

# Magnetic Monopole Searches

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## Monopole Production Cross Section — Accelerator Searches

<i>X-SECT</i> (cm <sup>2</sup> )	<i>MASS</i> (GeV)	<i>CHG</i> (g)	<i>ENERGY</i> (GeV)	<i>BEAM</i>	<i>DOCUMENT ID</i>	<i>TECN</i>
<5E-38	45-102	1	206	e <sup>+</sup> e <sup>-</sup>	1 ABBIENDI	08 OPAL
<0.2E-36	200-700	1	1960	p $\bar{p}$	2 ABULENCIA	06K CNTR
< 2.E-36		1	300	e <sup>+</sup> p	3,4 AKTAS	05A INDU
< 0.2 E-36		2	300	e <sup>+</sup> p	3,4 AKTAS	05A INDU
< 0.09E-36		3	300	e <sup>+</sup> p	3,4 AKTAS	05A INDU
< 0.05E-36		≥ 6	300	e <sup>+</sup> p	3,4 AKTAS	05A INDU
< 2.E-36		1	300	e <sup>+</sup> p	3,5 AKTAS	05A INDU
< 0.2E-36		2	300	e <sup>+</sup> p	3,5 AKTAS	05A INDU
< 0.07E-36		3	300	e <sup>+</sup> p	3,5 AKTAS	05A INDU
< 0.06E-36		≥ 6	300	e <sup>+</sup> p	3,5 AKTAS	05A INDU
< 0.6E-36	>265	1	1800	p $\bar{p}$	6 KALBFLEISCH 04	INDU
< 0.2E-36	>355	2	1800	p $\bar{p}$	6 KALBFLEISCH 04	INDU
< 0.07E-36	>410	3	1800	p $\bar{p}$	6 KALBFLEISCH 04	INDU
< 0.2E-36	>375	6	1800	p $\bar{p}$	6 KALBFLEISCH 04	INDU
< 0.7E-36	>295	1	1800	p $\bar{p}$	7,8 KALBFLEISCH 00	INDU
< 7.8E-36	>260	2	1800	p $\bar{p}$	7,8 KALBFLEISCH 00	INDU
< 2.3E-36	>325	3	1800	p $\bar{p}$	7,9 KALBFLEISCH 00	INDU
< 0.11E-36	>420	6	1800	p $\bar{p}$	7,9 KALBFLEISCH 00	INDU
<0.65E-33	<3.3	≥ 2	11A	197Au	10 HE	97
<1.90E-33	<8.1	≥ 2	160A	208Pb	10 HE	97
<3.E-37	<45.0	1.0	88-94	e <sup>+</sup> e <sup>-</sup>	PINFOLD	93 PLAS
<3.E-37	<41.6	2.0	88-94	e <sup>+</sup> e <sup>-</sup>	PINFOLD	93 PLAS
<7.E-35	<44.9	0.2-1.0	89-93	e <sup>+</sup> e <sup>-</sup>	KINOSHITA	92 PLAS
<2.E-34	<850	≥ 0.5	1800	p $\bar{p}$	BERTANI	90 PLAS
<1.2E-33	<800	≥ 1	1800	p $\bar{p}$	PRICE	90 PLAS
<1.E-37	<29	1	50-61	e <sup>+</sup> e <sup>-</sup>	KINOSHITA	89 PLAS
<1.E-37	<18	2	50-61	e <sup>+</sup> e <sup>-</sup>	KINOSHITA	89 PLAS
<1.E-38	<17	<1	35	e <sup>+</sup> e <sup>-</sup>	BRAUNSCH...	88B CNTR
<8.E-37	<24	1	50-52	e <sup>+</sup> e <sup>-</sup>	KINOSHITA	88 PLAS
<1.3E-35	<22	2	50-52	e <sup>+</sup> e <sup>-</sup>	KINOSHITA	88 PLAS
<9.E-37	<4	<0.15	10.6	e <sup>+</sup> e <sup>-</sup>	GENTILE	87 CLEO
<3.E-32	<800	≥ 1	1800	p $\bar{p}$	PRICE	87 PLAS
<3.E-38		<3	29	e <sup>+</sup> e <sup>-</sup>	FRYBERGER	84 PLAS
<1.E-31		1,3	540	p $\bar{p}$	AUBERT	83B PLAS
<4.E-38	<10	<6	34	e <sup>+</sup> e <sup>-</sup>	MUSSET	83 PLAS
<8.E-36	<20		52	pp	11 DELL	82 CNTR
<9.E-37	<30	<3	29	e <sup>+</sup> e <sup>-</sup>	KINOSHITA	82 PLAS
<1.E-37	<20	<24	63	pp	CARRIGAN	78 CNTR
<1.E-37	<30	<3	56	pp	HOFFMANN	78 PLAS
			62	pp	11 DELL	76 SPRK
<4.E-33			300	p	11 STEVENS	76B SPRK

