

**$\phi(1020)$**  $I^G(J^{PC}) = 0^-(1^{--})$  **$\phi(1020)$  MASS**

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>1019.461 \pm 0.016</math> OUR AVERAGE</b>				
1019.463 $\pm 0.061$	2.3M	<sup>1</sup> KOZYREV	18	CMD3 $e^+ e^- \rightarrow K^+ K^-, K_S^0 K_L^0$
1019.462 $\pm 0.042 \pm 0.056$	28k	<sup>2</sup> LEES	14H	BABR $e^+ e^- \rightarrow K_S^0 K_L^0 \gamma$
1019.51 $\pm 0.02 \pm 0.05$		<sup>3</sup> LEES	13Q	BABR $e^+ e^- \rightarrow K^+ K^- \gamma$
1019.30 $\pm 0.02 \pm 0.10$	105k	AKHMETSHIN 06	CMD2	$0.98-1.06 e^+ e^- \rightarrow \pi^+ \pi^- \pi^0$
1019.52 $\pm 0.05 \pm 0.05$	17.4k	AKHMETSHIN 05	CMD2	$0.60-1.38 e^+ e^- \rightarrow \eta \gamma$
1019.483 $\pm 0.011 \pm 0.025$	272k	<sup>4</sup> AKHMETSHIN 04	CMD2	$e^+ e^- \rightarrow K_L^0 K_S^0$
1019.42 $\pm 0.05$	1900k	<sup>5</sup> ACHASOV	01E	SND $e^+ e^- \rightarrow K^+ K^-, K_S K_L, \pi^+ \pi^- \pi^0$
1019.40 $\pm 0.04 \pm 0.05$	23k	AKHMETSHIN 01B	CMD2	$e^+ e^- \rightarrow \eta \gamma$
1019.36 $\pm 0.12$		<sup>6</sup> ACHASOV	00B	SND $e^+ e^- \rightarrow \eta \gamma$
1019.38 $\pm 0.07 \pm 0.08$	2200	<sup>7</sup> AKHMETSHIN 99F	CMD2	$e^+ e^- \rightarrow \pi^+ \pi^- \geq 2\gamma$
1019.51 $\pm 0.07 \pm 0.10$	11169	AKHMETSHIN 98	CMD2	$e^+ e^- \rightarrow \pi^+ \pi^- \pi^0$
1019.5 $\pm 0.4$		BARBERIS	98	OMEG $450 pp \rightarrow pp 2K^+ 2K^-$
1019.42 $\pm 0.06$	55600	AKHMETSHIN 95	CMD2	$e^+ e^- \rightarrow \text{hadrons}$
1019.7 $\pm 0.3$	2012	DAVENPORT	86	MPSF $400 \mu\text{A} \rightarrow 4KX$
1019.7 $\pm 0.1 \pm 0.1$	5079	ALBRECHT	85D	ARG $10 e^+ e^- \rightarrow K^+ K^- X$
1019.3 $\pm 0.1$	1500	ARENTON	82	AEMS $11.8 \text{ polar. } pp \rightarrow KK$
1019.67 $\pm 0.17$	25080	<sup>8</sup> PELLINEN	82	RVUE
1019.52 $\pm 0.13$	3681	BUKIN	78C	OLYA $e^+ e^- \rightarrow \text{hadrons}$
<b>• • • We do not use the following data for averages, fits, limits, etc. • • •</b>				
1018.4 $\pm 0.5 \pm 0.1$		<sup>9</sup> ALBRECHT	20	CBAR $0.9 \bar{p}p \rightarrow K^+ K^- \pi^0$
1019.21 $\pm 0.04 \pm 0.03$		<sup>10</sup> HOID	20	RVUE $e^+ e^- \rightarrow \pi^0 \gamma$
1019.54 $\pm 0.10 \pm 0.51$		<sup>11</sup> AAIJ	19H	LHCb $pp \rightarrow D^\pm X$
1019.20 $\pm 0.02 \pm 0.01$		<sup>12</sup> HOFERICHT...	19	RVUE $e^+ e^- \rightarrow \pi^+ \pi^- \pi^0$
1019.469 $\pm 0.061$	1.7M	KOZYREV	18	CMD3 $e^+ e^- \rightarrow K^+ K^-$
1019.457 $\pm 0.061$	610k	KOZYREV	16	CMD3 $e^+ e^- \rightarrow K_S^0 K_L^0$
1019.48 $\pm 0.01$		LEES	13F	BABR $D^+ \rightarrow K^+ K^- \pi^+$
1019.441 $\pm 0.008 \pm 0.080$	542k	<sup>13</sup> AKHMETSHIN 08	CMD2	$1.02 e^+ e^- \rightarrow K^+ K^-$
1019.63 $\pm 0.07$	12540	<sup>14</sup> AUBERT,B	05J	BABR $D^0 \rightarrow \bar{K}^0 K^+ K^-$
1019.8 $\pm 0.7$		ARMSTRONG	86	OMEG $85 \pi^+ / pp \rightarrow \pi^+ / p4Kp$
1020.1 $\pm 0.11$	5526	<sup>14</sup> ATKINSON	86	OMEG $20-70 \gamma p$
1019.7 $\pm 1.0$		BEBEK	86	CLEO $e^+ e^- \rightarrow \Upsilon(4S)$
1019.411 $\pm 0.008$	642k	<sup>15</sup> DIJKSTRA	86	SPEC $100-200 \pi^\pm, \bar{p}, p, K^\pm, \text{on Be}$

1020.9	$\pm 0.2$		<sup>14</sup> FRAME	86	OMEG	13	$K^+ p \rightarrow \phi K^+ p$
1021.0	$\pm 0.2$		<sup>14</sup> ARMSTRONG	83B	OMEG	18.5	$K^- p \rightarrow K^- K^+ \Lambda$
1020.0	$\pm 0.5$		<sup>14</sup> ARMSTRONG	83B	OMEG	18.5	$K^- p \rightarrow K^- K^+ \Lambda$
1019.7	$\pm 0.3$		<sup>14</sup> BARATE	83	GOLI	190	$\pi^- Be \rightarrow 2\mu X$
1019.8	$\pm 0.2$	$\pm 0.5$	766	IVANOV	81	OLYA	$1-1.4 e^+ e^- \rightarrow K^+ K^-$
1019.4	$\pm 0.5$		337	COOPER	78B	HBC	$0.7-0.8 \bar{p}p \rightarrow K_S^0 K_L^0 \pi^+ \pi^-$
1020	$\pm 1$		383	<sup>14</sup> BALDI	77	CNTR	$10 \pi^- p \rightarrow \pi^- \phi p$
1018.9	$\pm 0.6$		800	COHEN	77	ASPK	$6 \pi^\pm N \rightarrow K^+ K^- N$
1019.7	$\pm 0.5$		454	KALBFLEISCH	76	HBC	$2.18 K^- p \rightarrow \Lambda K\bar{K}$
1019.4	$\pm 0.8$		984	BESCH	74	CNTR	$2 \gamma p \rightarrow p K^+ K^-$
1020.3	$\pm 0.4$		100	BALLAM	73	HBC	$2.8-9.3 \gamma p$
1019.4	$\pm 0.7$			BINNIE	73B	CNTR	$\pi^- p \rightarrow \phi n$
1019.6	$\pm 0.5$		120	<sup>16</sup> AGUILAR-...	72B	HBC	$3.9, 4.6 K^- p \rightarrow \Lambda K^+ K^-$
1019.9	$\pm 0.5$		100	<sup>16</sup> AGUILAR-...	72B	HBC	$3.9, 4.6 K^- p \rightarrow K^- p K^+ K^-$
1020.4	$\pm 0.5$		131	COLLEY	72	HBC	$10 K^+ p \rightarrow K^+ p \phi$
1019.9	$\pm 0.3$		410	STOTTLE...	71	HBC	$2.9 K^- p \rightarrow \Sigma/\Lambda K\bar{K}$

<sup>1</sup> Average of KOZYREV 16 and KOZYREV 18 values taking into account the correlated uncertainties. Supersedes individual KOZYREV 16 and KOZYREV 18 results.

<sup>2</sup> Using a vector meson dominance model with contribution from  $\phi(1020)$  and higher mass excitations of  $\rho(770)$ ,  $\omega(782)$ , and  $\phi(1020)$ .

<sup>3</sup> Using a phenomenological model based on KUHN 90 with a sum of Breit-Wigner resonances for  $\rho(770)$ ,  $\omega(782)$ ,  $\phi(1020)$  and their higher mass excitations.

<sup>4</sup> Update of AKHMETSHIN 99D

<sup>5</sup> From the combined fit assuming that the total  $\phi(1020)$  production cross section is saturated by those of  $K^+ K^-$ ,  $K_S K_L$ ,  $\pi^+ \pi^- \pi^0$ , and  $\eta \gamma$  decays modes and using ACHASOV 00B for the  $\eta \gamma$  decay mode.

<sup>6</sup> Using a total width of  $4.43 \pm 0.05$  MeV. Systematic uncertainty included.

<sup>7</sup> Using a total width of  $4.43 \pm 0.05$  MeV.

<sup>8</sup> PELLINEN 82 review includes AKERLOF 77, DAUM 81, BALDI 77, AYRES 74, DEGROOT 74.

<sup>9</sup> Width fixed at 4.2 MeV.

<sup>10</sup> The values were extracted from a dispersively improved Breit-Wigner parameterization and do not include vacuum polarization. Inclusion of vacuum polarization gives  $1019.457 \pm 0.020$  MeV.

<sup>11</sup> From the  $D^\pm \rightarrow K^\pm K^+ K^-$  Dalitz plot fit with the Triple-M amplitude in the multi-meson model of AOUDE 18.

<sup>12</sup> The values were extracted from a dispersively improved Breit-Wigner parameterization and do not include vacuum polarization.

<sup>13</sup> Strongly correlated with AKHMETSHIN 04.

<sup>14</sup> Systematic errors not evaluated.

<sup>15</sup> Weighted and scaled average of 12 measurements of DIJKSTRA 86.

<sup>16</sup> Mass errors enlarged by us to  $\Gamma/\sqrt{N}$ ; see the note with the  $K^*(892)$  mass.

