

## **$V_{ud}$ , $V_{us}$ , THE CABIBBO ANGLE, AND CKM UNITARITY**

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The Cabibbo-Kobayashi-Maskawa (CKM) [1,2] three-generation quark mixing matrix written in terms of the Wolfenstein parameters  $(\lambda, A, \rho, \eta)$  [3] nicely illustrates the orthonormality constraint of unitarity and central role played by  $\lambda$ .

$$V_{\text{CKM}} = \begin{pmatrix} V_{ud} & V_{us} & V_{ub} \\ V_{cd} & V_{cs} & V_{cb} \\ V_{td} & V_{ts} & V_{tb} \end{pmatrix} = \begin{pmatrix} 1 - \lambda^2/2 & \lambda & A\lambda^3(\rho - i\eta) \\ -\lambda & 1 - \lambda^2/2 & A\lambda^2 \\ A\lambda^3(1 - \rho - i\eta) & -A\lambda^2 & 1 \end{pmatrix} + \mathcal{O}(\lambda^4). \quad (1)$$

That cornerstone is a carryover from the two-generation Cabibbo angle,  $\lambda = \sin(\theta_{\text{Cabibbo}}) = V_{us}$ . Its value is a critical ingredient in determinations of the other parameters and in tests of CKM unitarity.

Unfortunately, the precise value of  $\lambda$  has been somewhat controversial in the past, with kaon decays suggesting [4]  $\lambda \simeq 0.220$ , while hyperon decays [5] and indirect determinations via nuclear  $\beta$ -decays imply a somewhat larger  $\lambda \simeq 0.225 - 0.230$ . That discrepancy is often discussed in terms of a deviation from the unitarity requirement

$$|V_{ud}|^2 + |V_{us}|^2 + |V_{ub}|^2 = 1. \quad (2)$$

For many years, using a value of  $V_{us}$  derived from  $K \rightarrow \pi e \nu$  ( $K_{e3}$ ) decays, that sum was consistently 2–2.5 sigma below unity, a potential signal [6] for new physics effects. Below, we discuss the current status of  $V_{ud}$ ,  $V_{us}$ , and their associated unitarity test in Eq. (2). (Since  $|V_{ub}|^2 \simeq 1 \times 10^{-5}$  is negligibly small, it is ignored in this discussion.)

### **$V_{ud}$**

The value of  $V_{ud}$  has been obtained from superallowed nuclear, neutron, and pion decays. Currently, the most precise

determination of  $V_{ud}$  comes from superallowed nuclear beta-decays [6] ( $0^+ \rightarrow 0^+$  transitions). Measuring their half-lives,  $t$ , and  $Q$  values which give the decay rate factor,  $f$ , leads to a precise determination of  $V_{ud}$  via the master formula [7–9]

$$|V_{ud}|^2 = \frac{2984.48(5) \text{ sec}}{ft(1 + \text{RC})} \quad (3)$$

where RC denotes the entire effect of electroweak radiative corrections, nuclear structure, and isospin violating nuclear effects. RC is nucleus-dependent, ranging from about +3.0% to +3.6% for the best measured superallowed decays. The most recent analysis of Hardy and Towner [10, 11] gives a weighted average (with errors combined in quadrature) of

$$V_{ud} = 0.97425(22) \text{ (superallowed)} , \quad (4)$$

which, assuming unitarity, corresponds to  $\lambda = 0.2255(10)$ . The new average value of  $V_{ud}$  is shifted upward compared to our 2007 value of 0.97418(27) primarily because of improvements in the experimental  $ft$  values and nuclear isospin breaking corrections employed. We note, however, that the possibility of additional nuclear coulombic corrections has been raised recently [12].

