb } (95\%\text{CL}) \text{ for } m_{\chi 0} = 60 \pm 10^{-10} \text{ m}$ 2.5 GeV. If the process occurs via s-channel  $\gamma$  exchange, the limit translates to  $\Gamma(X^0) \cdot B(X^0 \rightarrow \gamma \gamma)^2 < 20 \text{ MeV for } m_{X^0} = 60 \pm 1 \text{ GeV.}$ <sup>181</sup> ABREU 92D give  $\sigma_Z \cdot B(Z \rightarrow \gamma X^0) \cdot B(X^0 \rightarrow \text{ hadrons}) < (3-10) \text{ pb for } m_{X^0} = 60 \pm 1 \text{ GeV.}$
- 10–78 GeV. A very similar limit is obtained for spin-1  $X^0$ .
- <sup>182</sup> ADRIANI 92F search for isolated  $\gamma$  in hadronic Z decays. The limit  $\sigma_Z \cdot B(Z \rightarrow \gamma X^0)$  $\cdot$  B(X^0 \rightarrow \, hadrons) <(2–10) pb (95%CL) is given for  $m_{\chi^0}=$  25–85 GeV.
- <sup>183</sup> ACTON 91 searches for  $Z \rightarrow Z^* X^0$ ,  $Z^* \rightarrow e^+ e^-$ ,  $\mu^+ \mu^-$ , or  $\nu \overline{\nu}$ . Excludes any new scalar  $X^0$  with  $m_{\chi 0} < 9.5 \text{ GeV}/c$  if it has the same coupling to  $ZZ^*$  as the MSM Higgs boson.
- <sup>184</sup> ACTON 91B limits are for  $m_{\chi^0} = 60-85$  GeV.
- $^{185}$  ADEVA 91D limits are for  $m_{\chi 0} =$  30–89 GeV.
- $^{186}$  ADEVA 91D limits are for  $m_{\chi 0}$  = 30–86 GeV.
- <sup>187</sup> AKRAWY 90J give  $\Gamma(Z \rightarrow \gamma X^0) \cdot B(X^0 \rightarrow \text{hadrons}) < 1.9 \text{ MeV}$  (95%CL) for  $m_{\chi 0}$ = 32–80 GeV. We divide by  $\Gamma(Z) = 2.5$  GeV to get product of branching ratios. For nonresonant transitions, the limit is  $B(Z \rightarrow \gamma q \overline{q}) < 8.2$  MeV assuming three-body phase space distribution.

		ivy neutral D050		uping	
VALUE (GeV)	<u>CL%</u>	DOCUMENT ID		TECN	COMMENT
• • • We do not use	the foll	owing data for avera	ages,	fits, limi	ts, etc. ● ● ●
none 55–61		<sup>188</sup> ODAKA	89	VNS	$\Gamma(X^0 \rightarrow e^+e^-)$
					$\cdot B(X^0 \rightarrow \text{hadrons}) \gtrsim 0.2 \text{ MeV}$
>45	95	<sup>189</sup> DERRICK	86	HRS	0.2  MeV $\Gamma(X^0 \rightarrow e^+ e^-) = 6 \text{ MeV}$
>46.6	95	<sup>190</sup> ADEVA			$\Gamma(X^0  ightarrow e^+ e^-) = 10  ext{ keV}$
>48	95	<sup>190</sup> ADEVA	85	MRKJ	$\Gamma(X^0 \rightarrow e^+e^-) = 4 \text{ MeV}$
		<sup>191</sup> BERGER		PLUT	
none 39.8–45.5		<sup>192</sup> ADEVA	84	MRKJ	$\Gamma(X^0_{-} \rightarrow e^+e^-)=10 \text{ keV}$
>47.8	95	<sup>192</sup> ADEVA	84	MRKJ	$\Gamma(X^0 \rightarrow e^+ e^-) = 4 \text{ MeV}$
none 39.8–45.2		<sup>192</sup> BEHREND		CELL	
>47	95	<sup>192</sup> BEHREND	84C	CELL	$\Gamma(X^0 \rightarrow e^+e^-) = 4 \text{ MeV}$

MASS LIMITS for a Heav	y Neutral Boson	Coupling to $e^+e^-$
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<sup>188</sup> ODAKA 89 looked for a narrow or wide scalar resonance in  $e^+e^- \rightarrow \text{hadrons at } E_{\text{cm}} = 55.0-60.8 \text{ GeV}$ .

<sup>189</sup> DERRICK 86 found no deviation from the Standard Model Bhabha scattering at  $E_{cm}$ = 29 GeV and set limits on the possible scalar boson  $e^+e^-$  coupling. See their figure 4 for excluded region in the  $\Gamma(X^0 \rightarrow e^+e^-) - m_{X^0}$  plane. Electronic chiral invariance requires a parity doublet of  $X^0$ , in which case the limit applies for  $\Gamma(X^0 \rightarrow e^+e^-) =$ 3 MeV.

<sup>190</sup> ADEVA 85 first limit is from  $2\gamma$ ,  $\mu^+ \mu^-$ , hadrons assuming  $X^0$  is a scalar. Second limit is from  $e^+ e^-$  channel.  $E_{\rm cm} = 40-47$  GeV. Supersedes ADEVA 84.

<sup>191</sup> BERGER 85B looked for effect of spin-0 boson exchange in  $e^+e^- \rightarrow e^+e^-$  and  $\mu^+\mu^-$  at  $E_{\rm cm} = 34.7$  GeV. See Fig. 5 for excluded region in the  $m_{\chi^0} - \Gamma(X^0)$  plane.

<sup>192</sup> ADEVA 84 and BEHREND 84C have  $E_{\rm cm} = 39.8-45.5$  GeV. MARK-J searched  $X^0$  in  $e^+e^- \rightarrow {\rm hadrons}$ ,  $2\gamma$ ,  $\mu^+\mu^-$ ,  $e^+e^-$  and CELLO in the same channels plus  $\tau$  pair. No narrow or broad  $X^0$  is found in the energy range. They also searched for the effect of  $X^0$  with  $m_X > E_{\rm cm}$ . The second limits are from Bhabha data and for spin-0 singlet. The same limits apply for  $\Gamma(X^0 \rightarrow e^+e^-) = 2$  MeV if  $X^0$  is a spin-0 doublet. The second limit of BEHREND 84C was read off from their figure 2. The original papers also list limits in other channels.

#### Search for $X^0$ Resonance in $e^+e^-$ Collisions

The limit is for  $\Gamma(X^0 \to e^+e^-) \cdot B(X^0 \to f)$ , where f is the specified final state. Spin 0 is assumed for  $X^0$ .

VALUE (keV)	<u>CL%</u>	DOCUMENT ID	7	TECN	COMMENT
$\bullet \bullet \bullet$ We do not use the	e following o	lata for averages	s, fits, I	limits, o	etc. ● ● ●
<10 <sup>3</sup>		<sup>3</sup> ABE	93c \	/NS	Γ(ee)
<(0.4–10)		<sup>1</sup> ABE	93c \	/NS	$f = \gamma \gamma$
<(0.3–5)	95 195,196		93d 1	ΓΟΡΖ	$f = \gamma \gamma$
<(2–12)	95 195,196		93d 1	ΓΟΡΖ	f = hadrons
<(4–200)	95 196,197	· ·	93d 1	ΓΟΡΖ	f = e e
<(0.1–6)	95 196,197		93d T	ΓΟΡΖ	$f = \mu \mu$
<(0.5–8)	90 198	<sup>3</sup> STERNER	93 A	λMY	$f=\gamma\gamma$

<sup>193</sup>Limit is for  $\Gamma(X^0 \to e^+e^-) m_{\chi 0} = 56-63.5$  GeV for  $\Gamma(X^0) = 0.5$  GeV.

 $^{194}$  Limit is for  $m_{\chi 0} =$  56–61.5 GeV and is valid for  $\Gamma(X^0) \ll$  100 MeV. See their Fig. 5 for limits for  $\Gamma = 1,2$  GeV.

<sup>195</sup>Limit is for  $m_{\chi^0} = 57.2-60$  GeV.

<sup>196</sup>Limit is valid for  $\Gamma(X^0) \ll 100$  MeV. See paper for limits for  $\Gamma = 1$  GeV and those for J = 2 resonances. <sup>197</sup>Limit is for  $m_{\chi^0} = 56.6-60$  GeV.