## $\phi(1020)$ WIDTH

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>4.249±0.013 OUR AVERAGE</b>				Error includes scale factor of 1.1.
4.245±0.013	2.3M	<sup>1</sup> KOZYREV	18	CMD3 $e^+ e^- \rightarrow K^+ K^-$ , $K_S^0 K_L^0$
4.205±0.103±0.067	28k	<sup>2</sup> LEES	14H	BABR $e^+ e^- \rightarrow K_S^0 K_L^0 \gamma$
4.29 ± 0.04 ± 0.07		<sup>3</sup> LEES	13Q	BABR $e^+ e^- \rightarrow K^+ K^- \gamma$
4.30 ± 0.06 ± 0.17	105k	AKHMETSHIN 06	CMD2	0.98–1.06 $e^+ e^- \rightarrow \pi^+ \pi^- \pi^0$
4.280±0.033±0.025	272k	<sup>4</sup> AKHMETSHIN 04	CMD2	$e^+ e^- \rightarrow K_L^0 K_S^0$
4.21 ± 0.04	1900k	<sup>5</sup> ACHASOV	01E	SND $e^+ e^- \rightarrow K^+ K^-$ , $K_S K_L, \pi^+ \pi^- \pi^0$
4.44 ± 0.09	55600	AKHMETSHIN 95	CMD2	$e^+ e^- \rightarrow$ hadrons
4.5 ± 0.7	1500	ARENTON	82	AEMS 11.8 polar. $p p \rightarrow K K$
4.2 ± 0.6	766	<sup>6</sup> IVANOV	81	OLYA 1–1.4 $e^+ e^- \rightarrow K^+ K^-$
4.3 ± 0.6		<sup>6</sup> CORDIER	80	DM1 $e^+ e^- \rightarrow \pi^+ \pi^- \pi^0$
4.36 ± 0.29	3681	<sup>6</sup> BUKIN	78C	OLYA $e^+ e^- \rightarrow$ hadrons
4.4 ± 0.6	984	<sup>6</sup> BESCH	74	CNTR $2 \gamma p \rightarrow p K^+ K^-$
4.67 ± 0.72	681	<sup>6</sup> BALAKIN	71	OSPK $e^+ e^- \rightarrow$ hadrons
4.09 ± 0.29		BIZOT	70	OSPK $e^+ e^- \rightarrow$ hadrons
• • • We do not use the following data for averages, fits, limits, etc. • • •				
4.07 ± 0.13 ± 0.01		<sup>7</sup> HOID	20	RVUE $e^+ e^- \rightarrow \pi^0 \gamma$
4.23 ± 0.04 ± 0.02		<sup>8</sup> HOFERICHT...	19	RVUE $e^+ e^- \rightarrow \pi^+ \pi^- \pi^0$
4.249±0.015	1.7M	KOZYREV	18	CMD3 $e^+ e^- \rightarrow K^+ K^-$
4.240±0.017	610k	KOZYREV	16	CMD3 $e^+ e^- \rightarrow K_S^0 K_L^0$
4.37 ± 0.02		LEES	13F	BABR $D^+ \rightarrow K^+ K^- \pi^+$
4.24 ± 0.02 ± 0.03	542k	<sup>9</sup> AKHMETSHIN 08	CMD2	1.02 $e^+ e^- \rightarrow K^+ K^-$
4.28 ± 0.13	12540	<sup>10</sup> AUBERT,B	05J	BABR $D^0 \rightarrow \bar{K}^0 K^+ K^-$
4.45 ± 0.06	271k	DIJKSTRA	86	SPEC 100 $\pi^-$ Be
3.6 ± 0.8	337	<sup>6</sup> COOPER	78B	HBC 0.7–0.8 $\bar{p} p \rightarrow$ $K_S^0 K_L^0 \pi^+ \pi^-$
4.5 ± 0.50	1300	<sup>6,10</sup> AKERLOF	77	SPEC 400 $p\bar{A} \rightarrow K^+ K^- X$
4.5 ± 0.8	500	<sup>6,10</sup> AYRES	74	ASPK 3–6 $\pi^- p \rightarrow$ $K^+ K^- n, K^- p \rightarrow$ $K^+ K^- \Lambda/\Sigma^0$
3.81 ± 0.37		COSME	74B	OSPK $e^+ e^- \rightarrow K_L^0 K_S^0$
3.8 ± 0.7	454	<sup>6</sup> BORENSTEIN	72	HBC 2.18 $K^- p \rightarrow K \bar{K} n$

<sup>1</sup> Average of KOZYREV 16 and KOZYREV 18 values taking into account the correlated uncertainties. Supersedes individual KOZYREV 16 and KOZYREV 18 results.

<sup>2</sup> Using a vector meson dominance model with contribution from  $\phi(1020)$  and higher mass excitations of  $\rho(770)$ ,  $\omega(782)$ , and  $\phi(1020)$ .

<sup>3</sup> Using a phenomenological model based on KUHN 90 with a sum of Breit-Wigner resonances for  $\rho(770)$ ,  $\omega(782)$ ,  $\phi(1020)$  and their higher mass excitations.

<sup>4</sup> Update of AKHMETSHIN 99D

<sup>5</sup> From the combined fit assuming that the total  $\phi(1020)$  production cross section is saturated by those of  $K^+ K^-$ ,  $K_S K_L$ ,  $\pi^+ \pi^- \pi^0$ , and  $\eta \gamma$  decays modes and using ACHASOV 00B for the  $\eta \gamma$  decay mode.

<sup>6</sup> Width errors enlarged by us to  $4\Gamma/\sqrt{N}$ ; see the note with the  $K^*(892)$  mass.

<sup>7</sup> The values were extracted from a dispersively improved Breit-Wigner parameterization.<sup>8</sup> The values were extracted from a dispersively improved Breit-Wigner parameterization and do not include vacuum polarization.<sup>9</sup> Strongly correlated with AKHMETSHIN 04.<sup>10</sup> Systematic errors not evaluated.

## **$\phi(1020)$ DECAY MODES**

Mode	Fraction ( $\Gamma_i/\Gamma$ )	Scale factor/ Confidence level
$\Gamma_1 K^+ K^-$	(49.1 $\pm$ 0.5 ) %	S=1.3
$\Gamma_2 K_L^0 K_S^0$	(33.9 $\pm$ 0.4 ) %	S=1.2
$\Gamma_3 \rho\pi + \pi^+\pi^-\pi^0$	(15.4 $\pm$ 0.4 ) %	S=1.2
$\Gamma_4 \pi^+\pi^-\pi^0$		
$\Gamma_5 \eta\gamma$	( 1.301 $\pm$ 0.024 ) %	S=1.2
$\Gamma_6 \pi^0\gamma$	( 1.32 $\pm$ 0.05 ) $\times 10^{-3}$	
$\Gamma_7 \ell^+\ell^-$	—	
$\Gamma_8 e^+e^-$	( 2.979 $\pm$ 0.033 ) $\times 10^{-4}$	S=1.2
$\Gamma_9 \mu^+\mu^-$	( 2.85 $\pm$ 0.22 ) $\times 10^{-4}$	S=1.2
$\Gamma_{10} \eta e^+e^-$	( 1.08 $\pm$ 0.04 ) $\times 10^{-4}$	
$\Gamma_{11} \pi^+\pi^-$	( 7.3 $\pm$ 1.3 ) $\times 10^{-5}$	
$\Gamma_{12} \omega\pi^0$	( 4.7 $\pm$ 0.5 ) $\times 10^{-5}$	
$\Gamma_{13} \omega\gamma$	< 5 %	CL=84%
$\Gamma_{14} \rho\gamma$	< 1.2 $\times 10^{-5}$	CL=90%
$\Gamma_{15} \pi^+\pi^-\gamma$	( 4.1 $\pm$ 1.3 ) $\times 10^{-5}$	
$\Gamma_{16} f_0(980)\gamma$	( 3.22 $\pm$ 0.19 ) $\times 10^{-4}$	S=1.1
$\Gamma_{17} \pi^0\pi^0\gamma$	( 1.12 $\pm$ 0.06 ) $\times 10^{-4}$	
$\Gamma_{18} \pi^+\pi^-\pi^+\pi^-$	( 3.9 $\pm$ 2.8 ) $\times 10^{-6}$	
$\Gamma_{19} \pi^+\pi^+\pi^-\pi^-\pi^0$	< 4.6 $\times 10^{-6}$	CL=90%
$\Gamma_{20} \pi^0e^+e^-$	( 1.33 $\pm$ 0.07 ) $\times 10^{-5}$	
$\Gamma_{21} \pi^0\eta\gamma$	( 7.27 $\pm$ 0.30 ) $\times 10^{-5}$	S=1.5
$\Gamma_{22} a_0(980)\gamma$	( 7.6 $\pm$ 0.6 ) $\times 10^{-5}$	
$\Gamma_{23} K^0\bar{K}^0\gamma$	< 1.9 $\times 10^{-8}$	CL=90%
$\Gamma_{24} \eta'(958)\gamma$	( 6.21 $\pm$ 0.20 ) $\times 10^{-5}$	
$\Gamma_{25} \eta\pi^0\pi^0\gamma$	< 2 $\times 10^{-5}$	CL=90%
$\Gamma_{26} \mu^+\mu^-\gamma$	( 1.4 $\pm$ 0.5 ) $\times 10^{-5}$	
$\Gamma_{27} \rho\gamma\gamma$	< 1.2 $\times 10^{-4}$	CL=90%
$\Gamma_{28} \eta\pi^+\pi^-$	< 1.8 $\times 10^{-5}$	CL=90%
$\Gamma_{29} \eta\mu^+\mu^-$	< 9.4 $\times 10^{-6}$	CL=90%
$\Gamma_{30} \eta U \rightarrow \eta e^+e^-$	< 1 $\times 10^{-6}$	CL=90%
$\Gamma_{31}$ invisible	< 1.7 $\times 10^{-4}$	CL=90%
<b>Lepton Family number (<i>LF</i>) violating modes</b>		
$\Gamma_{32} e^\pm\mu^\mp$	<i>LF</i> < 2 $\times 10^{-6}$	CL=90%

## CONSTRAINED FIT INFORMATION

An overall fit to 30 branching ratios uses 80 measurements and one constraint to determine 14 parameters. The overall fit has a  $\chi^2 = 61.8$  for 67 degrees of freedom.

The following *off-diagonal* array elements are the correlation coefficients  $\langle \delta x_i \delta x_j \rangle / (\delta x_i \cdot \delta x_j)$ , in percent, from the fit to the branching fractions,  $x_i \equiv \Gamma_i / \Gamma_{\text{total}}$ . The fit constrains the  $x_i$  whose labels appear in this array to sum to one.

