

# Free Quark Searches

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## Quark Production Cross Section — Accelerator Searches

<u>X-SECT</u> (cm <sup>2</sup> )	<u>CHG</u> (e/3)	<u>MASS</u> (GeV)	<u>ENERGY</u> (GeV)	<u>BEAM</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>
<1.3E-36	±2	45-84	130-172	e <sup>+</sup> e <sup>-</sup>	0	ABREU 97D	DLPH
<2.E-35	+2	250	1800	p $\bar{p}$	0	<sup>1</sup> ABE 92J	CDF
<1.E-35	+4	250	1800	p $\bar{p}$	0	<sup>1</sup> ABE 92J	CDF
<3.8E-28			14.5A	<sup>28</sup> Si-Pb	0	<sup>2</sup> HE 91	PLAS
<3.2E-28			14.5A	<sup>28</sup> Si-Cu	0	<sup>2</sup> HE 91	PLAS
<1.E-40	±1,2	<10		p, $\nu$ , $\bar{\nu}$	0	BERGSMA 84B	CHRM
<1.E-36	±1,2	<9	200	$\mu$	0	AUBERT 83C	SPEC
<2.E-10	±2,4	1-3	200	p	0	<sup>3</sup> BUSSIERE 80	CNTR
<5.E-38	+1,2	>5	300	p	0	<sup>4,5</sup> STEVENSON 79	CNTR
<1.E-33	±1	<20	52	pp	0	BASILE 78	SPEC
<9.E-39	±1,2	<6	400	p	0	<sup>4</sup> ANTREASYAN 77	SPEC
<8.E-35	+1,2	<20	52	pp	0	<sup>6</sup> FABJAN 75	CNTR
<5.E-38	-1,2	4-9	200	p	0	NASH 74	CNTR
<1.E-32	+2,4	4-24	52	pp	0	ALPER 73	SPEC
<5.E-31	+1,2,4	<12	300	p	0	LEIPUNER 73	CNTR
<6.E-34	±1,2	<13	52	pp	0	BOTT 72	CNTR
<1.E-36	-4	4	70	p	0	ANTIPOV 71	CNTR
<1.E-35	±1,2	2	28	p	0	<sup>7</sup> ALLABY 69B	CNTR
<4.E-37	-2	<5	70	p	0	<sup>3</sup> ANTIPOV 69	CNTR
<3.E-37	-1,2	2-5	70	p	0	<sup>7</sup> ANTIPOV 69B	CNTR
<1.E-35	+1,2	<7	30	p	0	DORFAN 65	CNTR
<2.E-35	-2	<2.5-5	30	p	0	<sup>8</sup> FRANZINI 65B	CNTR
<5.E-35	+1,2	<2.2	21	p	0	BINGHAM 64	HLBC
<1.E-32	+1,2	<4.0	28	p	0	BLUM 64	HBC
<1.E-35	+1,2	<2.5	31	p	0	<sup>8</sup> HAGOPIAN 64	HBC
<1.E-34	+1	<2	28	p	0	LEIPUNER 64	CNTR
<1.E-33	+1,2	<2.4	24	p	0	MORRISON 64	HBC

<sup>1</sup> ABE 92J flux limits decrease as the mass increases from 50 to 500 GeV.

<sup>2</sup> HE 91 limits are for charges of the form  $N \pm 1/3$  from 23/3 to 38/3.

<sup>3</sup> Hadronic or leptonic quarks.

<sup>4</sup> Cross section cm<sup>2</sup>/GeV<sup>2</sup>.

<sup>5</sup>  $3 \times 10^{-5} < \text{lifetime} < 1 \times 10^{-3}$  s.

<sup>6</sup> Includes BOTT 72 results.

<sup>7</sup> Assumes isotropic cm production.

<sup>8</sup> Cross section inferred from flux.