Combined measurements of the neutron lifetime,  $\tau_n$ , and the ratio of axial-vector/vector couplings,  $g_A \equiv G_A/G_V$ , via neutron decay asymmetries can also be used to determine  $V_{ud}$ :

$$|V_{ud}|^2 = \frac{4908.7(1.9) \text{ sec}}{\tau_n(1 + 3g_A^2)} , \quad (5)$$

where the error stems from uncertainties in the electroweak radiative corrections [8] due to hadronic loop effects. Those effects have been recently updated and their error was reduced by about a factor of 2 [9], leading to a  $\pm 0.0002$  theoretical uncertainty in  $V_{ud}$  (common to all  $V_{ud}$  extractions). Using the world averages from this *Review*

$$\begin{aligned} \tau_n^{\text{ave}} &= 885.7(8) \text{ sec} \\ g_A^{\text{ave}} &= 1.2695(29) \end{aligned} \quad (6)$$

leads to

$$V_{ud} = 0.9746(4)_{\tau_n(18)} g_A(2)_{\text{RC}} \quad (7)$$

with the error dominated by  $g_A$  uncertainties (which have been expanded due to experimental inconsistencies). We note that a recent precise measurement [13] of  $\tau_n = 878.5(7)(3)$  sec is also inconsistent with the world average from this *Review* and would lead to a considerably larger  $V_{ud} = 0.9786(4)(18)(2)$ . Alternatively, accepting the recent shorter lifetime measurement as correct, and employing it along with the value of  $V_{ud}$  in Eq. (4), leads to  $g_A = 1.2763(7)$ , which is outside of the range of Eq. (6) but in good accord with the most recent direct measurements of  $g_A$  [14]. Future neutron studies are expected to resolve these inconsistencies and significantly reduce the uncertainties in  $g_A$  and  $\tau_n$ , potentially making them the best way to determine  $V_{ud}$ .

The recently completed PIBETA experiment at PSI measured the very small ( $\mathcal{O}(10^{-8})$ ) branching ratio for  $\pi^+ \rightarrow \pi^0 e^+ \nu_e$  with about  $\pm 1/2\%$  precision. Their result gives [15]

$$V_{ud} = 0.9749(26) \left[ \frac{BR(\pi^+ \rightarrow e^+ \nu_e(\gamma))}{1.2352 \times 10^{-4}} \right]^{\frac{1}{2}} \quad (8)$$

which is normalized using the very precisely determined theoretical prediction for  $BR(\pi^+ \rightarrow e^+ \nu_e(\gamma)) = 1.2352(5) \times 10^{-4}$  [7], rather than the experimental branching ratio from this *Review* of  $1.230(4) \times 10^{-4}$  which would lower the value to  $V_{ud} = 0.9728(30)$ . Theoretical uncertainties in that determination are very small; however, much higher statistics would be required to make this approach competitive with others.

### $V_{us}$

$|V_{us}|$  may be determined from kaon decays, hyperon decays, and tau decays. Previous determinations have most often used  $K\ell 3$  decays:

$$\Gamma_{K\ell 3} = \frac{G_F^2 M_K^5}{192\pi^3} S_{EW} (1 + \delta_K^\ell + \delta_{SU2}) C^2 |V_{us}|^2 f_+^2(0) I_K^\ell. \quad (9)$$

Here,  $\ell$  refers to either  $e$  or  $\mu$ ,  $G_F$  is the Fermi constant,  $M_K$  is the kaon mass,  $S_{EW}$  is the short-distance radiative correction,  $\delta_K^\ell$  is the mode-dependent long-distance radiative correction,  $f_+(0)$  is the calculated form factor at zero momentum transfer for the  $\ell\nu$  system, and  $I_K^\ell$  is the phase-space integral, which

depends on measured semileptonic form factors. For charged kaon decays,  $\delta_{SU2}$  is the deviation from one of the ratio of  $f_+(0)$  for the charged to neutral kaon decay; it is zero for the neutral kaon.  $C^2$  is 1 (1/2) for neutral (charged) kaon decays. Most determinations of  $|V_{us}|$  have been based only on  $K \rightarrow \pi e \nu$  decays;  $K \rightarrow \pi \mu \nu$  decays have not been used because of large uncertainties in  $I_K^\mu$ . The experimental measurements are the semileptonic decay widths (based on the semileptonic branching fractions and lifetime) and form factors (allowing calculation of the phase space integrals). Theory is needed for  $S_{EW}$ ,  $\delta_K^\ell$ ,  $\delta_{SU2}$ , and  $f_+(0)$ .