$x_2$	-72									
$x_3$	-61 -12									
$x_5$	-19 17 2									
$x_6$	-11 11 1 8									
$x_8$	47 -50 -8 -35 -22									
$x_9$	-5 4 1 3 2 -9									
$x_{11}$	-3 3 0 2 1 -6 1									
$x_{12}$	-4 4 1 3 2 -8 1 0									
$x_{16}$	0 0 0 0 0 0 0 0 0									
$x_{17}$	-9 9 1 18 4 -18 2 1 1 0									
$x_{18}$	-1 1 0 1 0 -2 0 0 0 0									
$x_{22}$	0 0 0 0 0 0 0 0 0 0									
$x_{24}$	-6 5 1 32 2 -11 1 1 1 0									
	$x_1$	$x_2$	$x_3$	$x_5$	$x_6$	$x_8$	$x_9$	$x_{11}$	$x_{12}$	$x_{16}$
$x_{18}$	0									
$x_{22}$	0 0									
$x_{24}$	6 0 0									
	$x_{17}$	$x_{18}$	$x_{22}$							

## $\phi(1020)$ PARTIAL WIDTHS

### $\Gamma(\eta\gamma)$

VALUE (keV)	DOCUMENT ID	TECN	COMMENT	$\Gamma_5$
<b>• • •</b> We do not use the following data for averages, fits, limits, etc. <b>• • •</b>				
$58.9 \pm 0.5 \pm 2.4$	ACHASOV	00	SND	$e^+ e^- \rightarrow \eta\gamma$

### $\Gamma(\pi^0\gamma)$

VALUE (keV)	DOCUMENT ID	TECN	COMMENT	$\Gamma_6$
<b>• • •</b> We do not use the following data for averages, fits, limits, etc. <b>• • •</b>				
$5.40 \pm 0.16^{+0.43}_{-0.40}$	ACHASOV	00	SND	$e^+ e^- \rightarrow \pi^0\gamma$

$\Gamma(\ell^+\ell^-)$	$\Gamma_7$		
VALUE (keV)	DOCUMENT ID	TECN	COMMENT
$\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$			
$1.320 \pm 0.017 \pm 0.015$	<sup>1</sup> AMBROSINO 05	KLOE	$1.02 e^+ e^- \rightarrow \mu^+ \mu^-$
<sup>1</sup> Weighted average of $\Gamma_{ee}$ and $\sqrt{\Gamma_{ee}\Gamma_{\mu\mu}}$ from AMBROSINO 05 assuming lepton universality.			

$\Gamma(e^+e^-)$	$\Gamma_8$		
VALUE (keV)	DOCUMENT ID	TECN	COMMENT
<b>1.27 ± 0.04 OUR EVALUATION</b>			
<b>1.251 ± 0.021 OUR AVERAGE</b> Error includes scale factor of 1.1.			
$1.235 \pm 0.006 \pm 0.022$	<sup>1</sup> AKHMETSHIN 11	CMD2	$1.02 e^+ e^- \rightarrow \phi$
$1.32 \pm 0.05 \pm 0.03$	<sup>2</sup> AMBROSINO 05	KLOE	$1.02 e^+ e^- \rightarrow e^+ e^-$
$1.28 \pm 0.05$	AKHMETSHIN 95	CMD2	$1.02 e^+ e^- \rightarrow \phi$
<sup>1</sup> Combined analysis of the CMD-2 data on $\phi \rightarrow K^+ K^-$ , $K_S^0 K_L^0$ , $\pi^+ \pi^- \pi^0$ , $\eta \gamma$ assuming that the sum of their branching fractions is $0.99741 \pm 0.00007$ .			
<sup>2</sup> From forward-backward asymmetry and using $\Gamma_{\text{total}} = 4.26 \pm 0.05$ MeV from the 2004 edition of this Review.			

$(\Gamma(e^+e^-) \times \Gamma(\mu^+\mu^-))^{1/2}$	$(\Gamma_8 \Gamma_9)^{1/2}$		
VALUE (keV)	DOCUMENT ID	TECN	COMMENT
$1.320 \pm 0.018 \pm 0.017$	AMBROSINO 05	KLOE	$1.02 e^+ e^- \rightarrow \mu^+ \mu^-$

### $\phi(1020) \Gamma(i) \Gamma(e^+e^-)/\Gamma(\text{total})$

$\Gamma(K^+K^-) \times \Gamma(e^+e^-)/\Gamma_{\text{total}}$	$\Gamma_1 \Gamma_8/\Gamma$			
VALUE (keV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.6340 ± 0.0070 ± 0.0039</b>				
$1$	LEES	13Q	BABR	$e^+ e^- \rightarrow K^+ K^- \gamma$
$\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$				
$0.669 \pm 0.001 \pm 0.023$	1.7M	KOZYREV	18	CMD3 $e^+ e^- \rightarrow K^+ K^-$
<sup>1</sup> Using a phenomenological model based on KUHN 90 with a sum of Breit-Wigner resonances for $\rho(770)$ , $\omega(782)$ , $\phi(1020)$ and their higher mass excitations. The first error combines statistical and systematic uncertainties. The second one is due to the parametrization of the charged kaon form factor and mass calibration.				

$\Gamma(K_L^0 K_S^0) \times \Gamma(e^+e^-)/\Gamma_{\text{total}}$	$\Gamma_2 \Gamma_8/\Gamma$			
VALUE (keV)	EVTS	DOCUMENT ID	TECN	COMMENT
$0.4200 \pm 0.0033 \pm 0.0123$	28k	<sup>1</sup> LEES	14H	BABR $e^+ e^- \rightarrow K_S^0 K_L^0 \gamma$

<sup>1</sup> Using a vector meson dominance model with contribution from  $\phi(1020)$  and higher mass excitations of  $\rho(770)$ ,  $\omega(782)$ , and  $\phi(1020)$ .

$[\Gamma(\rho\pi) + \Gamma(\pi^+\pi^-\pi^0)] \times \Gamma(e^+e^-)/\Gamma_{\text{total}}$	$\Gamma_3 \Gamma_8/\Gamma$		
VALUE (eV)	DOCUMENT ID	TECN	COMMENT
$184.1 \pm 2.1 \pm 8.0$	1 LEES	21B	BABR $10.5 e^+ e^- \rightarrow \pi^+ \pi^- \pi^0 \gamma$

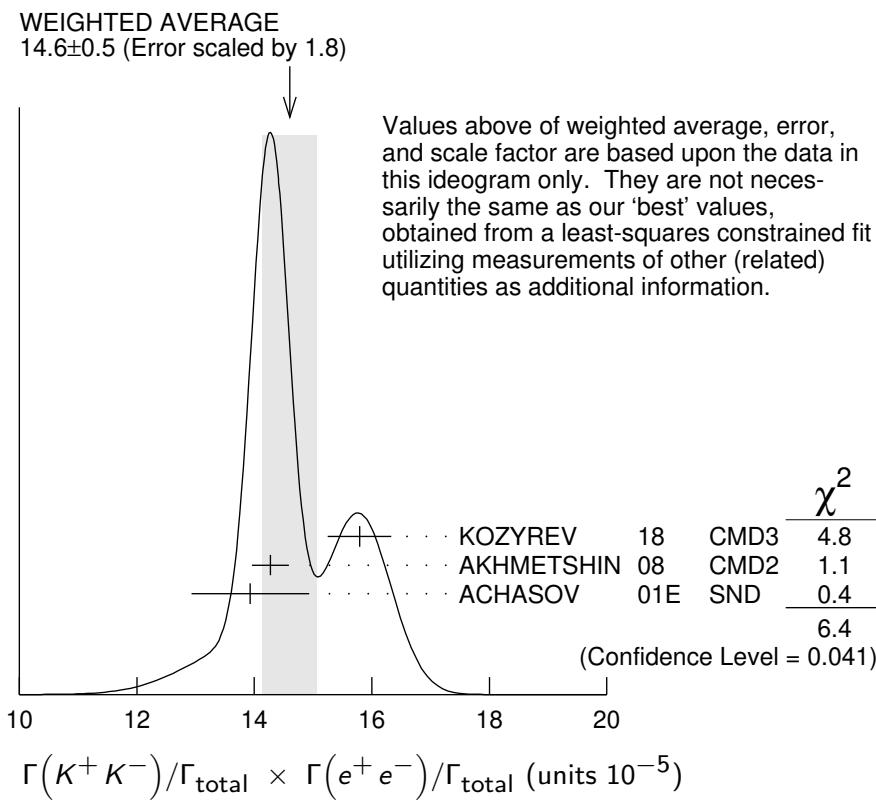
<sup>1</sup> From the cross section for  $e^+ e^- \rightarrow \pi^+ \pi^- \pi^0$  with contributions from  $\rho(770)$ ,  $\omega(782)$ ,  $\phi(1020)$ ,  $\omega(1420)$ , and  $\omega(1650)$ .

$$\phi(1020) \Gamma(i)\Gamma(e^+e^-)/\Gamma^2(\text{total})$$

$$\Gamma(K^+K^-)/\Gamma_{\text{total}} \times \Gamma(e^+e^-)/\Gamma_{\text{total}} \quad \Gamma_1/\Gamma \times \Gamma_8/\Gamma$$

VALUE (units $10^{-5}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>14.64 <math>\pm 0.27</math> OUR FIT</b>		Error includes scale factor of 1.4.		
<b>14.6 <math>\pm 0.5</math> OUR AVERAGE</b>		Error includes scale factor of 1.8. See the ideogram below.		
15.789 $\pm 0.541$	1.7M	KOZYREV 18	CMD3	$e^+e^- \rightarrow K^+K^-$
14.27 $\pm 0.05$ $\pm 0.31$	542k	AKHMETSHIN 08	CMD2	$1.02 e^+e^- \rightarrow K^+K^-$
13.93 $\pm 0.14$ $\pm 0.99$	1000k	<sup>1</sup> ACHASOV 01E	SND	$e^+e^- \rightarrow K^+K^-$ , $K_S K_L, \pi^+\pi^-\pi^0$

<sup>1</sup> From the combined fit assuming that the total  $\phi(1020)$  production cross section is saturated by those of  $K^+K^-$ ,  $K_S K_L$ ,  $\pi^+\pi^-\pi^0$ , and  $\eta\gamma$  decays modes and using ACHASOV 00B for the  $\eta\gamma$  decay mode.



$$\Gamma(K_L^0 K_S^0)/\Gamma_{\text{total}} \times \Gamma(e^+e^-)/\Gamma_{\text{total}}$$

$$\Gamma_2/\Gamma \times \Gamma_8/\Gamma$$

VALUE (units $10^{-5}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>10.11 <math>\pm 0.12</math> OUR FIT</b>				
<b>10.07 <math>\pm 0.13</math> OUR AVERAGE</b>				
10.078 $\pm 0.223$	610k	<sup>1</sup> KOZYREV 16	CMD3	$e^+e^- \rightarrow K_S^0 K_L^0$
10.01 $\pm 0.04$ $\pm 0.17$	272k	<sup>2</sup> AKHMETSHIN 04	CMD2	$e^+e^- \rightarrow K_L^0 K_S^0$
10.27 $\pm 0.07$ $\pm 0.34$	500k	<sup>3</sup> ACHASOV 01E	SND	$e^+e^- \rightarrow K^+K^-$ , $K_S K_L, \pi^+\pi^-\pi^0$

<sup>1</sup> KOZYREV 16 also reports  $\Gamma(e^+e^-) B(\phi \rightarrow K_S^0 K_L^0) = (0.428 \pm 0.001 \pm 0.009)$  keV.