## Quark Differential Production Cross Section — Accelerator Searches

<u>X-SECT</u> (cm <sup>2</sup> sr <sup>-1</sup> GeV <sup>-1</sup> )	<u>CHG</u> e/3	<u>MASS</u> (GeV)	<u>ENERGY</u> (GeV)	<u>BEAM</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>
<4.E-36	-2,4	1.5-6	70	p	0	BALDIN 76	CNTR
<2.E-33	±4	5-20	52	pp	0	ALBROW 75	SPEC
<5.E-34	<7	7-15	44	pp	0	JOVANOVA... 75	CNTR
<5.E-35			20	$\gamma$	0	<sup>9</sup> GALIK 74	CNTR

<9.E-35	-1,2		200	<i>p</i>	0	NASH	74	CNTR
<4.E-36	-4	2.3-2.7	70	<i>p</i>	0	ANTIPOV	71	CNTR
<3.E-35	±1,2	<2.7	27	<i>p</i>	0	ALLABY	69B	CNTR
<7.E-38	-1,2	<2.5	70	<i>p</i>	0	ANTIPOV	69B	CNTR

<sup>9</sup> Cross section in cm<sup>2</sup>/sr/equivalent quanta.

## Quark Flux — Accelerator Searches

The definition of FLUX depends on the experiment

- (a) is the ratio of measured free quarks to predicted free quarks if there is no “confinement.”
- (b) is the probability of fractional charge on nuclear fragments. Energy is in GeV/nucleon.
- (c) is the 90%CL upper limit on fractionally-charged particles produced per interaction.
- (d) is quarks per collision.
- (e) is inclusive quark-production cross-section ratio to  $\sigma(e^+e^- \rightarrow \mu^+\mu^-)$ .
- (f) is quark flux per charged particle.
- (g) is the flux per  $\nu$ -event.
- (h) is quark yield per  $\pi^-$  yield.
- (i) is 2-body exclusive quark-production cross-section ratio to  $\sigma(e^+e^- \rightarrow \mu^+\mu^-)$ .

<u>FLUX</u>		<u>CHG</u> ( <i>e</i> /3)	<u>MASS</u> (GeV)	<u>ENRGY</u> (GeV)	<u>BEAM</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>
<1.6E-3	b	see note		200	<sup>32</sup> S-Pb	0	10 HUENTRUP	96 PLAS
<6.2E-4	b	see note		10.6	<sup>32</sup> S-Pb	0	10 HUENTRUP	96 PLAS