<1.E-40	<5	<2	70	$p$	<sup>12</sup> ZRELOV	76	CNTR
<2.E-30			300	$n$	<sup>11</sup> BURKE	75	OSPK
<1.E-38			8	$\nu$	<sup>13</sup> CARRIGAN	75	HLBC
<5.E-43	<12	<10	400	$p$	EBERHARD	75 <sup>B</sup>	INDU
<2.E-36	<30	<3	60	$p\bar{p}$	GIACOMELLI	75	PLAS
<5.E-42	<13	<24	400	$p$	CARRIGAN	74	CNTR
<6.E-42	<12	<24	300	$p$	CARRIGAN	73	CNTR
<2.E-36		1	0.001	$\gamma$	<sup>12</sup> BARTLETT	72	CNTR
<1.E-41	<5		70	$p$	GUREVICH	72	EMUL
<1.E-40	<3	<2	28	$p$	AMALDI	63	EMUL
<2.E-40	<3	<2	30	$p$	PURCELL	63	CNTR
<1.E-35	<3	<4	28	$p$	FIDECARO	61	CNTR
<2.E-35	<1	1	6	$p$	BRADNER	59	EMUL

<sup>1</sup> ABBIENDI 08 assume production of spin 1/2 monopoles with effective charge  $g\beta$  ( $n=1$ ), via  $e^+e^- \rightarrow \gamma^* \rightarrow M\bar{M}$ , so that the cross section is proportional to  $(1 + \cos^2\theta)$ . There is no  $z$  information for such highly saturated tracks, so a parabolic track in the jet chamber is projected onto the  $xy$  plane. Charge per hit in the chamber produces a clean separation of signal and background.

<sup>2</sup> ABULENCIA 06K searches for high-ionizing signals in CDF central outer tracker and time-of-flight detector. For Drell-Yan  $M\bar{M}$  production, the cross section limit implies  $M > 360$  GeV at 95% CL.

<sup>3</sup> AKTAS 05A model-dependent limits as a function of monopole mass shown for arbitrary mass of 60 GeV. Based on search for stopped monopoles in the H1 Al beam pipe.

<sup>4</sup> AKTAS 05A limits with assumed elastic spin 0 monopole pair production.

<sup>5</sup> AKTAS 05A limits with assumed inelastic spin 1/2 monopole pair production.

<sup>6</sup> KALBFLEISCH 04 reports searches for stopped magnetic monopoles in Be, Al, and Pb samples obtained from discarded material from the upgrading of DØ and CDF. A large-aperture warm-bore cryogenic detector was used. The approach was an extension of the methods of KALBFLEISCH 00. Cross section results moderately model dependent; interpretation as a mass lower limit depends on possibly invalid perturbation expansion.

<sup>7</sup> KALBFLEISCH 00 used an induction method to search for stopped monopoles in pieces of the DØ (FNAL) beryllium beam pipe and in extensions to the drift chamber aluminum support cylinder. Results are model dependent.

<sup>8</sup> KALBFLEISCH 00 result is for aluminum.

<sup>9</sup> KALBFLEISCH 00 result is for beryllium.

<sup>10</sup> HE 97 used a lead target and barium phosphate glass detectors. Cross-section limits are well below those predicted via the Drell-Yan mechanism.

<sup>11</sup> Multiphoton events.

<sup>12</sup> Cherenkov radiation polarization.

<sup>13</sup> Re-examines CERN neutrino experiments.

## Monopole Production — Other Accelerator Searches

MASS (GeV)	CHG ( $g$ )	SPIN	ENERGY (GeV)	BEAM	DOCUMENT ID	TECN
> 610	$\geq 1$	0	1800	$p\bar{p}$	<sup>14</sup> ABBOTT	98K D0
> 870	$\geq 1$	1/2	1800	$p\bar{p}$	<sup>14</sup> ABBOTT	98K D0
>1580	$\geq 1$	1	1800	$p\bar{p}$	<sup>14</sup> ABBOTT	98K D0
> 510			88-94	$e^+e^-$	<sup>15</sup> ACCIARRI	95C L3

- <sup>14</sup> ABBOTT 98K search for heavy pointlike Dirac monopoles via central production of a pair of photons with high transverse energies.  
<sup>15</sup> ACCIARRI 95C finds a limit  $B(Z \rightarrow \gamma\gamma) < 0.8 \times 10^{-5}$  (which is possible via a monopole loop) at 95% CL and sets the mass limit via a cross section model.