<sup>2</sup> Update of AKHMETSHIN 99D

<sup>3</sup> From the combined fit assuming that the total  $\phi(1020)$  production cross section is saturated by those of  $K^+K^-$ ,  $K_S K_L$ ,  $\pi^+\pi^-\pi^0$ , and  $\eta\gamma$  decays modes and using ACHASOV 00B for the  $\eta\gamma$  decay mode.

$[\Gamma(\rho\pi) + \Gamma(\pi^+\pi^-\pi^0)]/\Gamma_{\text{total}} \times \Gamma(e^+e^-)/\Gamma_{\text{total}}$	$\Gamma_3/\Gamma \times \Gamma_8/\Gamma$			
VALUE (units $10^{-5}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>4.58 ±0.11 OUR FIT</b>		Error includes scale factor of 1.1.		
<b>4.51 ±0.14 OUR AVERAGE</b>				
4.51 ±0.16 ±0.11	105k	AKHMETSHIN 06	CMD2	$0.98-1.06 e^+e^- \rightarrow \pi^+\pi^-\pi^0$
4.665±0.042±0.261	400k	<sup>1</sup> ACHASOV 01E	SND	$e^+e^- \rightarrow K^+K^-$ , $K_S K_L$ , $\pi^+\pi^-\pi^0$
4.35 ±0.27 ±0.08	11169	<sup>2</sup> AKHMETSHIN 98	CMD2	$e^+e^- \rightarrow \pi^+\pi^-\pi^0$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
4.38 ±0.12		BENAYOUN 10	RVUE	$0.4-1.05 e^+e^-$
4.30 ±0.08 ±0.21		<sup>3</sup> AUBERT,B 04N	BABR	$10.6 e^+e^- \rightarrow \pi^+\pi^-\pi^0\gamma$

<sup>1</sup> From the combined fit assuming that the total  $\phi(1020)$  production cross section is saturated by those of  $K^+K^-$ ,  $K_S K_L$ ,  $\pi^+\pi^-\pi^0$ , and  $\eta\gamma$  decays modes and using ACHASOV 00B for the  $\eta\gamma$  decay mode.

<sup>2</sup> Recalculated by us from the cross section in the peak.

<sup>3</sup> Superseeded by LEES 21B.

$\Gamma(\eta\gamma)/\Gamma_{\text{total}} \times \Gamma(e^+e^-)/\Gamma_{\text{total}}$	$\Gamma_5/\Gamma \times \Gamma_8/\Gamma$			
VALUE (units $10^{-6}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>3.88 ±0.07 OUR FIT</b>		Error includes scale factor of 1.2.		
<b>3.93 ±0.09 OUR AVERAGE</b>		Error includes scale factor of 1.3. See the ideogram below.		
4.050±0.067±0.118	33k	<sup>1</sup> ACHASOV 07B	SND	$0.6-1.38 e^+e^- \rightarrow \eta\gamma$
4.093 <sup>+0.040</sup> <sub>-0.043</sub> ±0.247	17.4k	<sup>2</sup> AKHMETSHIN 05	CMD2	$0.60-1.38 e^+e^- \rightarrow \eta\gamma$
3.850±0.041±0.159	23k	<sup>3,4</sup> AKHMETSHIN 01B	CMD2	$e^+e^- \rightarrow \eta\gamma$
4.00 ±0.04 ±0.11		<sup>5</sup> ACHASOV 00	SND	$e^+e^- \rightarrow \eta\gamma$
3.53 ±0.08 ±0.17	2200	<sup>6,7</sup> AKHMETSHIN 99F	CMD2	$e^+e^- \rightarrow \eta\gamma$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
4.19 ±0.06		<sup>8</sup> BENAYOUN 10	RVUE	$0.4-1.05 e^+e^-$

<sup>1</sup> From a combined fit of  $\sigma(e^+e^- \rightarrow \eta\gamma)$  with  $\eta \rightarrow 3\pi^0$  and  $\eta \rightarrow \pi^+\pi^-\pi^0$ , and fixing  $B(\eta \rightarrow 3\pi^0) / B(\eta \rightarrow \pi^+\pi^-\pi^0) = 1.44 \pm 0.04$ . Recalculated by us from the cross section at the peak. Supersedes ACHASOV 00D and ACHASOV 06A.

<sup>2</sup> From the  $\eta \rightarrow 2\gamma$  decay and using  $B(\eta \rightarrow \gamma\gamma) = 39.43 \pm 0.26\%$ .

<sup>3</sup> From the  $\eta \rightarrow 3\pi^0$  decay and using  $B(\eta \rightarrow 3\pi^0) = (32.24 \pm 0.29) \times 10^{-2}$ .

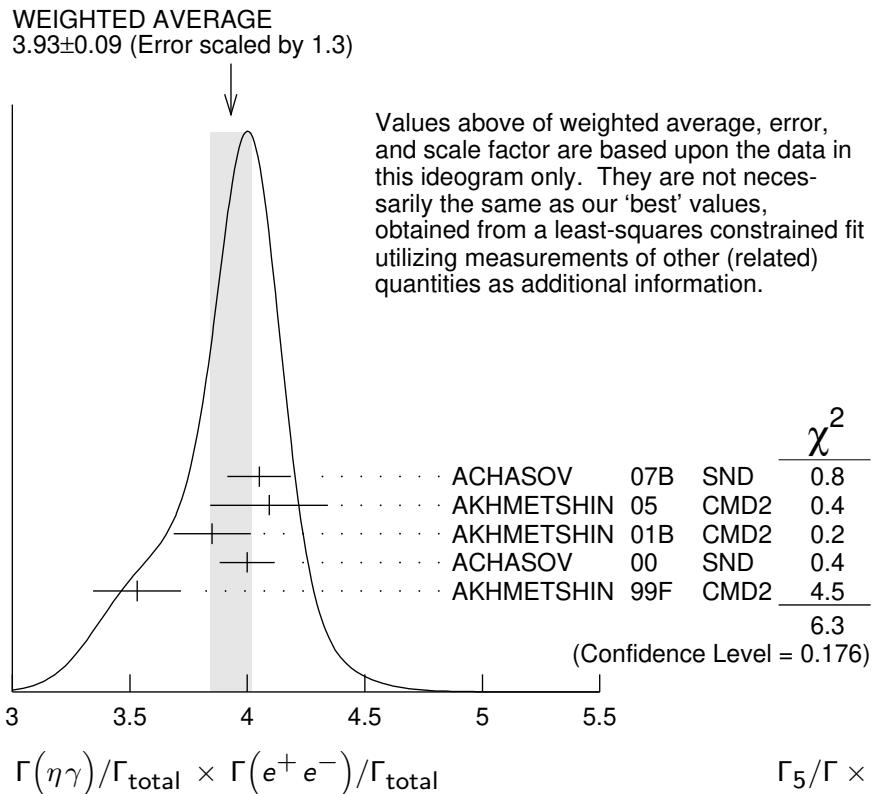
<sup>4</sup> The combined fit from 600 to 1380 MeV taking into account  $\rho(770)$ ,  $\omega(782)$ ,  $\phi(1020)$ , and  $\rho(1450)$  (mass and width fixed at 1450 MeV and 310 MeV respectively).

<sup>5</sup> From the  $\eta \rightarrow 2\gamma$  decay and using  $B(\eta \rightarrow 2\gamma) = (39.21 \pm 0.34) \times 10^{-2}$ .

<sup>6</sup> Recalculated by the authors from the cross section in the peak.

<sup>7</sup> From the  $\eta \rightarrow \pi^+\pi^-\pi^0$  decay and using  $B(\eta \rightarrow \pi^+\pi^-\pi^0) = (23.1 \pm 0.5) \times 10^{-2}$ .

<sup>8</sup> A simultaneous fit of  $e^+e^- \rightarrow \pi^+\pi^-$ ,  $\pi^+\pi^-\pi^0$ ,  $\pi^0\gamma$ ,  $\eta\gamma$  data.

 $\Gamma(\pi^0\gamma)/\Gamma_{\text{total}} \times \Gamma(e^+e^-)/\Gamma_{\text{total}}$  $\Gamma_6/\Gamma \times \Gamma_8/\Gamma$ 

VALUE (units $10^{-7}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>3.94±0.16 OUR FIT</b>				

**3.95±0.17 OUR AVERAGE**

4.04±0.09±0.19	1	ACHASOV 16A SND	0.60–1.38 $e^+e^- \rightarrow \pi^0\gamma$
3.75±0.11±0.29	18k	AKHMETSHIN 05 CMD2	0.60–1.38 $e^+e^- \rightarrow \pi^0\gamma$

• • • We do not use the following data for averages, fits, limits, etc. • • •

4.29±0.11	2	BENAYOUN 10 RVUE	0.4–1.05 $e^+e^-$
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3.67±0.10 <sup>+0.27</sup> <sub>-0.25</sub>	3	ACHASOV 00 SND	$e^+e^- \rightarrow \pi^0\gamma$
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<sup>1</sup> From the VMD model with the interfering  $\rho(770)$ ,  $\omega(782)$ ,  $\phi(1020)$  resonances, and an additional resonance describing the total contribution of the  $\rho(1450)$  and  $\omega(1420)$  states. Supersedes ACHASOV 00.

<sup>2</sup> A simultaneous fit of  $e^+e^- \rightarrow \pi^+\pi^-$ ,  $\pi^+\pi^-\pi^0$ ,  $\pi^0\gamma$ ,  $\eta\gamma$  data.

<sup>3</sup> From the  $\pi^0 \rightarrow 2\gamma$  decay and using  $B(\pi^0 \rightarrow 2\gamma) = (98.798 \pm 0.032) \times 10^{-2}$ .

 $\Gamma(\mu^+\mu^-)/\Gamma_{\text{total}} \times \Gamma(e^+e^-)/\Gamma_{\text{total}}$  $\Gamma_9/\Gamma \times \Gamma_8/\Gamma$ 

VALUE (units $10^{-8}$ )	DOCUMENT ID	TECN	COMMENT
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**8.5 ±0.6 OUR FIT** Error includes scale factor of 1.2.**8.8 ±0.9 OUR AVERAGE** Error includes scale factor of 1.5. See the ideogram below.