 $^{198}$  STERNER 93 limit is for  $m_{\chi 0}$  = 57–59.6 GeV and is valid for  $\Gamma(X^0){<}100$  MeV. See their Fig. 2 for limits for  $\Gamma = 1,3$  GeV.

#### Search for $X^0$ Resonance in Two-Photon Process

The limit is for $\Gamma(X^0) \cdot B(X^0  o \gamma \gamma)^2$ . Spin 0 is assumed for $X^0$ .						
VALUE (MeV)	<u>CL%</u>	DOCUMENT ID	TECN	COMMENT		
• • • We do not use	e the followin	g data for average	s, fits, limits,	etc. • • •		
<2.6	95	<sup>199</sup> ACTON	93e OPAL	$m_{oldsymbol{\chi}0}=$ 60 $\pm$ 1 GeV		
<2.9	95	BUSKULIC	93F ALEP	$m_{\chi^0}^{\prime}\sim$ 60 GeV		
<sup>199</sup> ACTON 93E limit for a $J = 2$ resonance is 0.8 MeV.						

Search for $X^0$ Resonance in $e^+e^- \rightarrow X^0\gamma$					
VALUE (GeV)	DOCUMENT ID	TECN	COMMENT		
• • • We do not use the follow	<sup>200</sup> ABREU	00z DLPH	etc. • • • $X^0$ decaying invisibly $X^0$ decaying invisibly		
200					

<sup>200</sup> ABREU 00Z is from the single photon cross section at  $\sqrt{s}$ =183, 189 GeV. The production cross section upper limit is less than 0.3 pb for  $X^0$  mass between 40 and 160 GeV. See their Fig. 4 for the limit in mass-cross section plane.

 $^{201}$  ADAM 96C is from the single photon production cross at  $\sqrt{s}$ =130, 136 GeV. The upper bound is less than 3 pb for  $X^0$  masses between 60 and 130 GeV. See their Fig. 5 for the exact bound on the cross section  $\sigma(e^+e^- \rightarrow \gamma X^0)$ .

Search for  $X^0$  Resonance in  $Z \to f\overline{f}X^0$ The limit is for  $B(Z \to f\overline{f}X^0) \cdot B(X^0 \to F)$  where f is a fermion and F is the specified final state. Spin 0 is assumed for  $X^0$ 

VALUE	<u>CL%</u>	DOCUMENT ID	TECN	COMMENT
• • • We do not use t	he follow	ing data for averages	s, fits, limits,	etc. • • •
		<sup>202</sup> ABREU	96⊤ DLPH	$f = e, \mu, \tau; F = \gamma \gamma$
$< 3.7 \times 10^{-6}$	95	<sup>203</sup> ABREU	96⊤ DLPH	$f=\nu; F=\gamma\gamma$
		<sup>204</sup> ABREU	96⊤ DLPH	$f=q; F=\gamma\gamma$
$< 6.8 \times 10^{-6}$	95	<sup>203</sup> ACTON	93e OPAL	$f = e, \mu, \tau; F = \gamma \gamma$
$< 5.5 \times 10^{-6}$	95	<sup>203</sup> ACTON	93e OPAL	$f = q; F = \gamma \gamma$
$< 3.1 \times 10^{-6}$	95	<sup>203</sup> ACTON	93e OPAL	$f = \nu; F = \gamma \gamma$
$< 6.5  imes 10^{-6}$	95	<sup>203</sup> ACTON	93e OPAL	$f = e, \mu; F = \ell \overline{\ell}, q \overline{q}, \nu \overline{\nu}$
$< 7.1 \times 10^{-6}$	95	<sup>203</sup> BUSKULIC	93F ALEP	$f = e, \mu; F = \ell \overline{\ell}, q \overline{q}, \nu \overline{\nu}$
		<sup>205</sup> ADRIANI	92F L3	$f = q; F = \gamma \gamma$

 $^{202}\mathrm{ABREU}$  96T obtain limit as a function of  $m_{\chi0}.$  See their Fig. 6.

 $^{203}\,{\rm Limit}$  is for  $m_{\chi 0}$  around 60 GeV.

 $^{204}\,{\rm ABREU}$  96T obtain limit as a function of  $m_{\chi0}.$  See their Fig. 15.

<sup>205</sup> ADRIANI 92F give  $\sigma_Z \cdot B(Z \rightarrow q \overline{q} X^0) \cdot B(X^0 \rightarrow \gamma \gamma) < (0.75-1.5) \text{ pb} (95\% \text{CL}) \text{ for } m_{X^0} = 10-70 \text{ GeV}$ . The limit is 1 pb at 60 GeV.

Search for $X^0$ Reso	nance in $p\overline{p} \to WX^0$	
VALUE (MeV)	DOCUMENT ID	D <u>TECN</u> <u>COMMENT</u>
• • • We do not use t	he following data for average	es, fits, limits, etc. • • •
	<sup>206</sup> ABE	97w CDF $X^0 \rightarrow b\overline{b}$
206	0	

<sup>206</sup>ABE 97W search for  $X^0$  production associated with W in  $p\overline{p}$  collisions at  $E_{\rm cm}$ =1.8 TeV. The 95%CL upper limit on the production cross section times the branching ratio for  $X^0 \rightarrow b\overline{b}$  ranges from 14 to 19 pb for  $X^0$  mass between 70 and 120 GeV. See their Fig. 3 for upper limits of the production cross section as a function of  $m_{\chi^0}$ .

#### Heavy Particle Production in Quarkonium Decays

Limits are for bran	ching r	atios to modes shown.		
VALUE	<u>CL%</u>	DOCUMENT ID	TECN	COMMENT
• • • We do not use the	e follow	ing data for averages, fits	s, limits,	etc. • • •
$< 1.5 \times 10^{-5}$	90	<sup>207</sup> BALEST 95	CLE2	$egin{array}{ll} \Upsilon(1S) & ightarrow X^{m 0} \gamma, \ m_{m  au 0} < 5  { m GeV} \end{array}$
$< 3  imes 10^{-5} - 6  imes 10^{-3}$	90	<sup>208</sup> BALEST 95	CLE2	$m_{\chi^0} < 5 \text{ GeV}$ $\Upsilon(15)  ightarrow X^0 \overline{X}^0 \gamma,$ $m_{\chi^0} < 3.9 \text{ GeV}$
$< 5.6 \times 10^{-5}$	90	<sup>209</sup> ANTREASYAN 900	CBAL	$ec{\gamma} \stackrel{\chi_0}{(1S)}  o X^0 \gamma, \ m_{\chi^0} < 7.2 \ { m GeV}$
		<sup>210</sup> ALBRECHT 89	ARG	Χ°

 $^{207}$  BALEST 95 two-body limit is for pseudoscalar  $X^0$ . The limit becomes  $< 10^{-4}$  for  $m_{\chi^0} < 7.7 \text{ GeV}.$ 

<sup>208</sup> BALEST 95 three-body limit is for phase-space photon energy distribution and angular distribution same as for  $\Upsilon \to gg\gamma$ . 209 ANTREASYAN 90C assume that  $X^0$  does not decay in the detector.

<sup>210</sup>ALBRECHT 89 give limits for B( $\Upsilon(1S), \Upsilon(2S) \rightarrow X^0 \gamma$ )·B( $X^0 \rightarrow \pi^+ \pi^-, K^+ K^-, K^- K^$  $p\overline{p}$ ) for  $m_{\chi 0}$  < 3.5 GeV.