Many new measurements during the last few years have resulted in a significant shift in  $V_{us}$ . Most importantly, recent measurements of the  $K \rightarrow \pi e \nu$  branching fractions are significantly different than earlier PDG averages, probably as a result of inadequate treatment of radiation in older experiments. This effect was first observed by BNL E865 [16] in the charged kaon system and then by KTeV [17,18] in the neutral kaon system; subsequent measurements were made by KLOE [19–22], NA48 [23–25], and ISTRA+ [26]. Current averages (*e.g.*, by the PDG [27] or Flavianet [28]) of the semileptonic branching fractions are based only on recent, high-statistics experiments where the treatment of radiation is clear. In addition to measurements of branching fractions, new measurements of lifetimes [29] and form factors [30–34], have resulted in improved precision for all of the experimental inputs to  $V_{us}$ . Precise measurements of form factors for  $K_{\mu 3}$  decay now make it possible to use both semileptonic decay modes to extract  $V_{us}$ .

Following the analysis of the Flavianet group [28], one finds the values of  $|V_{us}|f_+(0)$  in Table 1. The average of these measurements gives

$$f_+(0)|V_{us}| = 0.21664(48). \quad (10)$$

Figure 1 shows a comparison of these results with the PDG evaluation from 2002 [35], as well as  $f_+(0)(1 - |V_{ud}|^2 - |V_{ub}|^2)^{1/2}$ , the expectation for  $f_+(0)|V_{us}|$  assuming unitarity, based on  $|V_{ud}| = 0.9742 \pm 0.0003$ ,  $|V_{ub}| = (3.6 \pm 0.7) \times 10^{-3}$ , and the

lattice calculation of  $f_+(0) = 0.9644 \pm 0.0049$  [36] ( Lattice calculations of  $f_+(0)$  have improved significantly in recent years, and therefore replace the classic calculation of Leutwyler and Roos [37]. ) Combining the result in Eq. (10) with the above value of  $f_+(0)$  gives

$$|V_{us}| = \lambda = 0.2246 \pm 0.0012. \quad (11)$$

**Table 1:**  $|V_{us}|f_+(0)$  from  $K_{\ell 3}$ .

Decay Mode	$ V_{us} f_+(0)$
$K^\pm e 3$	$0.2173 \pm 0.0008$
$K^\pm \mu 3$	$0.2176 \pm 0.0011$
$K_L e 3$	$0.2163 \pm 0.0006$
$K_L \mu 3$	$0.2168 \pm 0.0007$
$K_S e 3$	$0.2154 \pm 0.0013$
Average	$0.2166 \pm 0.0005$

A value of  $V_{us}$  can also be obtained from a comparison of the radiative inclusive decay rates for  $K \rightarrow \mu\nu(\gamma)$  and  $\pi \rightarrow \mu\nu(\gamma)$  combined with a lattice gauge theory calculation of  $f_K/f_\pi$  via [42]

$$\frac{|V_{us}|f_K}{|V_{ud}|f_\pi} = 0.2387(4) \left[ \frac{\Gamma(K \rightarrow \mu\nu(\gamma))}{\Gamma(\pi \rightarrow \mu\nu(\gamma))} \right]^{\frac{1}{2}} \quad (12)$$

with the small error coming from electroweak radiative corrections. Employing

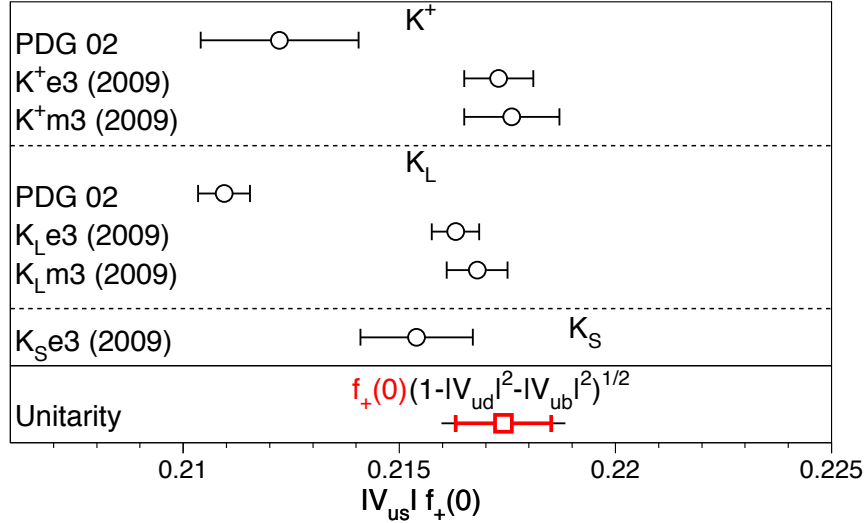
$$\frac{\Gamma(K \rightarrow \mu\nu(\gamma))}{\Gamma(\pi \rightarrow \mu\nu(\gamma))} = 1.3337(46), \quad (13)$$