8.36±0.59±0.37	ACHASOV 01G SND	$e^+e^- \rightarrow \mu^+\mu^-$
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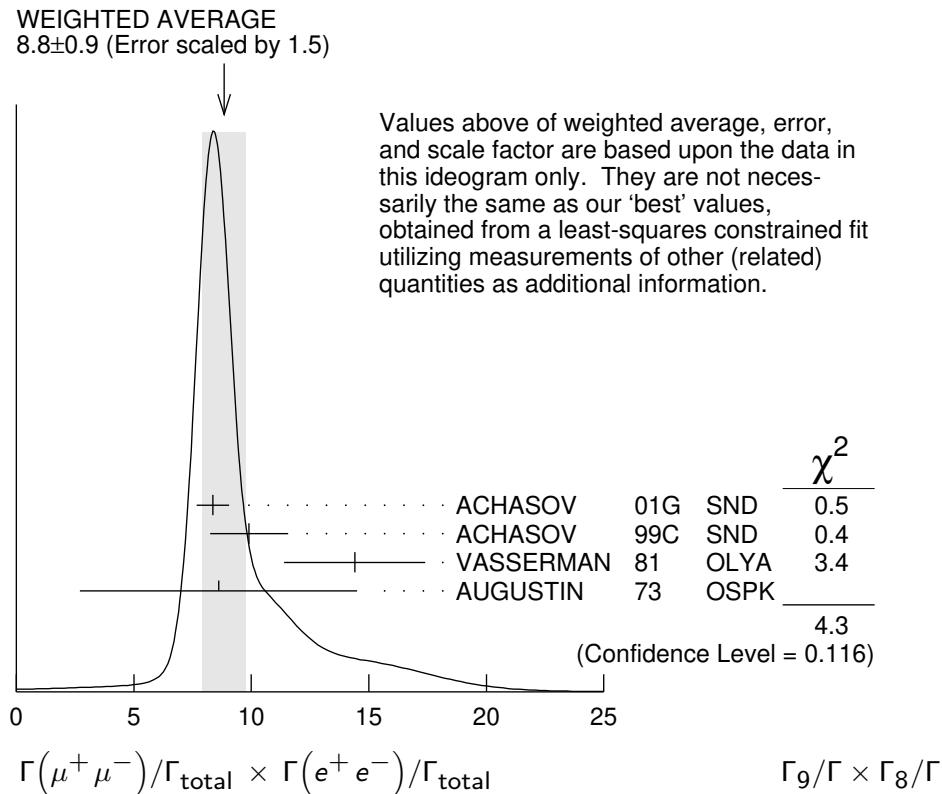
9.9 ±1.4 ±0.9	1 ACHASOV 99C SND	$e^+e^- \rightarrow \mu^+\mu^-$
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14.4 ±3.0	2 VASSERMAN 81 OLYA	$e^+e^- \rightarrow \mu^+\mu^-$
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8.6 ±5.9	2 AUGUSTIN 73 OSPK	$e^+e^- \rightarrow \mu^+\mu^-$
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<sup>1</sup> Recalculated by the authors from the cross section in the peak.

<sup>2</sup> Recalculated by us from the cross section in the peak.



### $\Gamma(\pi^+ \pi^-)/\Gamma_{\text{total}} \times \Gamma(e^+ e^-)/\Gamma_{\text{total}}$

VALUE (units  $10^{-8}$ )

**2.2 ±0.4 OUR FIT**

**2.2 ±0.4 OUR AVERAGE**

$2.1 \pm 0.3 \pm 0.3$

$1.95^{+1.15}_{-0.87}$

$6.01^{+3.19}_{-2.51}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$3.31 \pm 0.99$

DOCUMENT ID

TECN

COMMENT

<sup>1</sup> ACHASOV 00C SND  $e^+ e^- \rightarrow \pi^+ \pi^-$

<sup>2</sup> GOLUBEV 86 ND  $e^+ e^- \rightarrow \pi^+ \pi^-$

<sup>2</sup> VASSERMAN 81 OLYA  $e^+ e^- \rightarrow \pi^+ \pi^-$

### $\Gamma_{11}/\Gamma \times \Gamma_8/\Gamma$

### $\Gamma(\omega \pi^0)/\Gamma_{\text{total}} \times \Gamma(e^+ e^-)/\Gamma_{\text{total}}$

VALUE (units  $10^{-8}$ )

**1.40±0.15 OUR FIT**

**1.37±0.17±0.01**

DOCUMENT ID

TECN

COMMENT

### $\Gamma_{12}/\Gamma \times \Gamma_8/\Gamma$

<sup>1,2</sup> AMBROSINO 08G KLOE  $e^+ e^- \rightarrow \pi^+ \pi^- 2\pi^0, 2\pi^0 \gamma$

<sup>1</sup> Recalculated by the authors from the cross section at the peak.

<sup>2</sup> AMBROSINO 08G reports  $[\Gamma(\phi(1020) \rightarrow \omega \pi^0)/\Gamma_{\text{total}} \times \Gamma(\phi(1020) \rightarrow e^+ e^-)/\Gamma_{\text{total}}] \times [B(\omega(782) \rightarrow \pi^+ \pi^- \pi^0)] = (1.22 \pm 0.13 \pm 0.08) \times 10^{-8}$  which we divide

by our best value  $B(\omega(782) \rightarrow \pi^+ \pi^- \pi^0) = (89.2 \pm 0.7) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

$$\Gamma(\pi^0 \pi^0 \gamma)/\Gamma_{\text{total}} \times \Gamma(e^+ e^-)/\Gamma_{\text{total}} \quad \Gamma_{17}/\Gamma \times \Gamma_8/\Gamma$$

<u>VALUE (units <math>10^{-8}</math>)</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b><math>3.34 \pm 0.17</math> OUR FIT</b>			
<b><math>3.33^{+0.04}_{-0.09} {}^{+0.19}_{-0.20}</math></b>	<sup>1</sup> AMBROSINO 07	KLOE	$e^+ e^- \rightarrow \pi^0 \pi^0 \gamma$

<sup>1</sup> Calculated by the authors from the cross section at the peak.

$$\Gamma(\pi^+ \pi^- \pi^+ \pi^-)/\Gamma_{\text{total}} \times \Gamma(e^+ e^-)/\Gamma_{\text{total}} \quad \Gamma_{18}/\Gamma \times \Gamma_8/\Gamma$$

<u>VALUE (units <math>10^{-9}</math>)</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b><math>1.2^{+0.8}_{-0.7}</math> OUR FIT</b>				
<b><math>1.17 \pm 0.52 \pm 0.64</math></b>	3285	<sup>1</sup> AKHMETSHIN 00E	CMD2	$e^+ e^- \rightarrow \pi^+ \pi^- \pi^+ \pi^-$

<sup>1</sup> Recalculated by the authors from the cross section in the peak.

## $\phi(1020)$ BRANCHING RATIOS

$$\Gamma(K^+ K^-)/\Gamma_{\text{total}} \quad \Gamma_1/\Gamma$$

<u>VALUE</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b><math>0.491 \pm 0.005</math> OUR FIT</b>		Error includes scale factor of 1.3.		
<b><math>0.493 \pm 0.010</math> OUR AVERAGE</b>				
0.492 $\pm$ 0.012	2913	AKHMETSHIN 95	CMD2	$e^+ e^- \rightarrow K^+ K^-$
0.44 $\pm$ 0.05	321	KALBFLEISCH 76	HBC	$2.18 K^- p \rightarrow \Lambda K^+ K^-$
0.49 $\pm$ 0.06	270	DEGROOT 74	HBC	$4.2 K^- p \rightarrow \Lambda \phi$
0.540 $\pm$ 0.034	565	BALAKIN 71	OSPK	$e^+ e^- \rightarrow K^+ K^-$
0.48 $\pm$ 0.04	252	LINDSEY 66	HBC	$2.1-2.7 K^- p \rightarrow \Lambda K^+ K^-$
$\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$				
0.493 $\pm$ 0.003 $\pm$ 0.007		<sup>1</sup> AKHMETSHIN 11	CMD2	$1.02 e^+ e^- \rightarrow K^+ K^-$
0.476 $\pm$ 0.017	1000k	<sup>2</sup> ACHASOV 01E	SND	$e^+ e^- \rightarrow K^+ K^-, K_S K_L, \pi^+ \pi^- \pi^0$

<sup>1</sup> Combined analysis of the CMD-2 data on  $\phi \rightarrow K^+ K^-$ ,  $K_S^0 K_L^0$ ,  $\pi^+ \pi^- \pi^0$ ,  $\eta \gamma$  assuming that the sum of their branching fractions is  $0.99741 \pm 0.00007$ .

<sup>2</sup> Using  $B(\phi \rightarrow e^+ e^-) = (2.93 \pm 0.14) \times 10^{-4}$ .

$$\Gamma(K_L^0 K_S^0)/\Gamma_{\text{total}} \quad \Gamma_2/\Gamma$$

<u>VALUE</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b><math>0.339 \pm 0.004</math> OUR FIT</b>		Error includes scale factor of 1.2.		
<b><math>0.331 \pm 0.009</math> OUR AVERAGE</b>				
0.335 $\pm$ 0.010	40644	AKHMETSHIN 95	CMD2	$e^+ e^- \rightarrow K_L^0 K_S^0$
0.326 $\pm$ 0.035		DOLINSKY 91	ND	$e^+ e^- \rightarrow K_L^0 K_S^0$
0.310 $\pm$ 0.024		DRUZHININ 84	ND	$e^+ e^- \rightarrow K_L^0 K_S^0$
$\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$				
0.336 $\pm$ 0.002 $\pm$ 0.006		<sup>1</sup> AKHMETSHIN 11	CMD2	$1.02 e^+ e^- \rightarrow K_S^0 K_L^0$
0.351 $\pm$ 0.013	500k	<sup>2</sup> ACHASOV 01E	SND	$e^+ e^- \rightarrow K^+ K^-, K_S K_L, \pi^+ \pi^- \pi^0$
0.27 $\pm$ 0.03	133	KALBFLEISCH 76	HBC	$2.18 K^- p \rightarrow \Lambda K_L^0 K_S^0$

0.257±0.030	95	<sup>3</sup> BALAKIN	71	OSPK	$e^+ e^- \rightarrow K_L^0 K_S^0$
0.40 ± 0.04	167	LINDSEY	66	HBC	$2.1\text{--}2.7 K^- p \rightarrow \Lambda K_L^0 K_S^0$

<sup>1</sup> Combined analysis of the CMD-2 data on  $\phi \rightarrow K^+ K^-, K_S^0 K_L^0, \pi^+ \pi^- \pi^0, \eta \gamma$  assuming that the sum of their branching fractions is  $0.99741 \pm 0.00007$ .

<sup>2</sup> Using  $B(\phi \rightarrow e^+ e^-) = (2.93 \pm 0.14) \times 10^{-4}$ .

<sup>3</sup> Balakin error increased by Paul.