#### **REFERENCES FOR Searches for Heavy Bosons Other Than Higgs Bosons**

ABBOTT 00C PRL 84 2088 B. Abbott <i>et al.</i> (D0 Collab	ab.)́
AFFOLDER         011         PRL         87         231803         T.         Affolder et al.         (CDF         Collab           BREITWEG         01         PR         D63         052002         J.         Breitweg et al.         (ZEUS         Collab           CHEUNG         01B         PL         B517         167         K.         Cheung         (CDF         Collab           THOMAS         01         NP         A694         559         E.         Thomas et al.         (OPAL         Collab           ABBIENDI         00M         EPJ         C13         15         G.         Abbiendi et al.         (DO         Collab           ABBOTT         00C         PRL         84         2088         B.         Abbott et al.         (DO         Collab	,
CHEUNG         01B         PL         B517         167         K. Cheung           THOMAS         01         NP         A694         559         E. Thomas et al.           ABBIENDI         00M         EPJ         C13         15         G. Abbiendi et al.         (OPAL Collab	зb.)
THOMAS01NPA694559E.Thomas et al.ABBIENDI00MEPJC1315G.Abbiendi et al.(OPAL Collab	зb.)
ABBIENDI 00M EPJ C13 15 G. Abbiendi et al. (OPAL Collab	
ABBOTT 00C PRL 84 2088 B. Abbott <i>et al.</i> (D0 Collab	зb.)
	зb.)
ABE 00 PRL 84 5716 F. Abe <i>et al.</i> (CDF Collab	ab.)
ABREU 00S PL B485 45 P. Abreu et al. (DELPHI Collab	ab.)
ABREU 00Z EPJ C17 53 P. Abreu <i>et al.</i> (DELPHI Collab	ab.)