<0.94E-4	e	±2	2-30	88-94	<i>e</i> <sup>+</sup> <i>e</i> <sup>-</sup>	0	AKERS	95R OPAL
<1.7E-4	e	±2	30-40	88-94	<i>e</i> <sup>+</sup> <i>e</i> <sup>-</sup>	0	AKERS	95R OPAL
<3.6E-4	e	±4	5-30	88-94	<i>e</i> <sup>+</sup> <i>e</i> <sup>-</sup>	0	AKERS	95R OPAL
<1.9E-4	e	±4	30-45	88-94	<i>e</i> <sup>+</sup> <i>e</i> <sup>-</sup>	0	AKERS	95R OPAL
<2.E-3	e	+1	5-40	88-94	<i>e</i> <sup>+</sup> <i>e</i> <sup>-</sup>	0	11 BUSKULIC	93C ALEP
<6.E-4	e	+2	5-30	88-94	<i>e</i> <sup>+</sup> <i>e</i> <sup>-</sup>	0	11 BUSKULIC	93C ALEP
<1.2E-3	e	+4	15-40	88-94	<i>e</i> <sup>+</sup> <i>e</i> <sup>-</sup>	0	11 BUSKULIC	93C ALEP
<3.6E-4	i	+4	5.0-10.2	88-94	<i>e</i> <sup>+</sup> <i>e</i> <sup>-</sup>	0	BUSKULIC	93C ALEP
<3.6E-4	i	+4	16.5-26.0	88-94	<i>e</i> <sup>+</sup> <i>e</i> <sup>-</sup>	0	BUSKULIC	93C ALEP
<6.9E-4	i	+4	26.0-33.3	88-94	<i>e</i> <sup>+</sup> <i>e</i> <sup>-</sup>	0	BUSKULIC	93C ALEP
<9.1E-4	i	+4	33.3-38.6	88-94	<i>e</i> <sup>+</sup> <i>e</i> <sup>-</sup>	0	BUSKULIC	93C ALEP
<1.1E-3	i	+4	38.6-44.9	88-94	<i>e</i> <sup>+</sup> <i>e</i> <sup>-</sup>	0	BUSKULIC	93C ALEP
<1.6E-4	b	see note	see note			0	12 CECCHINI	93 PLAS
	b	4,5,7,8		2.1A	<sup>16</sup> O	0,2,0,6	13 GHOSH	92 EMUL
<6.4E-5	g	1			$\nu, \bar{\nu}$	1	14 BASILE	91 CNTR
<3.7E-5	g	2			$\nu, \bar{\nu}$	0	14 BASILE	91 CNTR
<3.9E-5	g	1			$\nu, \bar{\nu}$	1	15 BASILE	91 CNTR
<2.8E-5	g	2			$\nu, \bar{\nu}$	0	15 BASILE	91 CNTR
<1.9E-4	c			14.5A	<sup>28</sup> Si-Pb	0	16 HE	91 PLAS
<3.9E-4	c			14.5A	<sup>28</sup> Si-Cu	0	16 HE	91 PLAS
<1.E-9	c	±1,2,4		14.5A	<sup>16</sup> O-Ar	0	MATIS	91 MDRP
<5.1E-10	c	±1,2,4		14.5A	<sup>16</sup> O-Hg	0	MATIS	91 MDRP
<8.1E-9	c	±1,2,4		14.5A	Si-Hg	0	MATIS	91 MDRP