### $\Gamma(K_L^0 K_S^0)/\Gamma(K^+ K^-)$ $\Gamma_2/\Gamma_1$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.690±0.014 OUR FIT</b>	Error includes scale factor of 1.3.			
<b>0.740±0.031 OUR AVERAGE</b>				
0.70 ± 0.06	2732	BUKIN	78C OLYA	$e^+ e^- \rightarrow K_L^0 K_S^0$
0.82 ± 0.08		LOSTY	78 HBC	$4.2 K^- p \rightarrow \phi \text{hyperon}$
0.71 ± 0.05		LAVEN	77 HBC	$10 K^- p \rightarrow K^+ K^- \Lambda$
0.71 ± 0.08		LYONS	77 HBC	$3\text{--}4 K^- p \rightarrow \Lambda \phi$
0.89 ± 0.10	144	AGUILAR...	72B HBC	$3.9, 4.6 K^- p$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.638±0.022	2.3M	<sup>1</sup> KOZYREV	18	CMD3 $e^+ e^- \rightarrow K_L^0 K_S^0$ ,
0.68 ± 0.03		<sup>2</sup> AKHMETSHIN 95	CMD2	$e^+ e^- \rightarrow K_L^0 K_S^0, K^+ K^-$

<sup>1</sup> The prediction taking into account phase-space difference, radiative corrections, isospin breaking, and the Sommerfeld-Gamow-Sakharov factor gives 0.630.

<sup>2</sup> Theoretical analysis of BRAMON 00 taking into account phase-space difference, electromagnetic radiative corrections, as well as isospin breaking, predicts 0.62. FLOREZ-BAEZ 08 predicts 0.63 considering also structure-dependent radiative corrections. FIS-CHBACH 02 calculates additional corrections caused by the close threshold and predicts 0.68. See also BENAYOUN 01 and DUBYNSKIY 07. BENAYOUN 12 obtains  $0.71 \pm 0.01$  in the HLS model.

### $\Gamma(K_L^0 K_S^0)/\Gamma(K \bar{K})$ $\Gamma_2/(\Gamma_1 + \Gamma_2)$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.408±0.005 OUR FIT</b>	Error includes scale factor of 1.3.			
<b>0.45 ± 0.04 OUR AVERAGE</b>				
0.44 ± 0.07		<sup>1</sup> LONDON	66 HBC	$2.24 K^- p \rightarrow \Lambda K \bar{K}$
0.48 ± 0.07	52	BADIER	65B HBC	$3 K^- p$
0.40 ± 0.10	34	SCHLEIN	63 HBC	$1.95 K^- p \rightarrow \Lambda K \bar{K}$

<sup>1</sup> This is probably not affected by their controversial background subtraction; the value is from their numbers of  $K_1 K_2$  vs  $K^+ K^-$  events.

### $[\Gamma(\rho\pi) + \Gamma(\pi^+ \pi^- \pi^0)]/\Gamma_{\text{total}}$ $\Gamma_3/\Gamma$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.154±0.004 OUR FIT</b>	Error includes scale factor of 1.2.			
<b>0.151±0.009 OUR AVERAGE</b>	Error includes scale factor of 1.7.			
0.161±0.008	11761	AKHMETSHIN 95	CMD2	$e^+ e^- \rightarrow \pi^+ \pi^- \pi^0$
0.143±0.007		DOLINSKY	91 ND	$e^+ e^- \rightarrow \pi^+ \pi^- \pi^0$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.155±0.002±0.005		<sup>1</sup> AKHMETSHIN 11	CMD2	$1.02 e^+ e^- \rightarrow \pi^+ \pi^- \pi^0$
0.159±0.008	400k	<sup>2</sup> ACHASOV	01E SND	$e^+ e^- \rightarrow K^+ K^-, K_S K_L, \pi^+ \pi^- \pi^0$



• • • We do not use the following data for averages, fits, limits, etc. • • •

1.38 $\pm 0.02$ $\pm 0.02$	5 AKHMETSHIN 11	CMD2	1.02 $e^+ e^- \rightarrow \eta\gamma$
1.36 $\pm 0.05$ $\pm 0.01$	33k 6 ACHASOV 07B	SND	0.6–1.38 $e^+ e^- \rightarrow \eta\gamma$
1.373 $\pm 0.014$ $\pm 0.085$	17.4k 7,8 AKHMETSHIN 05	CMD2	0.60–1.38 $e^+ e^- \rightarrow \eta\gamma$
1.287 $\pm 0.013$ $\pm 0.063$	9,10 AKHMETSHIN 01B	CMD2	$e^+ e^- \rightarrow \eta\gamma$
1.338 $\pm 0.012$ $\pm 0.052$	11 ACHASOV 00	SND	$e^+ e^- \rightarrow \eta\gamma$
1.18 $\pm 0.03$ $\pm 0.06$	2200 12 AKHMETSHIN 99F	CMD2	$e^+ e^- \rightarrow \eta\gamma$
1.21 $\pm 0.07$	13 BENAYOUN 96	RVUE	0.54–1.04 $e^+ e^- \rightarrow \eta\gamma$

<sup>1</sup> Using  $B(\phi \rightarrow e^+ e^-) = (2.99 \pm 0.08) \times 10^{-4}$  and  $B(\eta \rightarrow 3\pi^0) = (32.2 \pm 0.4) \times 10^{-2}$ .

<sup>2</sup> From  $\pi^+ \pi^- \pi^0$  decay mode of  $\eta$ .

<sup>3</sup> From  $2\gamma$  decay mode of  $\eta$ .

<sup>4</sup> From  $3\pi^0$  decay mode of  $\eta$ .

<sup>5</sup> Combined analysis of the CMD-2 data on  $\phi \rightarrow K^+ K^-$ ,  $K_S^0 K_L^0$ ,  $\pi^+ \pi^- \pi^0$ ,  $\eta\gamma$  assuming that the sum of their branching fractions is  $0.99741 \pm 0.00007$ .

<sup>6</sup> ACHASOV 07B reports  $[\Gamma(\phi(1020) \rightarrow \eta\gamma)/\Gamma_{\text{total}}] \times [B(\phi(1020) \rightarrow e^+ e^-)] = (4.050 \pm 0.067 \pm 0.118) \times 10^{-6}$  which we divide by our best value  $B(\phi(1020) \rightarrow e^+ e^-) = (2.979 \pm 0.033) \times 10^{-4}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value. Supersedes ACHASOV 00D and ACHASOV 06A.

<sup>7</sup> Using  $B(\phi \rightarrow e^+ e^-) = (2.98 \pm 0.04) \times 10^{-4}$  and  $B(\eta \rightarrow \gamma\gamma) = 39.43 \pm 0.26\%$ .

<sup>8</sup> Not independent of the corresponding  $\Gamma(e^+ e^-) \times \Gamma(\eta\gamma)/\Gamma_{\text{total}}^2$ .

<sup>9</sup> Using  $B(\phi \rightarrow e^+ e^-) = (2.99 \pm 0.08) \times 10^{-4}$  and  $B(\eta \rightarrow 3\pi^0) = (32.24 \pm 0.29) \times 10^{-2}$ .

<sup>10</sup> The combined fit from 600 to 1380 MeV taking into account  $\rho(770)$ ,  $\omega(782)$ ,  $\phi(1020)$ , and  $\rho(1450)$  (mass and width fixed at 1450 MeV and 310 MeV respectively).

<sup>11</sup> From the  $\eta \rightarrow 2\gamma$  decay and using  $B(\phi \rightarrow e^+ e^-) = (2.99 \pm 0.08) \times 10^{-4}$ .

<sup>12</sup> From  $\pi^+ \pi^- \pi^0$  decay mode of  $\eta$  and using  $B(\phi \rightarrow e^+ e^-) = (2.99 \pm 0.08) \times 10^{-4}$ .

<sup>13</sup> Reanalysis of DRUZHININ 84, DOLINSKY 89, and DOLINSKY 91 taking into account a triangle anomaly contribution.

## $\Gamma(\pi^0\gamma)/\Gamma_{\text{total}}$

VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
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**1.32  $\pm 0.05$  OUR FIT**

**1.31  $\pm 0.13$  OUR AVERAGE**

1.30 $\pm 0.13$	DRUZHININ 84	ND	$e^+ e^- \rightarrow 3\gamma$
1.4 $\pm 0.5$	32 COSME 76	OSPK	$e^+ e^-$

• • • We do not use the following data for averages, fits, limits, etc. • • •

1.367 $\pm 0.072$	1 ACHASOV 16A	SND	$0.60–1.38 e^+ e^- \rightarrow \pi^0\gamma$
1.258 $\pm 0.037$ $\pm 0.077$	18k 2,3 AKHMETSHIN 05	CMD2	$0.60–1.38 e^+ e^- \rightarrow \pi^0\gamma$

1.226 $\pm 0.036$ $^{+0.096}_{-0.089}$	4 ACHASOV 00	SND	$e^+ e^- \rightarrow \pi^0\gamma$
1.26 $\pm 0.17$	5 BENAYOUN 96	RVUE	$0.54–1.04 e^+ e^- \rightarrow \pi^0\gamma$

<sup>1</sup> Using  $B(\phi \rightarrow e^+ e^-)$  from PDG 15. Supersedes ACHASOV 00.

<sup>2</sup> Using  $B(\phi \rightarrow e^+ e^-) = (2.98 \pm 0.04) \times 10^{-4}$ .

<sup>3</sup> Not independent of the corresponding  $\Gamma(e^+ e^-) \times \Gamma(\pi^0\gamma)/\Gamma_{\text{total}}^2$ .

<sup>4</sup> From the  $\pi^0 \rightarrow 2\gamma$  decay and using  $B(\phi \rightarrow e^+ e^-) = (2.99 \pm 0.08) \times 10^{-4}$ .

<sup>5</sup> Reanalysis of DRUZHININ 84, DOLINSKY 89, and DOLINSKY 91 taking into account a triangle anomaly contribution.