## Monopole Flux — Cosmic Ray Searches

“Caty” in the charge column indicates a search for monopole-catalyzed nucleon decay.

<u>FLUX</u> ( $\text{cm}^{-2}\text{sr}^{-1}\text{s}^{-1}$ )	<u>MASS</u> (GeV)	<u>CHG</u> (g)	<u>COMMENTS</u> ( $\beta = v/c$ )	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>
<1E-19		1	$\gamma > 1\text{E}10$	0	16 DETRIXHE	11 ANIT
<3.8E-17		1	$\beta > 0.76$	0	17 ABBASI	10A AMND
<1.3E-15	1E4<M<5E13	1	$\beta > 0.05$	0	18 BALESTRA	08 PLAS
<0.65E-15	>5E13	1	$\beta > 0.05$	0	18 BALESTRA	08 PLAS
<1E-18		1	$\gamma > 1\text{ E}8$	0	16 HOGAN	08 RICE
<1.4E-16		1	$1.1\text{E}-4 < \beta < 1$	0	19 AMBROSIO	02B MCRO
<3E-16		Caty	$1.1\text{E}-4 < \beta < 5\text{E}-3$	0	20 AMBROSIO	02C MCRO
<1.5E-15		1	$5\text{E}-3 < \beta < 0.99$	0	21 AMBROSIO	02D MCRO
<1E-15		1	$1.1 \times 10^{-4} - 0.1$	0	22 AMBROSIO	97 MCRO
<5.6E-15		1	$(0.18-3.0)\text{E}-3$	0	23 AHLEN	94 MCRO
<2.7E-15		Caty	$\beta \sim 1 \times 10^{-3}$	0	24 BECKER-SZ...	94 IMB
<8.7E-15		1	$> 2.\text{E}-3$	0	THRON	92 SOUD
<4.4E-12		1	all $\beta$	0	GARDNER	91 INDU
<7.2E-13		1	all $\beta$	0	HUBER	91 INDU
<3.7E-15	>E12	1	$\beta = 1.\text{E}-4$	0	25 ORITO	91 PLAS
<3.2E-16	>E10	1	$\beta > 0.05$	0	25 ORITO	91 PLAS
<3.2E-16	>E10-E12	2,3		0	25 ORITO	91 PLAS
<3.8E-13		1	all $\beta$	0	BERMON	90 INDU
<5.E-16		Caty	$\beta < 1.\text{E}-3$	0	24 BEZRUKOV	90 CHER
<1.8E-14		1	$\beta > 1.1\text{E}-4$	0	26 BUCKLAND	90 HEPT
<1E-18			$3.\text{E}-4 < \beta < 1.5\text{E}-3$	0	27 GHOSH	90 MICA
<7.2E-13		1	all $\beta$	0	HUBER	90 INDU
<5.E-12	>E7	1	$3.\text{E}-4 < \beta < 5.\text{E}-3$	0	BARISH	87 CNTR
<1.E-13		Caty	$1.\text{E}-5 < \beta < 1$	0	24 BARTELT	87 SOUD
<1.E-10		1	all $\beta$	0	EBISU	87 INDU
<2.E-13			$1.\text{E}-4 < \beta < 6.\text{E}-4$	0	MASEK	87 HEPT
<2.E-14			$4.\text{E}-5 < \beta < 2.\text{E}-4$	0	NAKAMURA	87 PLAS
<2.E-14			$1.\text{E}-3 < \beta < 1$	0	NAKAMURA	87 PLAS
<5.E-14			$9.\text{E}-4 < \beta < 1.\text{E}-2$	0	SHEPKO	87 CNTR
<2.E-13			$4.\text{E}-4 < \beta < 1$	0	TSUKAMOTO	87 CNTR
<5.E-14		1	all $\beta$	1	28 CAPLIN	86 INDU
<5.E-12		1		0	CROMAR	86 INDU
<1.E-13		1	$7.\text{E}-4 < \beta$	0	HARA	86 CNTR
<7.E-11		1	all $\beta$	0	INCANDELA	86 INDU
<1.E-18			$4.\text{E}-4 < \beta < 1.\text{E}-3$	0	27 PRICE	86 MICA
<5.E-12		1		0	BERMON	85 INDU
<6.E-12		1		0	CAPLIN	85 INDU
<6.E-10		1		0	EBISU	85 INDU
<3.E-15		Caty	$5.\text{E}-5 \leq \beta \leq 1.\text{E}-3$	0	24 KAJITA	85 KAMI
<2.E-21		Caty	$\beta < 1.\text{E}-3$	0	24,29 KAJITA	85 KAMI
<3.E-15		Caty	$1.\text{E}-3 < \beta < 1.\text{E}-1$	0	24 PARK	85B CNTR