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	000	DI D400 01	M Assister at al	(1.2  Callab)
ACCIARRI	00P	PL B489 81	M. Acciarri <i>et al.</i>	(L3 Collab.)
ADLOFF	00	PL B479 358	C. Adloff <i>et al.</i>	(H1 Collab.)
AFFOLDER	00K	PRL 85 2056	T. Affolder <i>et al.</i>	(CDF Collab.)
BARATE	001	EPJ C12 183	R. Barate <i>et al.</i>	(ALEPH Collab.)
BARGER	00	PL B480 149	V. Barger, K. Cheung	
BREITWEG	00E	EPJ C16 253	J. Breitweg <i>et al.</i>	(ZEUS Collab.)
CHAY	00	PR D61 035002	J. Chay, K.Y. Lee, S. Nam	(2200 00000)
СНО				
	00	MPL A15 311	G. Cho	
CORNET	00	PR D61 037701	F. Cornet, M. Relano, J. Rico	
DELGADO	00	JHEP 0001 030	A. Delgado, A. Pomarol, M. Quiros	
ERLER	00	PRL 84 212	J. Erler, P. Langacker	
GABRIELLI	00	PR D62 055009	E. Gabrielli	
RIZZO	00	PR D61 016007	T.G. Rizzo, J.D. Wells	
-				
ROSNER	00	PR D61 016006	J.L. Rosner	
ZARNECKI	00	EPJ C17 695	A. Zarnecki	
ABBIENDI	99	EPJ C6 1	G. Abbiendi <i>et al.</i>	(OPAL Collab.)
ABBOTT	99 J	PRL 83 2896	B. Abbott <i>et al.</i>	(D0 Collab.)
ABREU	99G	PL B446 62	P. Abreu <i>et al.</i>	(DELPHI Collab.)
ACKERSTAFF	99D	EPJ C8 3	K. Ackerstaff <i>et al.</i>	
				(OPAL Collab.)
ADLOFF	99	EPJ C11 447	C. Adloff <i>et al.</i>	(H1 Collab.)
Also	00C	EPJ C14 553 errata	C. Adloff <i>et al.</i>	(H1 Collab.)
CASALBUONI	99	PL B460 135	R. Casalbuoni <i>et al.</i>	
CZAKON	99	PL B458 355	M. Czakon, J. Gluza, M. Zralek	
ERLER	99	PL B456 68	J. Erler, P. Langacker	
MARCIANO	99	PR D60 093006	W. Marciano	
MASIP	99	PR D60 096005		
			M. Masip, A. Pomarol	
NATH	99	PR D60 116004	P. Nath, M. Yamaguchi	
STRUMIA	99	PL B466 107	A. Strumia	
ABBOTT	98E	PRL 80 2051	B. Abbott <i>et al.</i>	(D0 Collab.)
ABBOTT	98J	PRL 81 38	B. Abbott <i>et al.</i>	(D0 Collab.)
ABE	98S	PRL 81 4806	F. Abe <i>et al.</i>	(CDF Collab.)
ABE	98V	PRL 81 5742	F. Abe <i>et al.</i>	(CDF Collab.)
ACCIARRI	98J	PL B433 163	M. Acciarri <i>et al.</i>	(L3 Collab.)
ACKERSTAFF	98V	EPJ C2 441	K. Ackerstaff <i>et al.</i>	(OPAL Collab.)
BARATE	98U	EPJ C4 571	R. Barate <i>et al.</i>	(ALEPH Collab.)
BARENBOIM	98	EPJ C1 369	G. Barenboim	
СНО	98	EPJ C5 155	G. Cho, K. Hagiwara, S. Matsumoto	
CONRAD	98	RMP 70 1341	J.M. Conrad, M.H. Shaevitz, T. Boltor	
CONTORD				
	00			
DONCHESKI	98	PR D58 097702	M.A. Doncheski, R.W. Robinett	
GROSS-PILCH.	98	hep-ex/9810015	C. Grosso-Pilcher, G. Landsberg, M. Pa	
				aterno (CDF Collab.)
GROSS-PILCH.	98	hep-ex/9810015	C. Grosso-Pilcher, G. Landsberg, M. Pa	
GROSS-PILCH. ABE	98 97F	hep-ex/9810015 PRL 78 2906	C. Grosso-Pilcher, G. Landsberg, M. Pa F. Abe <i>et al.</i>	(CDF Collab.)
GROSS-PILCH. ABE ABE ABE	98 97F 97G 97S	hep-ex/9810015 PRL 78 2906 PR D55 R5263 PRL 79 2192	C. Grosso-Pilcher, G. Landsberg, M. Pa F. Abe <i>et al.</i> F. Abe <i>et al.</i>	(CDF Collab.) (CDF Collab.) (CDF Collab.)
GROSS-PILCH. ABE ABE ABE ABE	98 97F 97G 97S 97W	hep-ex/9810015 PRL 78 2906 PR D55 R5263 PRL 79 2192 PRL 79 3819	C. Grosso-Pilcher, G. Landsberg, M. Pa F. Abe <i>et al.</i> F. Abe <i>et al.</i> F. Abe <i>et al.</i> F. Abe <i>et al.</i> F. Abe <i>et al.</i>	(CDF Collab.) (CDF Collab.) (CDF Collab.) (CDF Collab.)
GROSS-PILCH. ABE ABE ABE ABE ABE	98 97F 97G 97S 97W 97W 97X	hep-ex/9810015 PRL 78 2906 PR D55 R5263 PRL 79 2192 PRL 79 3819 PRL 79 4327	<ul> <li>C. Grosso-Pilcher, G. Landsberg, M. Pa</li> <li>F. Abe <i>et al.</i></li> </ul>	(CDF Collab.) (CDF Collab.) (CDF Collab.) (CDF Collab.) (CDF Collab.)
GROSS-PILCH. ABE ABE ABE ABE ABE ACCIARRI	98 97F 97G 97S 97W 97W 97X 97Q	hep-ex/9810015 PRL 78 2906 PR D55 R5263 PRL 79 2192 PRL 79 3819 PRL 79 4327 PL B412 201	<ul> <li>C. Grosso-Pilcher, G. Landsberg, M. Pa</li> <li>F. Abe <i>et al.</i></li> <li>M. Acciarri <i>et al.</i></li> </ul>	(CDF Collab.) (CDF Collab.) (CDF Collab.) (CDF Collab.) (CDF Collab.) (L3 Collab.)
GROSS-PILCH. ABE ABE ABE ABE ACCIARRI ARIMA	98 97F 97G 97S 97W 97W 97X 97Q 97	hep-ex/9810015 PRL 78 2906 PR D55 R5263 PRL 79 2192 PRL 79 3819 PRL 79 4327 PL B412 201 PR D55 19	<ul> <li>C. Grosso-Pilcher, G. Landsberg, M. Pa</li> <li>F. Abe et al.</li> <li>F. Abe et al.</li> <li>F. Abe et al.</li> <li>F. Abe et al.</li> <li>M. Acciarri et al.</li> <li>T. Arima et al.</li> </ul>	(CDF Collab.) (CDF Collab.) (CDF Collab.) (CDF Collab.) (CDF Collab.) (L3 Collab.) (VENUS Collab.)
GROSS-PILCH. ABE ABE ABE ABE ACCIARRI ACCIARRI ARIMA BARENBOIM	98 97F 97G 97S 97W 97W 97X 97Q 97 97	hep-ex/9810015 PRL 78 2906 PR D55 R5263 PRL 79 2192 PRL 79 3819 PRL 79 4327 PL B412 201 PR D55 19 PR D55 4213	<ul> <li>C. Grosso-Pilcher, G. Landsberg, M. Pa</li> <li>F. Abe et al.</li> <li>F. Abe et al.</li> <li>F. Abe et al.</li> <li>F. Abe et al.</li> <li>M. Acciarri et al.</li> <li>T. Arima et al.</li> <li>G. Barenboim et al.</li> </ul>	(CDF Collab.) (CDF Collab.) (CDF Collab.) (CDF Collab.) (CDF Collab.) (L3 Collab.) (VENUS Collab.) (VALE, IFIC)
GROSS-PILCH. ABE ABE ABE ABE ACCIARRI ARIMA	98 97F 97G 97S 97W 97W 97X 97Q 97	hep-ex/9810015 PRL 78 2906 PR D55 R5263 PRL 79 2192 PRL 79 3819 PRL 79 4327 PL B412 201 PR D55 19	<ul> <li>C. Grosso-Pilcher, G. Landsberg, M. Pa</li> <li>F. Abe et al.</li> <li>F. Abe et al.</li> <li>F. Abe et al.</li> <li>F. Abe et al.</li> <li>M. Acciarri et al.</li> <li>T. Arima et al.</li> </ul>	(CDF Collab.) (CDF Collab.) (CDF Collab.) (CDF Collab.) (CDF Collab.) (L3 Collab.) (VENUS Collab.)
GROSS-PILCH. ABE ABE ABE ABE ACCIARRI ACCIARRI ARIMA BARENBOIM	98 97F 97G 97S 97W 97W 97X 97Q 97 97	hep-ex/9810015 PRL 78 2906 PR D55 R5263 PRL 79 2192 PRL 79 3819 PRL 79 4327 PL B412 201 PR D55 19 PR D55 4213	<ul> <li>C. Grosso-Pilcher, G. Landsberg, M. Pa</li> <li>F. Abe et al.</li> <li>F. Abe et al.</li> <li>F. Abe et al.</li> <li>F. Abe et al.</li> <li>M. Acciarri et al.</li> <li>T. Arima et al.</li> <li>G. Barenboim et al.</li> </ul>	(CDF Collab.) (CDF Collab.) (CDF Collab.) (CDF Collab.) (CDF Collab.) (L3 Collab.) (VENUS Collab.) (VALE, IFIC) (MARS)