which averages in the KLOE result [43],  $B(K \rightarrow \mu\nu(\gamma)) = 63.66(9)(15)\%$  and [44]

$$f_K/f_\pi = 1.189(7) \quad (14)$$

along with the value of  $V_{ud}$  in Eq. (4) leads to

$$|V_{us}| = 0.2259(5)(13). \quad (15)$$



**Figure 1:** Comparison of determinations of  $|V_{us}|f_+(0)$  from this review (labeled 2009), from the PDG 2002, and with the prediction from unitarity using  $|V_{ud}|$  and the lattice calculation of  $f_+(0)$  [36]. For  $f_+(0)(1 - |V_{ud}|^2 - |V_{ub}|^2)^{1/2}$ , the inner error bars are from the quoted uncertainty in  $f_+(0)$ ; the total uncertainties include the  $|V_{ud}|$  and  $|V_{ub}|$  errors.

It should be mentioned that hyperon decay fits suggest [5]

$$|V_{us}| = 0.2250(27) \text{ Hyperon Decays} \quad (16)$$

modulo SU(3) breaking effects that could shift that value up or down. We note that a recent representative effort [45] that incorporates SU(3) breaking found  $V_{us} = 0.226(5)$ . Similarly, inclusive strangeness changing tau decays give [46]

$$|V_{us}| = 0.2208(34) \text{ Tau Decays} \quad (17)$$

where the central value depends on the strange quark mass. However, a recent BaBar study [47] of  $\tau \rightarrow K\nu/\tau \rightarrow \pi\nu$  using the lattice value of  $f_K/f_\pi$  from Eq. (14) finds  $V_{us} = 0.2255(24)$ , in good agreement with other determinations.

Employing the value of  $V_{ud}$  in Eq. (4) and  $V_{us} = 0.2252(9)$ , the average of the  $K\ell 3$  (Eq. (11)) and  $K\mu 2$  (Eq. (15)) determinations of  $V_{us}$ , leads to the unitarity consistency check

$$|V_{ud}|^2 + |V_{us}|^2 + |V_{ub}|^2 = 0.9999(4)(4). \quad (18)$$

where the first error is the uncertainty from  $|V_{ud}|^2$  and the second error is the uncertainty from  $|V_{us}|^2$ .

### CKM Unitarity Constraints

The current good experimental agreement with unitarity,  $|V_{ud}|^2 + |V_{us}|^2 + |V_{ub}|^2 = 0.9999(6)$ , provides strong confirmation of Standard Model radiative corrections (which range between 3-4% depending on the nucleus used) at better than the 50 sigma level [48]. In addition, it implies constraints on “New Physics” effects at both the tree and quantum loop levels. Those effects could be in the form of contributions to nuclear beta decays,  $K$  decays and/or muon decays, with the last of these providing normalization via the muon lifetime [49], which is used to obtain the Fermi constant,  $G_\mu = 1.166371(6) \times 10^{-5} \text{GeV}^{-2}$ .

In the following sections, we illustrate the implications of CKM unitarity for (1) exotic muon decays [50] (beyond ordinary muon decay  $\mu^+ \rightarrow e^+ \nu_e \bar{\nu}_\mu$ ) and (2) new heavy quark mixing  $V_{uD}$  [51]. Other examples in the literature [52,53] include  $Z_\chi$  boson quantum loop effects, supersymmetry, leptoquarks, compositeness etc.