<5.E-12	1	$1.E-4 < \beta < 1$	0	BATTISTONI	84	NUSX
<7.E-12	1		0	INCANDELA	84	INDU
<7.E-13	1	$3.E-4 < \beta$	0	<sup>26</sup> KAJINO	84	CNTR
<2.E-12	1	$3.E-4 < \beta < 1.E-1$	0	KAJINO	84 <sup>B</sup>	CNTR
<6.E-13	1	$5.E-4 < \beta < 1$	0	KAWAGOE	84	CNTR
<2.E-14		$1.E-3 < \beta$	0	<sup>24</sup> KRISHNA...	84	CNTR
<4.E-13	1	$6.E-4 < \beta < 2.E-3$	0	LISS	84	CNTR
<1.E-16		$3.E-4 < \beta < 1.E-3$	0	<sup>27</sup> PRICE	84	MICA
<1.E-13	1	$1.E-4 < \beta$	0	PRICE	84 <sup>B</sup>	PLAS
<4.E-13	1	$6.E-4 < \beta < 2.E-3$	0	TARLE	84	CNTR
			7	<sup>30</sup> ANDERSON	83	EMUL
<4.E-13	1	$1.E-2 < \beta < 1.E-3$	0	BARTELT	83 <sup>B</sup>	CNTR
<1.E-12	1	$7.E-3 < \beta < 1$	0	BARWICK	83	PLAS
<3.E-13	1	$1.E-3 < \beta < 4.E-1$	0	BONARELLI	83	CNTR
<3.E-12	Caty	$5.E-4 < \beta < 5.E-2$	0	<sup>24</sup> BOSETTI	83	CNTR
<4.E-11	1		0	CABRERA	83	INDU
<5.E-15	1	$1.E-2 < \beta < 1$	0	DOKE	83	PLAS
<8.E-15	Caty	$1.E-4 < \beta < 1.E-1$	0	<sup>24</sup> ERREDE	83	IMB
<5.E-12	1	$1.E-4 < \beta < 3.E-2$	0	GROOM	83	CNTR
<2.E-12		$6.E-4 < \beta < 1$	0	MASHIMO	83	CNTR
<1.E-13	1	$\beta=3.E-3$	0	ALEXEYEV	82	CNTR
<2.E-12	1	$7.E-3 < \beta < 6.E-1$	0	BONARELLI	82	CNTR
6.E-10	1	all $\beta$	1	<sup>31</sup> CABRERA	82	INDU
<2.E-11		$1.E-2 < \beta < 1.E-1$	0	MASHIMO	82	CNTR
<2.E-15		concentrator	0	BARTLETT	81	PLAS
<1.E-13	>1	$1.E-3 < \beta$	0	KINOSHITA	81 <sup>B</sup>	PLAS
<5.E-11	<E17	$3.E-4 < \beta < 1.E-3$	0	ULLMAN	81	CNTR
<2.E-11		concentrator	0	BARTLETT	78	PLAS
1.E-1	>200	2	1	<sup>32</sup> PRICE	75	PLAS
<2.E-13		>2	0	FLEISCHER	71	PLAS
<1.E-19		>2 obsidian, mica	0	FLEISCHER	69 <sup>C</sup>	PLAS
<5.E-15	<15	<3 concentrator	0	CARITHERS	66	ELEC
<2.E-11		<1-3 concentrator	0	MALKUS	51	EMUL

<sup>16</sup> HOGAN 08 and DETRIXHE 11 limits on relativistic monopoles are based on nonobservation of radio Cherenkov signals at the South Pole. Limits are speed-dependent.

<sup>17</sup> ABBASI 10A was based on a Cherenkov signature in an array of optical modules which were sunk in the Antarctic ice cap. Limits are speed-dependent.