GROSS-PILCH. ABE ABE ABE ABE ACCIARRI ARIMA BARENBOIM DEANDREA DERRICK	98 97F 97G 97S 97W 97X 97Q 97 97 97 97	hep-ex/9810015 PRL 78 2906 PR D55 R5263 PRL 79 2192 PRL 79 3819 PRL 79 4327 PL B412 201 PR D55 19 PR D55 19 PR D55 4213 PL B409 277 ZPHY C73 613	<ul> <li>C. Grosso-Pilcher, G. Landsberg, M. Pa</li> <li>F. Abe et al.</li> <li>F. Abe et al.</li> <li>F. Abe et al.</li> <li>F. Abe et al.</li> <li>M. Acciarri et al.</li> <li>M. Acciarri et al.</li> <li>G. Barenboim et al.</li> <li>A. Deandrea</li> <li>M. Derrick et al.</li> </ul>	(CDF Collab.) (CDF Collab.) (CDF Collab.) (CDF Collab.) (CDF Collab.) (L3 Collab.) (VENUS Collab.) (VALE, IFIC) (MARS) (ZEUS Collab.)
GROSS-PILCH. ABE ABE ABE ABE ACCIARRI ARIMA BARENBOIM DEANDREA DERRICK GROSSMAN	98 97F 97G 97S 97W 97X 97Q 97 97 97 97 97	hep-ex/9810015 PRL 78 2906 PR D55 R5263 PRL 79 2192 PRL 79 3819 PRL 79 4327 PL B412 201 PR D55 19 PR D55 4213 PL B409 277 ZPHY C73 613 PR D55 2768	<ul> <li>C. Grosso-Pilcher, G. Landsberg, M. Pa</li> <li>F. Abe et al.</li> <li>M. Acciarri et al.</li> <li>M. Acciarri et al.</li> <li>G. Barenboim et al.</li> <li>A. Deandrea</li> <li>M. Derrick et al.</li> <li>Y. Grossman, Z. Ligeti, E. Nardi</li> </ul>	(CDF Collab.) (CDF Collab.) (CDF Collab.) (CDF Collab.) (CDF Collab.) (L3 Collab.) (VENUS Collab.) (VALE, IFIC) (MARS) (ZEUS Collab.) (REHO, CIT)
GROSS-PILCH. ABE ABE ABE ABE ACCIARRI ARIMA BARENBOIM DEANDREA DERRICK GROSSMAN JADACH	98 97F 97G 97S 97W 97X 97Q 97 97 97 97 97 97 97 97	hep-ex/9810015 PRL 78 2906 PR D55 R5263 PRL 79 2192 PRL 79 3819 PRL 79 4327 PL B412 201 PR D55 19 PR D55 4213 PL B409 277 ZPHY C73 613 PR D55 2768 PL B408 281	<ul> <li>C. Grosso-Pilcher, G. Landsberg, M. Pa</li> <li>F. Abe et al.</li> <li>M. Acciarri et al.</li> <li>G. Barenboim et al.</li> <li>A. Deandrea</li> <li>M. Derrick et al.</li> <li>Y. Grossman, Z. Ligeti, E. Nardi</li> <li>S. Jadach, B.F.L. Ward, Z. Was</li> </ul>	(CDF Collab.) (CDF Collab.) (CDF Collab.) (CDF Collab.) (CDF Collab.) (L3 Collab.) (VENUS Collab.) (VALE, IFIC) (MARS) (ZEUS Collab.) (REHO, CIT) (CERN, INPK+)
GROSS-PILCH. ABE ABE ABE ABE ACCIARRI ARIMA BARENBOIM DEANDREA DERRICK GROSSMAN JADACH STAHL	98 97F 97G 97S 97W 97X 97Q 97 97 97 97 97 97 97 97	hep-ex/9810015 PRL 78 2906 PR D55 R5263 PRL 79 2192 PRL 79 3819 PRL 79 4327 PL B412 201 PR D55 19 PR D55 4213 PL B409 277 ZPHY C73 613 PR D55 2768 PL B408 281 ZPHY C74 73	<ul> <li>C. Grosso-Pilcher, G. Landsberg, M. Pa</li> <li>F. Abe et al.</li> <li>M. Acciarri et al.</li> <li>T. Arima et al.</li> <li>G. Barenboim et al.</li> <li>A. Deandrea</li> <li>M. Derrick et al.</li> <li>Y. Grossman, Z. Ligeti, E. Nardi</li> <li>S. Jadach, B.F.L. Ward, Z. Was</li> <li>A. Stahl, H. Voss</li> </ul>	(CDF Collab.) (CDF Collab.) (CDF Collab.) (CDF Collab.) (CDF Collab.) (L3 Collab.) (VENUS Collab.) (VALE, IFIC) (MARS) (ZEUS Collab.) (REHO, CIT) (CERN, INPK+) (BONN)
GROSS-PILCH. ABE ABE ABE ACCIARRI ARIMA BARENBOIM DEANDREA DERRICK GROSSMAN JADACH STAHL ABACHI	98 97F 97G 97S 97W 97X 97Q 97 97 97 97 97 97 97 97 97	hep-ex/9810015 PRL 78 2906 PR D55 R5263 PRL 79 2192 PRL 79 3819 PRL 79 4327 PL B412 201 PR D55 19 PR D55 4213 PL B409 277 ZPHY C73 613 PR D55 2768 PL B408 281 ZPHY C74 73 PRL 76 3271	<ul> <li>C. Grosso-Pilcher, G. Landsberg, M. Pa</li> <li>F. Abe et al.</li> <li>M. Acciarri et al.</li> <li>T. Arima et al.</li> <li>G. Barenboim et al.</li> <li>A. Deandrea</li> <li>M. Derrick et al.</li> <li>Y. Grossman, Z. Ligeti, E. Nardi</li> <li>S. Jadach, B.F.L. Ward, Z. Was</li> <li>A. Stahl, H. Voss</li> <li>S. Abachi et al.</li> </ul>	(CDF Collab.) (CDF Collab.) (CDF Collab.) (CDF Collab.) (CDF Collab.) (L3 Collab.) (VENUS Collab.) (VALE, IFIC) (MARS) (ZEUS Collab.) (REHO, CIT) (CERN, INPK+) (BONN) (D0 Collab.)
GROSS-PILCH. ABE ABE ABE ABE ACCIARRI ARIMA BARENBOIM DEANDREA DERRICK GROSSMAN JADACH STAHL	98 97F 97G 97S 97W 97X 97Q 97 97 97 97 97 97 97 97	hep-ex/9810015 PRL 78 2906 PR D55 R5263 PRL 79 2192 PRL 79 3819 PRL 79 4327 PL B412 201 PR D55 19 PR D55 4213 PL B409 277 ZPHY C73 613 PR D55 2768 PL B408 281 ZPHY C74 73	<ul> <li>C. Grosso-Pilcher, G. Landsberg, M. Pa</li> <li>F. Abe et al.</li> <li>M. Acciarri et al.</li> <li>T. Arima et al.</li> <li>G. Barenboim et al.</li> <li>A. Deandrea</li> <li>M. Derrick et al.</li> <li>Y. Grossman, Z. Ligeti, E. Nardi</li> <li>S. Jadach, B.F.L. Ward, Z. Was</li> <li>A. Stahl, H. Voss</li> <li>S. Abachi et al.</li> <li>S. Abachi et al.</li> </ul>	(CDF Collab.) (CDF Collab.) (CDF Collab.) (CDF Collab.) (CDF Collab.) (L3 Collab.) (VENUS Collab.) (VALE, IFIC) (MARS) (ZEUS Collab.) (REHO, CIT) (CERN, INPK+) (BONN)
GROSS-PILCH. ABE ABE ABE ACCIARRI ARIMA BARENBOIM DEANDREA DERRICK GROSSMAN JADACH STAHL ABACHI	98 97F 97G 97S 97W 97X 97Q 97 97 97 97 97 97 97 97 97	hep-ex/9810015 PRL 78 2906 PR D55 R5263 PRL 79 2192 PRL 79 3819 PRL 79 4327 PL B412 201 PR D55 19 PR D55 4213 PL B409 277 ZPHY C73 613 PR D55 2768 PL B408 281 ZPHY C74 73 PRL 76 3271	<ul> <li>C. Grosso-Pilcher, G. Landsberg, M. Pa</li> <li>F. Abe et al.</li> <li>M. Acciarri et al.</li> <li>T. Arima et al.</li> <li>G. Barenboim et al.</li> <li>A. Deandrea</li> <li>M. Derrick et al.</li> <li>Y. Grossman, Z. Ligeti, E. Nardi</li> <li>S. Jadach, B.F.L. Ward, Z. Was</li> <li>A. Stahl, H. Voss</li> <li>S. Abachi et al.</li> </ul>	(CDF Collab.) (CDF Collab.) (CDF Collab.) (CDF Collab.) (CDF Collab.) (L3 Collab.) (VENUS Collab.) (VALE, IFIC) (MARS) (ZEUS Collab.) (REHO, CIT) (CERN, INPK+) (BONN) (D0 Collab.)
GROSS-PILCH. ABE ABE ABE ACCIARRI ARIMA BARENBOIM DEANDREA DERRICK GROSSMAN JADACH STAHL ABACHI ABACHI	98 97F 97G 97S 97W 97X 97Q 97 97 97 97 97 97 97 97 96C 96D	hep-ex/9810015 PRL 78 2906 PR D55 R5263 PRL 79 2192 PRL 79 3819 PRL 79 4327 PL B412 201 PR D55 19 PR D55 4213 PL B409 277 ZPHY C73 613 PR D55 2768 PL B408 281 ZPHY C74 73 PRL 76 3271 PL B385 471	<ul> <li>C. Grosso-Pilcher, G. Landsberg, M. Pa</li> <li>F. Abe et al.</li> <li>M. Acciarri et al.</li> <li>T. Arima et al.</li> <li>G. Barenboim et al.</li> <li>A. Deandrea</li> <li>M. Derrick et al.</li> <li>Y. Grossman, Z. Ligeti, E. Nardi</li> <li>S. Jadach, B.F.L. Ward, Z. Was</li> <li>A. Stahl, H. Voss</li> <li>S. Abachi et al.</li> <li>S. Abachi et al.</li> </ul>	(CDF Collab.) (CDF Collab.) (CDF Collab.) (CDF Collab.) (CDF Collab.) (L3 Collab.) (VENUS Collab.) (VALE, IFIC) (MARS) (ZEUS Collab.) (REHO, CIT) (CERN, INPK+) (BONN) (D0 Collab.)