#### Exotic Muon Decays

If additional lepton flavor violating decays such as  $\mu^+ \rightarrow e^+ \bar{\nu}_e \nu_\mu$  (wrong neutrinos) occur, they would cause confusion in searches for neutrino oscillations at, for example, muon storage rings/neutrino factories or other neutrino sources from muon decays. Calling the rate for all such decays  $\Gamma(\text{exotic } \mu \text{ decays})$ , they should be subtracted before the extraction of  $G_\mu$  and normalization of the CKM matrix. Since that is not done and unitarity works, one has (at one-sided 95% CL)

$$|V_{ud}|^2 + |V_{us}|^2 + |V_{ub}|^2 = 1 - BR(\text{exotic } \mu \text{ decays}) \geq 0.9989 \quad (19)$$

or

$$BR(\text{exotic } \mu \text{ decays}) < 0.001 . \quad (20)$$

This bound is a factor of 10 better than the direct experimental bound on  $\mu^+ \rightarrow e^+ \bar{\nu}_e \nu_\mu$ .

### New Heavy Quark Mixing

Heavy  $D$  quarks naturally occur in fourth quark generation models and some heavy quark “new physics” scenarios such as  $E_6$  grand unification. Their mixing with ordinary quarks gives rise to  $V_{ud}$  which is constrained by unitarity (one sided 95% CL)

$$\begin{aligned} |V_{ud}|^2 + |V_{us}|^2 + |V_{ub}|^2 &= 1 - |V_{uD}|^2 > 0.9989 \\ |V_{uD}| &< 0.03 . \end{aligned} \quad (21)$$

A similar constraint applies to heavy neutrino mixing and the couplings  $V_{\mu N}$  and  $V_{eN}$ .

### References

1. N. Cabibbo, Phys. Rev. Lett. **10**, 531 (1963).
2. M. Kobayashi and T. Maskawa, Prog. Theor. Phys. **49**, 652 (1973).
3. L. Wolfenstein, Phys. Rev. Lett. **51**, 1945 (1983).
4. S. Eidelman *et al.*, [Particle Data Group], Phys. Lett. **B592**, 1 (2004).
5. N. Cabibbo, E.C. Swallow, and R. Winston, Phys. Rev. Lett. **92**, 251803 (2004) [[hep-ph/0307214](#)].
6. J.C. Hardy and I.S. Towner, Phys. Rev. Lett. **94**, 092502 (2005) [[nucl-th/0412050](#)].
7. W.J. Marciano and A. Sirlin, Phys. Rev. Lett. **71**, 3629 (1993).
8. A. Czarnecki, W.J. Marciano, and A. Sirlin, Phys. Rev. **D70**, 093006 (2004) [[hep-ph/0406324](#)].
9. W.J. Marciano and A. Sirlin, Phys. Rev. Lett. **96**, 032002 (2006) [[hep-ph/0510099](#)].
10. J.C. Hardy and I.S. Towner, Phys. Rev. **C77**, 025501 (2008).
11. J.C. Hardy and I.S. Towner, Phys. Rev. **C79**, 055502 (2009).
12. G.A. Miller and A. Schwenk, Phys. Rev. **C78**, 035501 (2008); N. Auerbach, Phys. Rev. **C79**, 035502 (2009); H.