<1.7E-6	c	±1,2,4		60A	<sup>16</sup> O-Hg	0		MATIS	91	MDRP
<3.5E-7	c	±1,2,4		200A	<sup>16</sup> O-Hg	0		MATIS	91	MDRP
<1.3E-6	c	±1,2,4		200A	S-Hg	0		MATIS	91	MDRP
<5E-2	e	2	19-27	52-60	e <sup>+</sup> e <sup>-</sup>	0		ADACHI	90C	TOPZ
<5E-2	e	4	<24	52-60	e <sup>+</sup> e <sup>-</sup>	0		ADACHI	90C	TOPZ
<1.E-4	e	+2	<3.5	10	e <sup>+</sup> e <sup>-</sup>	0		BOWCOCK	89B	CLEO
<1.E-6	d	±1,2		60	<sup>16</sup> O-Hg	0		CALLOWAY	89	MDRP
<3.5E-7	d	±1,2		200	<sup>16</sup> O-Hg	0		CALLOWAY	89	MDRP
<1.3E-6	d	±1,2		200	S-Hg	0		CALLOWAY	89	MDRP
<1.2E-10	d	±1	1	800	p-Hg	0		MATIS	89	MDRP
<1.1E-10	d	±2	1	800	p-Hg	0		MATIS	89	MDRP
<1.2E-10	d	±1	1	800	p-N <sub>2</sub>	0		MATIS	89	MDRP
<7.7E-11	d	±2	1	800	p-N <sub>2</sub>	0		MATIS	89	MDRP
<6.E-9	h	-5	0.9-2.3	12	p	0		NAKAMURA	89	SPEC
<5.E-5	g	1,2	<0.5		ν,ν̄d	0		ALLASIA	88	BEBC
<3.E-4	b	See note		14.5	<sup>16</sup> O-Pb	0	17	HOFFMANN	88	PLAS
<2.E-4	b	See note		200	<sup>16</sup> O-Pb	0	18	HOFFMANN	88	PLAS
<8E-5	b	19,20,22,23		200A				GERBIER	87	PLAS
<2.E-4	a	±1,2	<300	320	p̄p	0		LYONS	87	MLEV
<1.E-9	c	±1,2,4,5		14.5	<sup>16</sup> O-Hg	0		SHAW	87	MDRP
<3.E-3	d	-1,2,3,4,6	<5	2	Si-Si	0	19	ABACHI	86C	CNTR
<1.E-4	e	±1,2,4	<4	10	e <sup>+</sup> e <sup>-</sup>	0		ALBRECHT	85G	ARG
<6.E-5	b	±1,2	1	540	p̄p̄	0		BANNER	85	UA2
<5.E-3	e	-4	1-8	29	e <sup>+</sup> e <sup>-</sup>	0		AIHARA	84	TPC
<1.E-2	e	±1,2	1-13	29	e <sup>+</sup> e <sup>-</sup>	0		AIHARA	84B	TPC
<2.E-4	b	±1		72	<sup>40</sup> Ar	0	20	BARWICK	84	CNTR
<1.E-4	e	±2	<0.4	1.4	e <sup>+</sup> e <sup>-</sup>	0		BONDAR	84	OLYA
<5.E-1	e	±1,2	<13	29	e <sup>+</sup> e <sup>-</sup>	0		GURYN	84	CNTR
<3.E-3	b	±1,2	<2	540	p̄p̄	0		BANNER	83	CNTR
<1.E-4	b	±1,2		106	<sup>56</sup> Fe	0		LINDGREN	83	CNTR
<3.E-3	b	>  ±0.1		74	<sup>40</sup> Ar	0	20	PRICE	83	PLAS
<1.E-2	e	±1,2	<14	29	e <sup>+</sup> e <sup>-</sup>	0		MARINI	82B	CNTR
<8.E-2	e	±1,2	<12	29	e <sup>+</sup> e <sup>-</sup>	0		ROSS	82	CNTR
<3.E-4	e	±2	1.8-2	7	e <sup>+</sup> e <sup>-</sup>	0		WEISS	81	MRK2
<5.E-2	e	+1,2,4,5	2-12	27	e <sup>+</sup> e <sup>-</sup>	0		BARTEL	80	JADE
<2.E-5	g	1,2			ν	0	14,15	BASILE	80	CNTR
<3.E-10	f	±2,4	1-3	200	p	0	21	BOZZOLI	79	CNTR
<6.E-11	f	±1	<21	52	pp	0		BASILE	78	SPEC
<5.E-3	g				νμ	0		BASILE	78B	CNTR
<2.E-9	f	±1	<26	62	pp	0		BASILE	77	SPEC
<7.E-10	f	+1,2	<20	52	p	0	22	FABJAN	75	CNTR
		+1,2	>4.5		γ	0	14,15	GALIK	74	CNTR
		+1,2	>1.5	12	e <sup>-</sup>	0	14,15	BELLAMY	68	CNTR
		+1,2	>0.9		γ	0	15	BATHOW	67	CNTR
		+1,2	>0.9	6	γ	0	15	FOSS	67	CNTR