GROSS-PILCH. ABE ABE ABE ACCIARRI ARIMA BARENBOIM DEANDREA DERRICK GROSSMAN JADACH STAHL ABACHI ABACHI ABREU ADAM	98 97F 97G 97S 97W 97X 97 97 97 97 97 97 97 97 97 97 97 97 97	hep-ex/9810015 PRL 78 2906 PR D55 R5263 PRL 79 2192 PRL 79 3819 PRL 79 4327 PL B412 201 PR D55 19 PR D55 19 PR D55 4213 PL B409 277 ZPHY C73 613 PR D55 2768 PL B408 281 ZPHY C74 73 PRL 76 3271 PL B385 471 ZPHY C72 179 PL B380 471	<ul> <li>C. Grosso-Pilcher, G. Landsberg, M. Pa</li> <li>F. Abe et al.</li> <li>M. Acciarri et al.</li> <li>M. Acciarri et al.</li> <li>G. Barenboim et al.</li> <li>A. Deandrea</li> <li>M. Derrick et al.</li> <li>Y. Grossman, Z. Ligeti, E. Nardi</li> <li>S. Jadach, B.F.L. Ward, Z. Was</li> <li>A. Stahl, H. Voss</li> <li>S. Abachi et al.</li> <li>P. Aberu et al.</li> <li>W. Adam et al.</li> </ul>	(CDF Collab.) (CDF Collab.) (CDF Collab.) (CDF Collab.) (CDF Collab.) (L3 Collab.) (VENUS Collab.) (VALE, IFIC) (MARS) (ZEUS Collab.) (REHO, CIT) (CERN, INPK+) (BONN) (D0 Collab.) (D0 Collab.) (DELPHI Collab.)
GROSS-PILCH. ABE ABE ABE ABE ACCIARRI ARIMA BARENBOIM DEANDREA DERRICK GROSSMAN JADACH STAHL ABACHI ABACHI ABREU ADAM AID	98 97F 97G 97S 97W 97X 97Q 97 97 97 97 97 97 97 97 97 96C 96D 96T 96C	hep-ex/9810015 PRL 78 2906 PR D55 R5263 PRL 79 2192 PRL 79 3819 PRL 79 4327 PL B412 201 PR D55 19 PR D55 4213 PL B409 277 ZPHY C73 613 PR D55 2768 PL B408 281 ZPHY C74 73 PRL 76 3271 PL B385 471 ZPHY C72 179 PL B380 471 PL B369 173	<ul> <li>C. Grosso-Pilcher, G. Landsberg, M. Pa</li> <li>F. Abe et al.</li> <li>M. Acciarri et al.</li> <li>M. Acciarri et al.</li> <li>G. Barenboim et al.</li> <li>A. Deandrea</li> <li>M. Derrick et al.</li> <li>Y. Grossman, Z. Ligeti, E. Nardi</li> <li>S. Jadach, B.F.L. Ward, Z. Was</li> <li>A. Stahl, H. Voss</li> <li>S. Abachi et al.</li> <li>P. Abreu et al.</li> <li>W. Adam et al.</li> <li>S. Aid et al.</li> </ul>	(CDF Collab.) (CDF Collab.) (CDF Collab.) (CDF Collab.) (CDF Collab.) (L3 Collab.) (VENUS Collab.) (VALE, IFIC) (MARS) (ZEUS Collab.) (REHO, CIT) (CERN, INPK+) (BONN) (D0 Collab.) (D0 Collab.) (DELPHI Collab.) (DELPHI Collab.) (H1 Collab.)
GROSS-PILCH. ABE ABE ABE ABE ACCIARRI ARIMA BARENBOIM DEANDREA DERRICK GROSSMAN JADACH STAHL ABACHI ABACHI ABACHI ABREU ADAM AID ALLET	98 97F 97G 97S 97W 97X 97Q 97 97 97 97 97 97 97 97 96C 96D 96T 96C 96B	hep-ex/9810015 PRL 78 2906 PR D55 R5263 PRL 79 2192 PRL 79 3819 PRL 79 4327 PL B412 201 PR D55 19 PR D55 19 PR D55 4213 PL B409 277 ZPHY C73 613 PR D55 2768 PL B408 281 ZPHY C74 73 PRL 76 3271 PL B385 471 ZPHY C72 179 PL B380 471 PL B369 173 PL B383 139	C. Grosso-Pilcher, G. Landsberg, M. Pa F. Abe et al. F. Abe et al. M. Acciarri et al. M. Acciarri et al. G. Barenboim et al. A. Deandrea M. Derrick et al. Y. Grossman, Z. Ligeti, E. Nardi S. Jadach, B.F.L. Ward, Z. Was A. Stahl, H. Voss S. Abachi et al. S. Abachi et al. P. Abreu et al. W. Adam et al. S. Aid et al. M. Allet et al. (VILL, LE	(CDF Collab.) (CDF Collab.) (CDF Collab.) (CDF Collab.) (CDF Collab.) (L3 Collab.) (VENUS Collab.) (VALE, IFIC) (MARS) (ZEUS Collab.) (ZEUS Collab.) (REHO, CIT) (CERN, INPK+) (BONN) (D0 Collab.) (D0 Collab.) (DELPHI Collab.) (DELPHI Collab.) (H1 Collab.) (UV, LOUV, WISC)
GROSS-PILCH. ABE ABE ABE ABE ACCIARRI ARIMA BARENBOIM DEANDREA DERRICK GROSSMAN JADACH STAHL ABACHI ABACHI ABREU ADAM AID ALLET ABACHI	98 97F 97G 97S 97W 97X 97 97 97 97 97 97 97 97 97 96C 96D 96C 96B 96 95E	hep-ex/9810015 PRL 78 2906 PR D55 R5263 PRL 79 2192 PRL 79 3819 PRL 79 4327 PL B412 201 PR D55 19 PR D55 4213 PL B409 277 ZPHY C73 613 PR D55 2768 PL B408 281 ZPHY C74 73 PRL 76 3271 PL B385 471 ZPHY C72 179 PL B380 471 PL B369 173 PL B383 139 PL B358 405	C. Grosso-Pilcher, G. Landsberg, M. Pa F. Abe et al. F. Abe et al. F. Abe et al. F. Abe et al. F. Abe et al. M. Acciarri et al. T. Arima et al. G. Barenboim et al. A. Deandrea M. Derrick et al. Y. Grossman, Z. Ligeti, E. Nardi S. Jadach, B.F.L. Ward, Z. Was A. Stahl, H. Voss S. Abachi et al. P. Abreu et al. W. Adam et al. S. Aid et al. M. Allet et al. S. Abachi et al.	(CDF Collab.) (CDF Collab.) (CDF Collab.) (CDF Collab.) (CDF Collab.) (L3 Collab.) (VENUS Collab.) (VALE, IFIC) (MARS) (ZEUS Collab.) (ZEUS Collab.) (REHO, CIT) (CERN, INPK+) (BONN) (D0 Collab.) (D0 Collab.) (DELPHI Collab.) (DELPHI Collab.) (H1 Collab.) (UV, LOUV, WISC) (D0 Collab.)
GROSS-PILCH. ABE ABE ABE ABE ACCIARRI ARIMA BARENBOIM DEANDREA DERRICK GROSSMAN JADACH STAHL ABACHI ABACHI ABACHI ABACHI ADAM AID ALLET ABACHI ABE	98 97F 97G 97S 97W 97X 97 97 97 97 97 97 97 97 97 97 97 96C 96D 96C 96B 96 95E 95M	hep-ex/9810015 PRL 78 2906 PR D55 R5263 PRL 79 2192 PRL 79 3819 PRL 79 4327 PL B412 201 PR D55 19 PR D55 4213 PL B409 277 ZPHY C73 613 PR D55 2768 PL B408 281 ZPHY C74 73 PRL 76 3271 PL B385 471 ZPHY C72 179 PL B380 471 PL B369 173 PL B383 139 PL B358 405 PRL 74 2900	C. Grosso-Pilcher, G. Landsberg, M. Pa F. Abe et al. F. Abe et al. M. Acciarri et al. T. Arima et al. G. Barenboim et al. A. Deandrea M. Derrick et al. Y. Grossman, Z. Ligeti, E. Nardi S. Jadach, B.F.L. Ward, Z. Was A. Stahl, H. Voss S. Abachi et al. F. Abreu et al. M. Allet et al. S. Abachi et al. M. Allet et al. F. Abe et al. F. Abe et al.	(CDF Collab.) (CDF Collab.) (CDF Collab.) (CDF Collab.) (CDF Collab.) (L3 Collab.) (VENUS Collab.) (VALE, IFIC) (MARS) (ZEUS Collab.) (ZEUS Collab.) (REHO, CIT) (CERN, INPK+) (BONN) (D0 Collab.) (D0 Collab.) (DELPHI Collab.) (DELPHI Collab.) (H1 Collab.) (UV, LOUV, WISC) (D0 Collab.) (CDF Collab.)