<sup>18</sup> BALESTRA 08 exposed of nuclear track detector modules totaling 400 m<sup>2</sup> for 4 years at the Chacaltaya Laboratory (5230 m) in search for intermediate-mass monopoles with  $\beta > 0.05$ . The analysis is mainly based on three CR39 modules. For  $M > 5 \times 10^{13}$  GeV there can be upward-going monopoles as well, hence the flux limit is half that obtained for less massive monopoles. Previous experiments (e.g. MACRO and OHYA (ORITO 91)) had set limits only for  $M > 1 \times 10^9$  GeV.

<sup>19</sup> AMBROSIO 02B direct search final result for  $m \geq 10^{17}$  GeV, based upon 4.2 to 9.5 years of running, depending upon the subsystem. Limit with CR39 track-etch detector extends the limit from  $\beta=4 \times 10^{-5}$  ( $3.1 \times 10^{-16}$  cm<sup>-2</sup> sr<sup>-1</sup> s<sup>-1</sup>) to  $\beta=1 \times 10^{-4}$  ( $2.1 \times 10^{-16}$  cm<sup>-2</sup> sr<sup>-1</sup> s<sup>-1</sup>). Limit curve in paper is piecewise continuous due to different detection techniques for different  $\beta$  ranges.

<sup>20</sup> AMBROSIO 02C limit for catalysis of nucleon decay with catalysis cross section of  $\approx 1$  mb. The flux limit increases by  $\sim 3$  at the higher  $\beta$  limit, and increases to  $1 \times 10^{-14}$  cm<sup>-2</sup> sr<sup>-1</sup> s<sup>-1</sup> if the catalysis cross section is 0.01 mb. Based upon 71193 hr of data with the streamer detector, with an acceptance of 4250 m<sup>2</sup> sr.

- <sup>21</sup> AMBROSIO 02D result for “more than two years of data.” Ionization search using several subsystems. Limit curve as a function of  $\beta$  not given. Included in AMBROSIO 02B.
- <sup>22</sup> AMBROSIO 97 global MACRO 90%CL is  $0.78 \times 10^{-15}$  at  $\beta=1.1 \times 10^{-4}$ , goes through a minimum at  $0.61 \times 10^{-15}$  near  $\beta=(1.1-2.7) \times 10^{-3}$ , then rises to  $0.84 \times 10^{-15}$  at  $\beta=0.1$ . The global limit in this region is below the Parker bound at  $10^{-15}$ . Less stringent limits are established for  $4 \times 10^{-5} < \beta < 1 \times 10^{-4}$ . Limits set by various triggers and different subdetectors are given in the paper. All limits assume a catalysis cross section smaller than a few mb.
- <sup>23</sup> AHLEN 94 limit for dyons extends down to  $\beta=0.9E-4$  and a limit of  $1.3E-14$  extends to  $\beta = 0.8E-4$ . Also see comment by PRICE 94 and reply of BARISH 94. One loophole in the AHLEN 94 result is that in the case of monopoles catalyzing nucleon decay, relativistic particles could veto the events. See AMBROSIO 97 for additional results.
- <sup>24</sup> Catalysis of nucleon decay; sensitive to assumed catalysis cross section.
- <sup>25</sup> ORITO 91 limits are functions of velocity. Lowest limits are given here.
- <sup>26</sup> Used DKMPR mechanism and Penning effect.
- <sup>27</sup> Assumes monopole attaches fermion nucleus.
- <sup>28</sup> Limit from combining data of CAPLIN 86, BERMON 85, INCANDELA 84, and CABRERA 83. For a discussion of controversy about CAPLIN 86 observed event, see GUY 87. Also see SCHOUTEN 87.
- <sup>29</sup> Based on lack of high- energy solar neutrinos from catalysis in the sun.
- <sup>30</sup> Anomalous long-range  $\alpha$  ( $^4\text{He}$ ) tracks.
- <sup>31</sup> CABRERA 82 candidate event has single Dirac charge within  $\pm 5\%$ .
- <sup>32</sup> ALVAREZ 75, FLEISCHER 75, and FRIEDLANDER 75 explain as fragmenting nucleus. EBERHARD 75 and ROSS 76 discuss conflict with other experiments. HAGSTROM 77 reinterprets as antinucleus. PRICE 78 reassesses.