<sup>10</sup> HUENTRUP 96 quote 95% CL limits for production of fragments with charge differing by as much as ±1/3 (in units of e) for charge 6 ≤ Z ≤ 10.

<sup>11</sup> BUSKULIC 93C limits for inclusive quark production are more conservative if the ALEPH hadronic fragmentation function is assumed.

- <sup>12</sup> CECCHINI 93 limit at 90%CL for  $23/3 \leq Z \leq 40/3$ , for 16A GeV O, 14.5A Si, and 200A S incident on Cu target. Other limits are  $2.3 \times 10^{-4}$  for  $17/3 \leq Z \leq 20/3$  and  $1.2 \times 10^{-4}$  for  $20/3 \leq Z \leq 23/3$ .
- <sup>13</sup> GHOSH 92 reports measurement of spallation fragment charge based on ionization in emulsion. Out of 650 measured tracks, 2 were consistent with charge  $5e/3$ , and 4 with  $7e/3$ .
- <sup>14</sup> Hadronic quark.
- <sup>15</sup> Leptonic quark.
- <sup>16</sup> HE 91 limits are for charges of the form  $N \pm 1/3$  from  $23/3$  to  $38/3$ , and correspond to cross-section limits of  $380 \mu\text{b}$  (Pb) and  $320 \mu\text{b}$  (Cu).
- <sup>17</sup> The limits apply to projectile fragment charges of 17, 19, 20, 22, 23 in units of  $e/3$ .
- <sup>18</sup> The limits apply to projectile fragment charges of 16, 17, 19, 20, 22, 23 in units of  $e/3$ .
- <sup>19</sup> Flux limits and mass range depend on charge.
- <sup>20</sup> Bound to nuclei.
- <sup>21</sup> Quark lifetimes  $> 1 \times 10^{-8}$  s.
- <sup>22</sup> One candidate  $m < 0.17$  GeV.

### Quark Flux — Cosmic Ray Searches

Shielding values followed with an asterisk indicate altitude in km. Shielding values not followed with an asterisk indicate sea level in  $\text{kg}/\text{cm}^2$ .

$FLUX$ ( $\text{cm}^{-2}\text{sr}^{-1}\text{s}^{-1}$ )	$CHG$ ( $e/3$ )	$MASS$ (GeV)	$SHIELDING$	$EVTS$	$DOCUMENT ID$	$TECN$
$< 9.2E-15$	$\pm 1$		3800	0	<sup>23</sup> AMBROSIO	00C MCRO
$< 2.1E-15$	$\pm 1$			0	MORI	91 KAM2
$< 2.3E-15$	$\pm 2$			0	MORI	91 KAM2
$< 2.E-10$	$\pm 1, 2$		0.3	0	WADA	88 CNTR
	$\pm 4$		0.3	12	<sup>24</sup> WADA	88 CNTR
	$\pm 4$		0.3	9	<sup>25</sup> WADA	86 CNTR
$< 1.E-12$	$\pm 2, 3/2$		-70.	0	<sup>26</sup> KAWAGOE	84B PLAS
$< 9.E-10$	$\pm 1, 2$		0.3	0	WADA	84B CNTR
$< 4.E-9$	$\pm 4$		0.3	7	WADA	84B CNTR
$< 2.E-12$	$\pm 1, 2, 3$		-0.3 *	0	MASHIMO	83 CNTR
$< 3.E-10$	$\pm 1, 2$		0.3	0	MARINI	82 CNTR
$< 2.E-11$	$\pm 1, 2$			0	MASHIMO	82 CNTR
$< 8.E-10$	$\pm 1, 2$		0.3	0	<sup>26</sup> NAPOLITANO	82 CNTR
				3	<sup>27</sup> YOCK	78 CNTR
$< 1.E-9$				0	<sup>28</sup> BRIATORE	76 ELEC
$< 2.E-11$	+1			0	<sup>29</sup> HAZEN	75 CC
$< 2.E-10$	+1, 2			0	KRISOR	75 CNTR
$< 1.E-7$	+1, 2			0	<sup>29,30</sup> CLARK	74B CC
$< 3.E-10$	+1	>20		0	KIFUNE	74 CNTR
$< 8.E-11$	+1			0	<sup>29</sup> ASHTON	73 CNTR
$< 2.E-8$	+1, 2			0	HICKS	73B CNTR
$< 5.E-10$	+4		2.8 *	0	BEAUCHAMP	72 CNTR
$< 1.E-10$	+1, 2			0	<sup>29</sup> BOHM	72B CNTR
$< 1.E-10$	+1, 2		2.8 *	0	COX	72 ELEC
$< 3.E-10$	+2			0	CROUCH	72 CNTR
$< 3.E-8$			7	0	<sup>28</sup> DARDO	72 CNTR
$< 4.E-9$	+1			0	<sup>29</sup> EVANS	72 CC
$< 2.E-9$		>10		0	<sup>28</sup> TONWAR	72 CNTR
$< 2.E-10$	+1		2.8 *	0	CHIN	71 CNTR