GROSS-PILCH. ABE ABE ABE ABE ACCIARRI ARIMA BARENBOIM DEANDREA DERRICK GROSSMAN JADACH STAHL ABACHI ABACHI ABREU ADAM AID ALLET ABACHI ABE ABE	98 97F 97G 97S 97W 97X 977 97 97 97 97 97 97 97 97 97 97 96C 96D 96T 96C 96B 95E 95M 95N	hep-ex/9810015 PRL 78 2906 PR D55 R5263 PRL 79 2192 PRL 79 3819 PRL 79 4327 PL B412 201 PR D55 19 PR D55 4213 PL B409 277 ZPHY C73 613 PR D55 2768 PL B408 281 ZPHY C74 73 PRL 76 3271 PL B385 471 ZPHY C72 179 PL B380 471 PL B380 471 PL B383 139 PL B358 405 PRL 74 2900 PRL 74 3538	C. Grosso-Pilcher, G. Landsberg, M. Pa F. Abe et al. F. Abe et al. M. Acciarri et al. T. Arima et al. G. Barenboim et al. A. Deandrea M. Derrick et al. Y. Grossman, Z. Ligeti, E. Nardi S. Jadach, B.F.L. Ward, Z. Was A. Stahl, H. Voss S. Abachi et al. S. Abachi et al. S. Abachi et al. M. Allet et al. S. Abachi et al. F. Abe et al. F. Abe et al.	(CDF Collab.) (CDF Collab.) (CDF Collab.) (CDF Collab.) (CDF Collab.) (L3 Collab.) (VENUS Collab.) (VALE, IFIC) (VALE, IFIC) (ZEUS Collab.) (ZEUS Collab.) (REHO, CIT) (CERN, INPK+) (BONN) (D0 Collab.) (D0 Collab.) (DELPHI Collab.) (DELPHI Collab.) (DELPHI Collab.) (DU Collab.) (DO Collab.) (DO Collab.) (CDF Collab.) (CDF Collab.)
GROSS-PILCH. ABE ABE ABE ABE ACCIARRI ARIMA BARENBOIM DEANDREA DERRICK GROSSMAN JADACH STAHL ABACHI ABACHI ABACHI ABACHI ADAM AID ALLET ABACHI ABE	98 97F 97G 97S 97W 97X 97 97 97 97 97 97 97 97 97 97 97 96C 96D 96C 96B 96 95E 95M	hep-ex/9810015 PRL 78 2906 PR D55 R5263 PRL 79 2192 PRL 79 3819 PRL 79 4327 PL B412 201 PR D55 19 PR D55 4213 PL B409 277 ZPHY C73 613 PR D55 2768 PL B408 281 ZPHY C74 73 PRL 76 3271 PL B385 471 ZPHY C72 179 PL B380 471 PL B369 173 PL B383 139 PL B358 405 PRL 74 2900	C. Grosso-Pilcher, G. Landsberg, M. Pa F. Abe et al. F. Abe et al. M. Acciarri et al. T. Arima et al. G. Barenboim et al. A. Deandrea M. Derrick et al. Y. Grossman, Z. Ligeti, E. Nardi S. Jadach, B.F.L. Ward, Z. Was A. Stahl, H. Voss S. Abachi et al. F. Abreu et al. M. Allet et al. S. Abachi et al. M. Allet et al. F. Abe et al. F. Abe et al.	(CDF Collab.) (CDF Collab.) (CDF Collab.) (CDF Collab.) (CDF Collab.) (L3 Collab.) (VENUS Collab.) (VALE, IFIC) (MARS) (ZEUS Collab.) (ZEUS Collab.) (REHO, CIT) (CERN, INPK+) (BONN) (D0 Collab.) (D0 Collab.) (DELPHI Collab.) (DELPHI Collab.) (H1 Collab.) (UV, LOUV, WISC) (D0 Collab.) (CDF Collab.)
GROSS-PILCH. ABE ABE ABE ABE ACCIARRI ARIMA BARENBOIM DEANDREA DERRICK GROSSMAN JADACH STAHL ABACHI ABACHI ABREU ADAM AID ALLET ABACHI ABE ABE	98 97F 97G 97S 97W 97X 977 97 97 97 97 97 97 97 97 97 97 96C 96D 96T 96C 96B 95E 95M 95N	hep-ex/9810015 PRL 78 2906 PR D55 R5263 PRL 79 2192 PRL 79 3819 PRL 79 4327 PL B412 201 PR D55 19 PR D55 4213 PL B409 277 ZPHY C73 613 PR D55 2768 PL B408 281 ZPHY C74 73 PRL 76 3271 PL B385 471 ZPHY C72 179 PL B380 471 PL B380 471 PL B383 139 PL B358 405 PRL 74 2900 PRL 74 3538	C. Grosso-Pilcher, G. Landsberg, M. Pa F. Abe et al. F. Abe et al. M. Acciarri et al. T. Arima et al. G. Barenboim et al. A. Deandrea M. Derrick et al. Y. Grossman, Z. Ligeti, E. Nardi S. Jadach, B.F.L. Ward, Z. Was A. Stahl, H. Voss S. Abachi et al. S. Abachi et al. S. Abachi et al. S. Abachi et al. S. Aid et al. M. Allet et al. F. Abe et al. F. Abe et al. R. Balest et al.	(CDF Collab.) (CDF Collab.) (CDF Collab.) (CDF Collab.) (CDF Collab.) (L3 Collab.) (VENUS Collab.) (VALE, IFIC) (VALE, IFIC) (ZEUS Collab.) (ZEUS Collab.) (REHO, CIT) (CERN, INPK+) (BONN) (D0 Collab.) (D0 Collab.) (DELPHI Collab.) (DELPHI Collab.) (DELPHI Collab.) (DU Collab.) (DO Collab.) (DO Collab.) (CDF Collab.) (CDF Collab.)
GROSS-PILCH. ABE ABE ABE ABE ABE ACCIARRI ARIMA BARENBOIM DEANDREA DERRICK GROSSMAN JADACH STAHL ABACHI ABACHI ABACHI ABACHI ABREU ADAM AID ALLET ABACHI ABE BALEST KUZNETSOV	98 97F 97G 97S 97W 97X 97Q 97 97 97 97 97 97 97 97 97 97 96C 96D 96C 96B 96C 96B 95M 95N 95 95	hep-ex/9810015 PRL 78 2906 PR D55 R5263 PRL 79 2192 PRL 79 3819 PRL 79 4327 PL B412 201 PR D55 19 PR D55 4213 PL B409 277 ZPHY C73 613 PR D55 2768 PL B408 281 ZPHY C74 73 PRL 76 3271 PL B385 471 ZPHY C72 179 PL B380 471 PL B389 173 PL B383 139 PL B358 405 PRL 74 2900 PRL 74 3538 PR D51 2053 PRL 75 794	C. Grosso-Pilcher, G. Landsberg, M. Pa F. Abe et al. F. Abe et al. M. Acciarri et al. T. Arima et al. G. Barenboim et al. A. Deandrea M. Derrick et al. Y. Grossman, Z. Ligeti, E. Nardi S. Jadach, B.F.L. Ward, Z. Was A. Stahl, H. Voss S. Abachi et al. S. Abachi et al. P. Abreu et al. S. Aid et al. M. Allet et al. F. Abe et al. F. Abe et al. F. Abe et al. R. Balest et al. I.A. Kuznetsov et al. (PN	(CDF Collab.) (CDF Collab.) (CDF Collab.) (CDF Collab.) (CDF Collab.) (L3 Collab.) (VENUS Collab.) (VALE, IFIC) (VALE, IFIC) (MARS) (ZEUS Collab.) (ZEUS Collab.) (REHO, CIT) (CERN, INPK+) (BONN) (D0 Collab.) (D0 Collab.) (D0 Collab.) (DELPHI Collab.) (DELPHI Collab.) (DELPHI Collab.) (D1 Collab.) (D0 Collab.) (CDF Collab.) (CDF Collab.) (CDF Collab.) (CDF Collab.)
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GROSS-PILCH. ABE ABE ABE ABE ABE ACCIARRI ARIMA BARENBOIM DEANDREA DERRICK GROSSMAN JADACH STAHL ABACHI ABACHI ABACHI ABACHI ABREU ADAM AID ALLET ABACHI ABE BALEST KUZNETSOV KUZNETSOV MIZUKOSHI ABREU BHATTACH Also	98 97F 97G 97S 97W 97X 97Q 97 97 97 97 97 97 97 97 97 96C 96D 96C 96B 96C 96B 95M 95N 95 95 95B 95 95B 95 940 94 94 83	hep-ex/9810015 PRL 78 2906 PR D55 R5263 PRL 79 2192 PRL 79 3819 PRL 79 4327 PL B412 201 PR D55 19 PR D55 4213 PL B409 277 ZPHY C73 613 PR D55 2768 PL B408 281 ZPHY C74 73 PRL 76 3271 PL B385 471 ZPHY C72 179 PL B380 471 PL B369 173 PL B383 139 PL B358 405 PRL 74 2900 PRL 74 3538 PR D51 2053 PRL 75 794 PAN 58 2113 Translated from YAF 58 NP B443 20 ZPHY C64 183 PL B336 100 PL B338 522 (erratum)	C. Grosso-Pilcher, G. Landsberg, M. Pa F. Abe et al. F. Abe et al. M. Acciarri et al. T. Arima et al. G. Barenboim et al. A. Deandrea M. Derrick et al. Y. Grossman, Z. Ligeti, E. Nardi S. Jadach, B.F.L. Ward, Z. Was A. Stahl, H. Voss S. Abachi et al. S. Abachi et al. S. Abachi et al. S. Abachi et al. S. Aid et al. M. Allet et al. S. Abac et al. F. Abe et al. R. Balest et al. R. Balest et al. R. Balest et al. I.A. Kuznetsov et al. M. Alter et al. R. Balest et al. G. Bhattacharyya, J. Ellis, K. Sridhar G. Bhattacharyya, J. Ellis, K. Sridhar	(CDF Collab.) (CDF Collab.) (CDF Collab.) (CDF Collab.) (CDF Collab.) (L3 Collab.) (VENUS Collab.) (VALE, IFIC) (MARS) (ZEUS Collab.) (ZEUS Collab.) (REHO, CIT) (CERN, INPK+) (BONN) (D0 Collab.) (D0 Collab.) (D0 Collab.) (D0 Collab.) (DELPHI Collab.) (D1 Collab.) (D0 Collab.) (D0 Collab.) (D0 Collab.) (D0 Collab.) (CDF Collab.) (CERN) (CERN)