## $\Gamma_6/\Gamma$

$\Gamma(\eta\gamma)/\Gamma(\pi^0\gamma)$		$\Gamma_5/\Gamma_6$	
<u>VALUE</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
$\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$			
$10.9 \pm 0.3^{+0.7}_{-0.8}$	ACHASOV 00	SND	$e^+ e^- \rightarrow \eta\gamma, \pi^0\gamma$

$\Gamma(e^+ e^-)/\Gamma_{\text{total}}$		$\Gamma_8/\Gamma$		
<u>VALUE (units <math>10^{-4}</math>)</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b><math>2.979 \pm 0.033</math> OUR FIT</b>		Error includes scale factor of 1.2.		
<b><math>2.98 \pm 0.07</math> OUR AVERAGE</b>		Error includes scale factor of 1.1.		
2.93 $\pm 0.14$	1900k	1 ACHASOV 01E	SND	$e^+ e^- \rightarrow K^+ K^-, K_S K_L, \pi^+ \pi^- \pi^0$
2.88 $\pm 0.09$	55600	AKHMETSHIN 95	CMD2	$e^+ e^- \rightarrow \text{hadrons}$
3.00 $\pm 0.21$	3681	BUKIN 78C	OLYA	$e^+ e^- \rightarrow \text{hadrons}$
3.10 $\pm 0.14$		2 PARROUR 76	OSPK	$e^+ e^-$
3.3 $\pm 0.3$		COSME 74	OSPK	$e^+ e^- \rightarrow \text{hadrons}$
2.81 $\pm 0.25$	681	BALAKIN 71	OSPK	$e^+ e^- \rightarrow \text{hadrons}$
3.50 $\pm 0.27$		CHATELUS 71	OSPK	$e^+ e^-$

<sup>1</sup> From the combined fit assuming that the total  $\phi(1020)$  production cross section is saturated by those of  $K^+ K^-$ ,  $K_S K_L$ ,  $\pi^+ \pi^- \pi^0$ , and  $\eta\gamma$  decays modes and using ACHASOV 00B for the  $\eta\gamma$  decay mode.

<sup>2</sup> Using total width 4.2 MeV. They detect  $3\pi$  mode and observe significant interference with  $\omega$  tail. This is accounted for in the result quoted above.

$\Gamma(\mu^+ \mu^-)/\Gamma_{\text{total}}$		$\Gamma_9/\Gamma$	
<u>VALUE (units <math>10^{-4}</math>)</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b><math>2.85 \pm 0.22</math> OUR FIT</b>	Error includes scale factor of 1.2.		
<b><math>2.5 \pm 0.4</math> OUR AVERAGE</b>			
2.69 $\pm 0.46$	1 HAYES 71	CNTR	$8.3, 9.8 \gamma C \rightarrow \mu^+ \mu^- X$
2.17 $\pm 0.60$	1 EARLES 70	CNTR	$6.0 \gamma C \rightarrow \mu^+ \mu^- X$
$\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$			
2.87 $\pm 0.20 \pm 0.14$	2 ACHASOV 01G	SND	$e^+ e^- \rightarrow \mu^+ \mu^-$
3.30 $\pm 0.45 \pm 0.32$	3 ACHASOV 99C	SND	$e^+ e^- \rightarrow \mu^+ \mu^-$
4.83 $\pm 1.02$	4 VASSERMAN 81	OLYA	$e^+ e^- \rightarrow \mu^+ \mu^-$
2.87 $\pm 1.98$	4 AUGUSTIN 73	OSPK	$e^+ e^- \rightarrow \mu^+ \mu^-$

<sup>1</sup> Neglecting interference between resonance and continuum.

<sup>2</sup> Using  $B(\phi \rightarrow e^+ e^-) = (2.91 \pm 0.07) \times 10^{-4}$ .

<sup>3</sup> Using  $B(\phi \rightarrow e^+ e^-) = (2.99 \pm 0.08) \times 10^{-4}$ .

<sup>4</sup> Recalculated by us using  $B(\phi \rightarrow e^+ e^-) = (2.99 \pm 0.08) \times 10^{-4}$ .

$\Gamma(\eta e^+ e^-)/\Gamma_{\text{total}}$		$\Gamma_{10}/\Gamma$		
<u>VALUE (units <math>10^{-4}</math>)</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b><math>1.08 \pm 0.04</math> OUR AVERAGE</b>				
1.075 $\pm 0.007 \pm 0.038$	30k	1 BABUSCI 15	KLOE	$1.02 e^+ e^- \rightarrow \eta e^+ e^-$
1.19 $\pm 0.19 \pm 0.12$	213	2 ACHASOV 01B	SND	$e^+ e^- \rightarrow \eta e^+ e^-$
1.14 $\pm 0.10 \pm 0.06$	355	3 AKHMETSHIN 01	CMD2	$e^+ e^- \rightarrow \eta e^+ e^-$
$\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$				
1.13 $\pm 0.14 \pm 0.07$	183	4 AKHMETSHIN 01	CMD2	$e^+ e^- \rightarrow \eta e^+ e^-$
1.21 $\pm 0.14 \pm 0.09$	130	5 AKHMETSHIN 01	CMD2	$e^+ e^- \rightarrow \eta e^+ e^-$

1.04	$\pm 0.20$	$\pm 0.08$	42	<sup>6</sup> AKHMETSHIN 01	CMD2	$e^+ e^- \rightarrow \eta e^+ e^-$	
1.3	$+0.8$	$-0.6$	7	GOLUBEV	85	ND	$e^+ e^- \rightarrow \eta e^+ e^-$

<sup>1</sup> Using  $B(\eta \rightarrow 3\pi^0) = (32.57 \pm 0.23)\%$  from PDG 12.

<sup>2</sup> Using  $B(\eta \rightarrow \gamma\gamma) = (39.25 \pm 0.32)\%$ ,  $B(\phi \rightarrow \eta\gamma) = (1.26 \pm 0.06)\%$ , and  $B(\phi \rightarrow e^+ e^-) = (3.00 \pm 0.06) \times 10^{-4}$ .

<sup>3</sup> The average of the branching ratios separately obtained from the  $\eta \rightarrow \gamma\gamma$ ,  $3\pi^0$ ,  $\pi^+\pi^-\pi^0$  decays.

<sup>4</sup> From  $\eta \rightarrow \gamma\gamma$  decays and using  $B(\eta \rightarrow \gamma\gamma) = (39.33 \pm 0.25) \times 10^{-2}$ ,  $B(\eta \rightarrow \pi^+\pi^-\gamma) = (4.75 \pm 11) \times 10^{-2}$ , and  $B(\phi \rightarrow \eta\gamma) = (1.297 \pm 0.033) \times 10^{-2}$ .

<sup>5</sup> From  $\eta \rightarrow 3\pi^0$  decays and using  $B(\pi^0 \rightarrow \gamma\gamma) = (98.798 \pm 0.033) \times 10^{-2}$ ,  $B(\eta \rightarrow 3\pi^0) = (32.24 \pm 0.29) \times 10^{-2}$ ,  $B(\eta \rightarrow \pi^+\pi^-\gamma) = (4.75 \pm 0.11) \times 10^{-2}$ , and  $B(\phi \rightarrow \eta\gamma) = (1.297 \pm 0.033) \times 10^{-2}$ .

<sup>6</sup> From  $\eta \rightarrow \pi^+\pi^-\pi^0$  decays and using  $B(\pi^0 \rightarrow \gamma\gamma) = (98.798 \pm 0.033) \times 10^{-2}$ ,  $B(\pi^0 \rightarrow e^+e^-\gamma) = (1.198 \pm 0.032) \times 10^{-2}$ ,  $B(\eta \rightarrow \pi^+\pi^-\pi^0) = (23.0 \pm 0.4) \times 10^{-2}$ ,  $B(\phi \rightarrow \pi^+\pi^-\pi^0) = (15.5 \pm 0.6) \times 10^{-2}$ , and  $B(\phi \rightarrow \eta\gamma) = (1.297 \pm 0.033) \times 10^{-2}$ .

### $\Gamma(\pi^+\pi^-)/\Gamma_{\text{total}}$

### $\Gamma_{11}/\Gamma$

VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>• • • We do not use the following data for averages, fits, limits, etc. • • •</b>				
0.71 $\pm 0.11 \pm 0.09$		<sup>1</sup> ACHASOV 00C	SND	$e^+ e^- \rightarrow \pi^+\pi^-$
0.65 $^{+0.38}_{-0.29}$		<sup>1</sup> GOLUBEV 86	ND	$e^+ e^- \rightarrow \pi^+\pi^-$
2.01 $^{+1.07}_{-0.84}$		<sup>1</sup> VASSERMAN 81	OLYA	$e^+ e^- \rightarrow \pi^+\pi^-$
<6.6	95	BUKIN 78B	OLYA	$e^+ e^- \rightarrow \pi^+\pi^-$
<2.7	95	ALVENSLEB... 72	CNTR	$6.7 \gamma C \rightarrow C\pi^+\pi^-$

<sup>1</sup> Using  $B(\phi \rightarrow e^+e^-) = (2.99 \pm 0.08) \times 10^{-4}$ .

### $\Gamma(\omega\pi^0)/\Gamma_{\text{total}}$

### $\Gamma_{12}/\Gamma$

VALUE (units $10^{-5}$ )	DOCUMENT ID	TECN	COMMENT
<b>4.7 <math>\pm 0.5</math> OUR FIT</b>			
<b>5.2 <math>^{+1.3}_{-1.1}</math></b>	<sup>1,2</sup> AULCHENKO 00A	SND	$e^+ e^- \rightarrow \pi^+\pi^-\pi^0\pi^0$
4.4 $\pm 0.6$	<sup>3</sup> AMBROSINO 08G	KLOE	$e^+ e^- \rightarrow \pi^+\pi^-2\pi^0, 2\pi^0\gamma$
$\sim 5.4$	<sup>4</sup> ACHASOV 00E	SND	$e^+ e^- \rightarrow \pi^0\pi^0\gamma$
5.5 $^{+1.6}_{-1.4} \pm 0.3$	<sup>2,5</sup> AULCHENKO 00A	SND	$e^+ e^- \rightarrow \pi^+\pi^-\pi^0\pi^0$
4.8 $^{+1.9}_{-1.7} \pm 0.8$	<sup>4</sup> ACHASOV 99	SND	$e^+ e^- \rightarrow \pi^+\pi^-\pi^0\pi^0$

<sup>1</sup> Using the 1996 and 1998 data.

<sup>2</sup> ( $2.3 \pm 0.3$ )% correction for other decay modes of the  $\omega(782)$  applied.

<sup>3</sup> Not independent of the corresponding  $\Gamma(\omega\pi^0) \times \Gamma(e^+e^-) / \Gamma^2(\text{total})$ .

<sup>4</sup> Using the 1996 data.

<sup>5</sup> Using the 1998 data.

### $\Gamma(\omega\gamma)/\Gamma_{\text{total}}$

### $\Gamma_{13}/\Gamma$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<0.05	84	LINDSEY 66	HBC	2.1–2.7 $K^- p \rightarrow \Lambda\pi^+\pi^-$ neutrals

$\Gamma(\rho\gamma)/\Gamma_{\text{total}}$			$\Gamma_{14}/\Gamma$		
VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
< <b>0.12</b>	90	1 AKHMETSHIN 99B	CMD2	$e^+ e^- \rightarrow \pi^+ \pi^- \gamma$	
$\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$					
< 7	90	AKHMETSHIN 97C	CMD2	$e^+ e^- \rightarrow \pi^+ \pi^- \gamma$	
<200	84	LINDSEY	66 HBC	2.1–2.7 $K^- p \rightarrow \Lambda \pi^+ \pi^-$ neutrals	

<sup>1</sup> Supersedes AKHMETSHIN 97C.