<3.E-10	+1,2			0	<sup>29</sup> CLARK	71B	CC
<1.E-10	+1,2			0	<sup>29</sup> HAZEN	71	CC
<5.E-10	+1,2		3.5 *	0	BOSIA	70	CNTR
	+1,2	<6.5		1	<sup>29</sup> CHU	70	HLBC
<2.E-9	+1			0	FAISSNER	70B	CNTR
<2.E-10	+1,2		0.8 *	0	KRIDER	70	CNTR
<5.E-11	+2			4	CAIRNS	69	CC
<8.E-10	+1,2	<10		0	FUKUSHIMA	69	CNTR
	+2			1	<sup>29,31</sup> MCCUSKER	69	CC
<1.E-10		>5	1.7,3.6	0	<sup>28</sup> BJORNBOE	68	CNTR
<1.E-8	±1,2,4		6.3,.2 *	0	<sup>26</sup> BRIATORE	68	CNTR
<3.E-8		>2		0	FRANZINI	68	CNTR
<9.E-11	±1,2			0	GARMIRE	68	CNTR
<4.E-10	±1			0	HANAYAMA	68	CNTR
<3.E-8		>15		0	KASHA	68	OSPK
<2.E-10	+2			0	KASHA	68B	CNTR
<2.E-10	+4			0	KASHA	68C	CNTR
<2.E-10	+2		6	0	BARTON	67	CNTR
<2.E-7	+4		0.008,0.5 *	0	BUHLER	67	CNTR
<5.E-10	1,2		0.008,0.5 *	0	BUHLER	67B	CNTR
<4.E-10	+1,2			0	GOMEZ	67	CNTR
<2.E-9	+2			0	KASHA	67	CNTR
<2.E-10	+2		220	0	BARTON	66	CNTR
<2.E-9	+1,2		0.5 *	0	BUHLER	66	CNTR
<3.E-9	+1,2			0	KASHA	66	CNTR
<2.E-9	+1,2			0	LAMB	66	CNTR
<2.E-8	+1,2	>7	2.8 *	0	DELISE	65	CNTR
<5.E-8	+2	>2.5	0.5 *	0	MASSAM	65	CNTR
<2.E-8	+1		2.5 *	0	BOWEN	64	CNTR
<2.E-7	+1		0.8	0	SUNYAR	64	CNTR

<sup>23</sup> AMBROSIO 00C limit is below  $11 \times 10^{-15}$  for  $0.25 < q/e < 0.5$ , and is changing rapidly near  $q/e=2/3$ , where it is  $2 \times 10^{-14}$ .

<sup>24</sup> Distribution in celestial sphere was described as anisotropic.

<sup>25</sup> With telescope axis at zenith angle  $40^\circ$  to the south.

<sup>26</sup> Leptonic quarks.

<sup>27</sup> Lifetime  $> 10^{-8}$  s; charge  $\pm 0.70, 0.68, 0.42$ ; and mass  $> 4.4, 4.8, \text{ and } 20$  GeV, respectively.

<sup>28</sup> Time delayed air shower search.

<sup>29</sup> Prompt air shower search.

<sup>30</sup> Also  $e/4$  and  $e/6$  charges.

<sup>31</sup> No events in subsequent experiments.

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### Quark Density — Matter Searches

<u>QUARKS/ NUCLEON</u>	<u>CHG (e/3)</u>	<u>MASS (GeV)</u>	<u>MATERIAL/METHOD</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>
<1.17E-22			silicone oil drops	0	<sup>32</sup> LEE 02
<4.71E-22			silicone oil drops	1	<sup>33</sup> HALYO 00
<4.7E-21	±1,2		silicone oil drops	0	MAR 96
<8.E-22	+2		Si/infrared photoionization	0	PERERA 93
<5.E-27	±1,2		sea water/levitation	0	HOMER 92
<4.E-20	±1,2		meteorites/mag. levitation	0	JONES 89