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DAVIDSON	94	ZPHY C61 613	S. Davidson, D. Bailey, B.A. Campbell (CFPA+)
KUZNETSOV	94	PL B329 295	A.V. Kuznetsov, N.V. Mikheev (YARO)
KUZNETSOV	94B	JETPL 60 315 Translated from ZETFP	I.A. Kuznetsov <i>et al.</i> (PNPI, KIAE, HARV+)
LEURER	94	PR D50 536	M. Leurer (REHO)
LEURER	94B	PR D49 333	M. Leurer (REHO)
Also	93	PRL 71 1324	M. Leurer (REHO)
MAHANTA	94	PL B337 128	U. Mahanta (MEHTA)
SEVERIJNS	94	PRL 73 611 (erratum)	N. Severijns <i>et al.</i> (LOUV, WISC, LEUV+)
VILAIN	94B	PL B332 465	P. Vilain <i>et al.</i> (CHARM II Collab.)
ABE	93C	PL B302 119	K. Abe <i>et al.</i> (VENUS Collab.)
ABE	93D	PL B304 373	T. Abe <i>et al.</i> (TOPAZ Collab.)
ABE	93G	PRL 71 2542	F. Abe <i>et al.</i> (CDF Collab.)
ABREU	93J	PL B316 620	P. Abreu <i>et al.</i> (DELPHI Collab.)
ACTON	93E	PL B311 391	P.D. Acton <i>et al.</i> (OPAL Collab.)
ADRIANI ALITTI	93ivi 93	PRPL 236 1 NP B400 3	O. Adriani <i>et al.</i> (L3 Collab.) J. Alitti <i>et al.</i> (UA2 Collab.)
BHATTACH	93 93	PR D47 R3693	G. Bhattacharyya <i>et al.</i> (CALC, JADA, ICTP+)
BUSKULIC	93F	PL B308 425	D. Buskulic <i>et al.</i> (ALEPH Collab.)
DERRICK	93	PL B306 173	M. Derrick <i>et al.</i> (ZEUS Collab.)
RIZZO	93	PR D48 4470	T.G. Rizzo (ANL)
SEVERIJNS	93	PRL 70 4047	N. Severijns <i>et al.</i> (LOUV, WISC, $LEUV+$ )
Also	94	PRL 73 611 (erratum)	N. Severijns <i>et al.</i> (LOUV, WISC, LEUV+)
STERNER	93	PL B303 385	K.L. Sterner <i>et al.</i> (AMY Collab.)
ABREU	92D	ZPHY C53 555	P. Abreu <i>et al.</i> (DELPHI Collab.)
ADRIANI	92F	PL B292 472	O. Adriani <i>et al.</i> (L3 Collab.)
DECAMP	92	PRPL 216 253	D. Decamp <i>et al.</i> (ALEPH Collab.)
	92	PRL 69 877	J. Imazato <i>et al.</i> (KEK, INUS, TOKY+)
MISHRA POLAK	92 92B	PRL 68 3499 PR D46 3871	S.R. Mishra <i>et al.</i> (COLU, CHIC, FNAL+) J. Polak, M. Zralek (SILES)
ACTON	92Б 91	PL B268 122	J. Polak, M. Zralek (SILES) D.P. Acton <i>et al.</i> (OPAL Collab.)
ACTON	91B	PL B273 338	D.P. Acton <i>et al.</i> (OFAL Collab.)
ADEVA	91D	PL B262 155	B. Adeva <i>et al.</i> (L3 Collab.)
AQUINO	91	PL B261 280	M. Aquino, A. Fernandez, A. Garcia (CINV, PUEB)
COLANGELO	91	PL B253 154	P. Colangelo, G. Nardulli (BARI)
CUYPERS	91	PL B259 173	F. Cuypers, A.F. Falk, P.H. Frampton (DURH, HÀRV+)
FARAGGI	91	MPL A6 61	A.E. Faraggi, D.V. Nanopoulos (TAMU)
POLAK	91	NP B363 385	J. Polak, M. Zralek (SILES)
RIZZO	91	PR D44 202	T.G. Rizzo (WISC, ISU)
WALKER	91	APJ 376 51	T.P. Walker <i>et al.</i> (HSCA, OSU, CHIC+)
ABE	90F	PL B246 297	K. Abe <i>et al.</i> (VENUS Collab.)
ABE AKRAWY	90H 90J	PR D41 1722 PL B246 285	F. Abe <i>et al.</i> (CDF Collab.) M.Z. Akrawy <i>et al.</i> (OPAL Collab.)
ANTREASYAN		PL B251 204	D. Antreasyan <i>et al.</i> (Crystal Ball Collab.)
GONZALEZ-G.		PL B240 163	M.C. Gonzalez-Garcia, J.W.F. Valle (VALE)
GRIFOLS	90	NP B331 244	J.A. Grifols, E. Masso (BARC)
GRIFOLS	90D	PR D42 3293	J.A. Grifols, E. Masso, T.G. Rizzo (BARC, CERN+)
KIM	90	PL B240 243	G.N. Kim <i>et al.</i> (AMY Collab.)
LOPEZ	90	PL B241 392	J.L. Lopez, D.V. Nanopoulos (TAMU)
ALBAJAR	89	ZPHY C44 15	C. Albajar <i>et al.</i> (UA1 Collab.)
ALBRECHT	89	ZPHY C42 349	H. Albrecht <i>et al.</i> (ARGUS Collab.)
BARBIERI	89B	PR D39 1229	R. Barbieri, R.N. Mohapatra (PISA, UMD)
	89B	PR D40 1569	P. Langacker, S. Uma Sankar (PENN)
ODAKA ROBINETT	89 89	JPSJ 58 3037 PR D39 834	S. Odaka <i>et al.</i> (VENUS Collab.) R.W. Robinett (PSU)
ALBAJAR	88B	PL B209 127	C. Albajar <i>et al.</i> (UA1 Collab.)
BAGGER	88	PR D37 1188	J. Bagger, C. Schmidt, S. King (HARV, BOST)
BALKE	88	PR D37 587	B. Balke <i>et al.</i> (LBL, UCB, COLO, NWES+)
BERGSTROM	88	PL B212 386	L. Bergstrom (STOH)
CUYPERS	88	PRL 60 1237	F. Cuypers, P.H. Frampton (UNCCH)
DONCHESKI	88	PL B206 137	M.A. Doncheski, H. Grotch, R. Robinett (PSU)
DONCHESKI	88B	PR D38 412	M.A. Doncheski, H. Grotch, R.W. Robinett (PSU)
ANSARI	87D	PL B195 613	R. Ansari <i>et al.</i> (UA2 Collab.)
BARTEL	87B	ZPHY C36 15	W. Bartel <i>et al.</i> (JADE Collab.)
BEHREND	86B	PL B178 452 PL 166B 463	H.J. Behrend <i>et al.</i> (CELLO Collab.)
DERRICK Also	86 86B	PL 166B 463 PR D34 3286	M. Derrick <i>et al.</i> (HRS Collab.) M. Derrick <i>et al.</i> (HRS Collab.)
JODIDIO	86	PR D34 1967	A. Jodidio <i>et al.</i> (LBL, NWES, TRIU)
Also	88	PR D37 237 erratum	A. Jodidio <i>et al.</i> (LBL, NWES, TRIU)
MOHAPATRA	86	PR D34 909	R.N. Mohapatra (UMD)
ADEVA	85	PL 152B 439	B. Adeva <i>et al.</i> (Mark-J Collab.)

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BERGER	85B	ZPHY C27 341	C. Berger <i>et al.</i>	(PLUTO Collab.)
STOKER	85	PRL 54 1887	D.P. Stoker <i>et al.</i>	(LBÌ, NWES, TRIÚ)
ADEVA	84	PRL 53 134	B. Adeva <i>et al.</i>	(Mark-J Collab.)
BEHREND	84C	PL 140B 130	H.J. Behrend et al.	(CELLO Collab.)
BERGSMA	83	PL 122B 465	F. Bergsma <i>et al.</i>	(CHARM Collab.)
CARR	83	PRL 51 627	J. Carr <i>et al.</i>	(LBL, NWES, TRIU)
BEALL	82	PRL 48 848	G. Beall, M. Bander, A. Soni	(UCI, UCLA)
SHANKER	82	NP B204 375	O. Shanker	(TRIU)
STEIGMAN	79	PRL 43 239	G. Steigman, K.A. Olive, D.N. Sch	hramm (BART+)

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