### Monopole Flux — Astrophysics

<i>FLUX</i> ( $\text{cm}^{-2}\text{sr}^{-1}\text{s}^{-1}$ )	<i>MASS</i> (GeV)	<i>CHG</i> (g)	<i>COMMENTS</i> ( $\beta = v/c$ )	<i>EVTS</i>	<i>DOCUMENT ID</i>	<i>TECN</i>
<1.3E-20			faint white dwarf		<sup>33</sup> FREESE 99	ASTR
<1.E-16	E17	1	galactic field	0	<sup>34</sup> ADAMS 93	COSM
<1.E-23			Jovian planets		<sup>33</sup> ARAFUNE 85	ASTR
<1.E-16	E15		solar trapping	0	BRACCI 85B	ASTR
<1.E-18		1		0	<sup>33</sup> HARVEY 84	COSM
<3.E-23			neutron stars		KOLB 84	ASTR
<7.E-22			pulsars	0	<sup>33</sup> FREESE 83B	ASTR
<1.E-18	<E18	1	intergalactic field	0	<sup>33</sup> REPHAELI 83	COSM
<1.E-23			neutron stars	0	<sup>33</sup> DIMOPOUL... 82	COSM
<5.E-22			neutron stars	0	<sup>33</sup> KOLB 82	COSM
<5.E-15	>E21		galactic halo		SALPETER 82	COSM
<1.E-12	E19	1	$\beta=3.E-3$	0	<sup>35</sup> TURNER 82	COSM
<1.E-16		1	galactic field	0	PARKER 70	COSM

- <sup>33</sup> Catalysis of nucleon decay.
- <sup>34</sup> ADAMS 93 limit based on “survival and growth of a small galactic seed field” is  $10^{-16} (m/10^{17} \text{ GeV}) \text{ cm}^{-2} \text{ s}^{-1} \text{ sr}^{-1}$ . Above  $10^{17} \text{ GeV}$ , limit  $10^{-16} (10^{17} \text{ GeV}/m) \text{ cm}^{-2} \text{ s}^{-1} \text{ sr}^{-1}$  (from requirement that monopole density does not overclose the universe) is more stringent.
- <sup>35</sup> Re-evaluates PARKER 70 limit for GUT monopoles.

## Monopole Density — Matter Searches

<u>DENSITY</u>	<u>CHG (g)</u>	<u>MATERIAL</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>
<6.9E-6/gram	>1/3	Meteorites and other	0	JEON 95	INDU
<2.E-7/gram	>0.6	Fe ore	0	<sup>36</sup> EBISU 87	INDU
<4.6E-6/gram	> 0.5	deep schist	0	KOVALIK 86	INDU
<1.6E-6/gram	> 0.5	manganese nodules	0	<sup>37</sup> KOVALIK 86	INDU
<1.3E-6/gram	> 0.5	seawater	0	KOVALIK 86	INDU
>1.E+14/gram	>1/3	iron aerosols	>1	MIKHAILOV 83	SPEC
<6.E-4/gram		air, seawater	0	CARRIGAN 76	CNTR
<5.E-1/gram	>0.04	11 materials	0	CABRERA 75	INDU
<2.E-4/gram	>0.05	moon rock	0	ROSS 73	INDU
<6.E-7/gram	<140	seawater	0	KOLM 71	CNTR
<1.E-2/gram	<120	manganese nodules	0	FLEISCHER 69	PLAS
<1.E-4/gram	>0	manganese	0	FLEISCHER 69B	PLAS
<2.E-3/gram	<1-3	magnetite, meteor	0	GOTO 63	EMUL
<2.E-2/gram		meteorite	0	PETUKHOV 63	CNTR

<sup>36</sup> Mass  $1 \times 10^{14}$ – $1 \times 10^{17}$  GeV.

<sup>37</sup> KOVALIK 86 examined 498 kg of schist from two sites which exhibited clear mineralogical evidence of having been buried at least 20 km deep and held below the Curie temperature.

## Monopole Density — Astrophysics

<u>DENSITY</u>	<u>CHG (g)</u>	<u>MATERIAL</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>
<1.E-9/gram	1	sun, catalysis	0	<sup>38</sup> ARAFUNE 83	COSM
<6.E-33/nucl	1	moon wake	0	SCHATTEN 83	ELEC
<2.E-28/nucl		earth heat	0	CARRIGAN 80	COSM
<2.E-4/prot		42cm absorption	0	BRODERICK 79	COSM
<2.E-13/m <sup>3</sup>		moon wake	0	SCHATTEN 70	ELEC

<sup>38</sup> Catalysis of nucleon decay.

## REFERENCES FOR Magnetic Monopole Searches

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ABBIENDI 08	PL B663 37	G. Abbiendi <i>et al.</i>	(OPAL Collab.)
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