<1.E-19	±1,2		various/spectrometer	0	MILNER	87
<5.E-22	±1,2		W/levitation	0	SMITH	87
<3.E-20	+1,2		org liq/droplet tower	0	VANPOLEN	87
<6.E-20	-1,2		org liq/droplet tower	0	VANPOLEN	87
<3.E-21	±1		Hg drops-untreated	0	SAVAGE	86
<3.E-22	±1,2		levitated niobium	0	SMITH	86
<2.E-26	±1,2		<sup>4</sup> He/levitation	0	SMITH	86B
<2.E-20	>±1	0.2-250	niobium+tungs/ion	0	MILNER	85
<1.E-21	±1		levitated niobium	0	SMITH	85
	+1,2	<100	niobium/mass spec	0	KUTSCHERA	84
<5.E-22			levitated steel	0	MARINELLI	84
<9.E-20	± <13		water/oil drop	0	JOYCE	83
<2.E-21	>   ± 1/2		levitated steel	0	LIEBOWITZ	83
<1.E-19	±1,2		photo ion spec	0	VANDESTEEG	83
<2.E-20			mercury/oil drop	0	<sup>34</sup> HODGES	81
1.E-20	+1		levitated niobium	4	<sup>35</sup> LARUE	81
1.E-20	-1		levitated niobium	4	<sup>35</sup> LARUE	81
<1.E-21			levitated steel	0	MARINELLI	80B
<6.E-16			helium/mass spec	0	BOYD	79
1.E-20	+1		levitated niobium	2	<sup>35</sup> LARUE	79
<4.E-28			earth+/ion beam	0	OGOROD...	79
<5.E-15	+1		tungs./mass spec	0	BOYD	78
<5.E-16	+3	<1.7	hydrogen/mass spec	0	BOYD	78B
<1.E-21	±2,4		water/ion beam	0	LUND	78
<6.E-15	>1/2		levitated tungsten	0	PUTT	78
<1.E-22			metals/mass spec	0	SCHIFFER	78
<5.E-15			levitated tungsten ox	0	BLAND	77
<3.E-21			levitated iron	0	GALLINARO	77
2.E-21	-1		levitated niobium	1	<sup>35</sup> LARUE	77
4.E-21	+1		levitated niobium	2	<sup>35</sup> LARUE	77
<1.E-13	+3	<7.7	hydrogen/mass spec	0	MULLER	77
<5.E-27			water+/ion beam	0	OGOROD...	77
<1.E-21			lunar+/ion spec	0	STEVENS	76
<1.E-15	+1	<60	oxygen+/ion spec	0	ELBERT	70
<5.E-19			levitated graphite	0	MORPURGO	70
<5.E-23			water+/atom beam	0	COOK	69
<1.E-17	±1,2		levitated graphite	0	BRAGINSK	68
<1.E-17			water+/uv spec	0	RANK	68
<3.E-19	±1		levitated iron	0	STOVER	67
<1.E-10			sun/uv spec	0	<sup>36</sup> BENNETT	66
<1.E-17	+1,2		meteorites+/ion beam	0	CHUPKA	66
<1.E-16	±1		levitated graphite	0	GALLINARO	66
<1.E-22			argon/electrometer	0	HILLAS	59
	-2		levitated oil	0	MILLIKAN	10

<sup>32</sup> 95% CL limit for fractional charge particles with  $0.18e \leq |Q_{residual}| \leq 0.82e$  in total of 70.1 mg of silicone oil.

<sup>33</sup> 95% CL limit for particles with fractional charge  $|Q_{residual}| > 0.16e$  in total of 17.4 mg of silicone oil.

<sup>34</sup> Also set limits for  $Q = \pm e/6$ .

<sup>35</sup> Note that in PHILLIPS 88 these authors report a subtle magnetic effect which could account for the apparent fractional charges.

36 Limit inferred by JONES 77B.

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