- Liang, N. Van Giai and J. Meng, Phys. Rev. **C79**, 064316 (2009).
13. A. Serebrov *et al.*, Phys. Lett. **B605**, 72 (2005) [nucl-ex/0408009].
  14. H. Abele, Prog. in Part. Nucl. Phys. **60**, 1 (2008).
  15. D. Pocanic *et al.*, Phys. Rev. Lett. **93**, 181803 (2004) [hep-ex/0312030].
  16. A. Sher *et al.*, Phys. Rev. Lett. **91**, 261802 (2003).
  17. T. Alexopoulos *et al.*, [KTeV Collab.], Phys. Rev. Lett. **93**, 181802 (2004) [hep-ex/0406001].
  18. T. Alexopoulos *et al.*, [KTeV Collab.], Phys. Rev. **D70**, 092006 (2004) [hep-ex/0406002].
  19. F. Ambrosino *et al.*, [KLOE Collab.], Phys. Lett. **B632**, 43 (2006) [hep-ex/0508027].
  20. F. Ambrosino *et al.*, [KLOE Collab.], Phys. Lett. **B638**, 140 (2006) [hep-ex/0603041].
  21. F. Ambrosino *et al.*, [KLOE Collab.], Phys. Lett. **B636**, 173 (2006) [hep-ex/0601026].
  22. F. Ambrosino *et al.*, [KLOE Collab.], PoS **HEP2005**, 287 (2006) [Frascati Phys. Ser. **41**, 69 (2006)] [hep-ex/0510028].
  23. A. Lai *et al.*, [NA48 Collab.], Phys. Lett. **B602**, 41 (2004) [hep-ex/0410059].
  24. A. Lai *et al.*, [NA48 Collab.], Phys. Lett. **B645**, 26 (2007) [hep-ex/0611052].
  25. J.R. Batley *et al.*, [NA48/2 Collab.], Eur. Phys. J. C **50**, 329 (2007) [hep-ex/0702015].
  26. V.I. Romanovsky *et al.*, [hep-ex/0704.2052].
  27. C. Amsler *et al.*, [Particle Data Group], Phys. Lett. **B667**, 1 (2008).
  28. Flavianet Working Group on Precise SM Tests in K Decays, <http://www.lnf.infn.it/wg/vus>. For a recent detailed review, see M. Antonelli *et al.*, [hep-ph/0907.5386].
  29. F. Ambrosino *et al.*, [KLOE Collab.], Phys. Lett. **B626**, 15 (2005) [hep-ex/0507088].
  30. T. Alexopoulos *et al.*, [KTeV Collab.], Phys. Rev. **D70**, 092007 (2004) [hep-ex/0406003].
  31. E. Abouzaid *et al.*, [KTeV Collab.], Phys. Rev. **D74**, 097101 (2006) [hep-ex/0608058].
  32. F. Ambrosino *et al.*, [KLOE Collab.], Phys. Lett. **B636**, 166 (2006) [hep-ex/0601038].
  33. A. Lai *et al.*, [NA48 Collab.], Phys. Lett. **B604**, 1 (2004) [hep-ex/0410065].

34. O.P. Yushchenko *et al.*, Phys. Lett. **B589**, 111 (2004) [hep-ex/0404030].
35. K. Hagiwara *et al.*, [Particle Data Group], Phys. Rev. **D66**, 1 (2002).
36. P.A. Boyle *et al.*, Phys. Rev. Lett. **100**, 141601 (2008).
37. H. Leutwyler and M. Roos, Z. Phys. **C25**, 91 (1984).
38. D. Becirevic *et al.*, Nucl. Phys. **B705**, 339 (2005) [hep-ph/0403217].
39. J. Bijnens and P. Talavera, Nucl. Phys. **B669**, 341 (2003).
40. V. Cirigliano *et al.*, JHEP **0504**, 006 (2005) [hep-ph/0503108].
41. M. Jamin, J.A. Oller, and A. Pich, JHEP **02**, 047 (2004).
42. W.J. Marciano, Phys. Rev. Lett. **93**, 231803 (2004) [hep-ph/0402299].
43. F. Ambrosino *et al.*, [KLOE Collab.], Phys. Lett. **B632**, 76 (2006) [hep-ex/0509045].
44. E. Follana *et al.*, Phys. Rev. Lett. **100**, 062002 (2008).
45. V. Mateu and A. Pich, JHEP **0510**, 041 (2005) [hep-ph/0509045].
46. E. Gamiz *et al.*, Phys. Rev. Lett. **94**, 011803 (2005) [hep-ph/0408044]; E. Gamiz *et al.*, arXiv:0709.0282 [hep-ph].
47. B. Aubert *et al.*, [BaBar Collaboration], [hep-ex/0912.0242].
48. A. Sirlin, Rev. Mod. Phys. **50**, 573 (1978).
49. D.B. Chitwood *et al.*, Phys. Rev. Lett. **99**, 032001 (2007).
50. K.S. Babu and S. Pakvasa, hep-ph/0204236.
51. W. Marciano and A. Sirlin, Phys. Rev. Lett. **56**, 22 (1986); P. Langacker and D. London, Phys. Rev. **D38**, 886 (1988).
52. W. Marciano and A. Sirlin, Phys. Rev. **D35**, 1672 (1987).
53. R. Barbieri *et al.*, Phys. Lett. **156B**, 348 (1985); K. Hagiwara *et al.*, Phys. Rev. Lett. **75**, 3605 (1995); A. Kurylov and M. Ramsey-Musolf, Phys. Rev. Lett. **88**, 071804 (2000).