$\Gamma(\pi^+ \pi^- \gamma)/\Gamma_{\text{total}}$			$\Gamma_{15}/\Gamma$		
VALUE (units $10^{-4}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.41±0.12±0.04</b>	30175	1 AKHMETSHIN 99B	CMD2	$e^+ e^- \rightarrow \pi^+ \pi^- \gamma$	
$\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$					
< 0.3	90	2 AKHMETSHIN 97C	CMD2	$e^+ e^- \rightarrow \pi^+ \pi^- \gamma$	
<600	90	KALBFLEISCH 75	HBC	2.18 $K^- p \rightarrow \Lambda \pi^+ \pi^- \gamma$	
< 70	90	COSME	74 OSPK	$e^+ e^- \rightarrow \pi^+ \pi^- \gamma$	
<400	90	LINDSEY	65 HBC	2.1–2.7 $K^- p \rightarrow \Lambda \pi^+ \pi^-$ neutrals	

<sup>1</sup> For  $E_\gamma > 20$  MeV and assuming that  $B(\phi(1020) \rightarrow f_0(980)\gamma)$  is negligible. Supersedes AKHMETSHIN 97C.

<sup>2</sup> For  $E_\gamma > 20$  MeV and assuming that  $B(\phi(1020) \rightarrow f_0(980)\gamma)$  is negligible.

$\Gamma(f_0(980)\gamma)/\Gamma_{\text{total}}$			$\Gamma_{16}/\Gamma$		
VALUE (units $10^{-4}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>3.22±0.19 OUR FIT</b>	Error includes scale factor of 1.1.				
<b>3.21±0.19 OUR AVERAGE</b>					
$\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$					
3.21 <sup>+0.03</sup> <sub>-0.09</sub>	18	1 AMBROSINO 07	KLOE	$e^+ e^- \rightarrow \pi^0 \pi^0 \gamma$	
2.90±0.21±1.54		2 AKHMETSHIN 99C	CMD2	$e^+ e^- \rightarrow \pi^+ \pi^- \gamma, \pi^0 \pi^0 \gamma$	
4.47±0.21	2438	3 ALOISIO 02D	KLOE	$e^+ e^- \rightarrow \pi^0 \pi^0 \gamma$	
3.5 ± 0.3 <sup>+1.3</sup> <sub>-0.5</sub>	419	4,5 ACHASOV 00H	SND	$e^+ e^- \rightarrow \pi^0 \pi^0 \gamma$	
1.93±0.46±0.50	27188	6 AKHMETSHIN 99B	CMD2	$e^+ e^- \rightarrow \pi^+ \pi^- \gamma$	
3.05±0.25±0.72	268	7 AKHMETSHIN 99C	CMD2	$e^+ e^- \rightarrow \pi^0 \pi^0 \gamma$	
1.5 ± 0.5	268	8 AKHMETSHIN 99C	CMD2	$e^+ e^- \rightarrow \pi^0 \pi^0 \gamma$	
3.42±0.30±0.36	164	4 ACHASOV 98I	SND	$e^+ e^- \rightarrow 5\gamma$	
< 1	90	9 AKHMETSHIN 97C	CMD2	$e^+ e^- \rightarrow \pi^+ \pi^- \gamma$	
< 7	90	10 AKHMETSHIN 97C	CMD2	$e^+ e^- \rightarrow \pi^+ \pi^- \gamma$	
< 20	90	DRUZHININ 87	ND	$e^+ e^- \rightarrow \pi^0 \pi^0 \gamma$	

<sup>1</sup> Obtained by the authors taking into account the  $\pi^+ \pi^-$  decay mode. Includes a component due to  $\pi\pi$  production via the  $f_0(500)$  meson. Supersedes ALOISIO 02D.

<sup>2</sup> From the combined fit of the photon spectra in the reactions  $e^+ e^- \rightarrow \pi^+ \pi^- \gamma, \pi^0 \pi^0 \gamma$ .

<sup>3</sup> From the negative interference with the  $f_0(500)$  meson of AITALA 01B using the ACHASOV 89 parameterization for the  $f_0(980)$ , a Breit-Wigner for the  $f_0(500)$ , and ACHASOV 01F for the  $\rho\pi$  contribution. Superseded by AMBROSINO 07.

<sup>4</sup> Assuming that the  $\pi^0\pi^0\gamma$  final state is completely determined by the  $f_0\gamma$  mechanism, neglecting the decay  $B(\phi \rightarrow K\bar{K}\gamma)$  and using  $B(f_0 \rightarrow \pi^+\pi^-) = 2B(f_0 \rightarrow \pi^0\pi^0)$ .

<sup>5</sup> Using the value  $B(\phi \rightarrow \eta\gamma) = (1.338 \pm 0.053) \times 10^{-2}$ .

<sup>6</sup> For  $E_\gamma > 20$  MeV. Supersedes AKHMETSHIN 97c.

<sup>7</sup> Neglecting other intermediate mechanisms ( $\rho\pi$ ,  $\sigma\gamma$ ).

<sup>8</sup> A narrow pole fit taking into account  $f_0(980)$  and  $f_0(1200)$  intermediate mechanisms.

<sup>9</sup> For destructive interference with the Bremsstrahlung process

<sup>10</sup> For constructive interference with the Bremsstrahlung process

### $\Gamma(f_0(980)\gamma)/\Gamma(\eta\gamma)$

### $\Gamma_{16}/\Gamma_5$

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>2.48 \pm 0.15</math> OUR FIT</b>	Error includes scale factor of 1.1.			

**$2.6 \pm 0.2$**  <sup>+0.8</sup><sub>-0.3</sub>      419      <sup>1</sup> ACHASOV      00H SND       $e^+e^- \rightarrow \pi^0\pi^0\gamma$

<sup>1</sup> Assuming that the  $\pi^0\pi^0\gamma$  final state is completely determined by the  $f_0\gamma$  mechanism, neglecting the decay  $B(\phi \rightarrow K\bar{K}\gamma)$  and using  $B(f_0 \rightarrow \pi^+\pi^-) = 2B(f_0 \rightarrow \pi^0\pi^0)$ .

### $\Gamma(\pi^0\pi^0\gamma)/\Gamma_{\text{total}}$

### $\Gamma_{17}/\Gamma$

VALUE (units $10^{-4}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>1.07 \pm 0.06</math> OUR AVERAGE</b>					

$1.07 \pm 0.01$	$+0.06$		<sup>1</sup> AMBROSINO	07	KLOE	$e^+e^- \rightarrow \pi^0\pi^0\gamma$
$-0.03$	$-0.06$					
$1.08 \pm 0.17$	$\pm 0.09$	268	AKHMETSHIN 99c	CMD2	$e^+e^- \rightarrow \pi^0\pi^0\gamma$	
<b>• • • We do not use the following data for averages, fits, limits, etc. • • •</b>						
$1.09 \pm 0.03$	$\pm 0.05$	2438	ALOISIO	02D	KLOE	$e^+e^- \rightarrow \pi^0\pi^0\gamma$
$1.158 \pm 0.093 \pm 0.052$		419	<sup>2,3</sup> ACHASOV	00H	SND	$e^+e^- \rightarrow \pi^0\pi^0\gamma$
$< 10$		90	DRUZHININ	87	ND	$e^+e^- \rightarrow 5\gamma$

<sup>1</sup> Supersedes ALOISIO 02D.

<sup>2</sup> Using the value  $B(\phi \rightarrow \eta\gamma) = (1.338 \pm 0.053) \times 10^{-2}$ .

<sup>3</sup> Supersedes ACHASOV 98l. Excluding  $\omega\pi^0$ .

### $\Gamma(\pi^0\pi^0\gamma)/\Gamma(\eta\gamma)$

### $\Gamma_{17}/\Gamma_5$

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>0.86 \pm 0.04</math> OUR FIT</b>				

**$0.865 \pm 0.070 \pm 0.017$**       419      <sup>1</sup> ACHASOV      00H SND       $e^+e^- \rightarrow \pi^0\pi^0\gamma$

**• • • We do not use the following data for averages, fits, limits, etc. • • •**

$0.90 \pm 0.08 \pm 0.07$       164      ACHASOV      98l SND       $e^+e^- \rightarrow 5\gamma$

<sup>1</sup> Supersedes ACHASOV 98l. Excluding  $\omega\pi^0$ .

### $\Gamma(\pi^+\pi^-\pi^+\pi^-)/\Gamma_{\text{total}}$

### $\Gamma_{18}/\Gamma$

VALUE (units $10^{-6}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>6.5 \pm 2.7 \pm 1.6</math></b>		6.8k	<sup>1</sup> AKHMETSHIN 17	CMD3	$e^+e^- \rightarrow \pi^+\pi^-\pi^+\pi^-$

**• • • We do not use the following data for averages, fits, limits, etc. • • •**

$3.93 \pm 1.74 \pm 2.14$	3.3k	AKHMETSHIN 00E	CMD2	$e^+e^- \rightarrow \pi^+\pi^-\pi^+\pi^-$	
$< 870$	90	CORDIER	79	WIRE	$e^+e^- \rightarrow \pi^+\pi^-\pi^+\pi^-$

<sup>1</sup> Using the cross section at the  $\phi$  meson peak  $\sigma(\phi) = 4172 \pm 42$  nb, the nonresonant cross section  $\sigma(0) = 1.263 \pm 0.027$  nb and  $\text{Re}(Z) = 0.146 \pm 0.030$ ,  $\text{Im}(Z) = -0.002 \pm 0.024$  for the complex amplitude of the  $\phi \rightarrow \pi^+\pi^-\pi^+\pi^-$  transition.

$\Gamma(\pi^+\pi^+\pi^-\pi^-\pi^0)/\Gamma_{\text{total}}$   $\Gamma_{19}/\Gamma$ 

<u>VALUE (units <math>10^{-6}</math>)</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
< <b>4.6</b>	90	AKHMETSHIN 00E	CMD2	$e^+e^- \rightarrow \pi^+\pi^-\pi^+\pi^-\pi^0$
<b>• • • We do not use the following data for averages, fits, limits, etc. • • •</b>				
<150	95	BARKOV	88	CMD $e^+e^- \rightarrow \pi^+\pi^-\pi^+\pi^-\pi^0$

 $\Gamma(\pi^0 e^+ e^-)/\Gamma_{\text{total}}$   $\Gamma_{20}/\Gamma$ 

<u>VALUE (units <math>10^{-5}</math>)</u>	<u>CL%</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>1.33<sup>+0.07</sup><sub>-0.10</sub> OUR AVERAGE</b>					
1.35 $\pm$ 0.05 <sup>+0.05</sup> <sub>-0.10</sub>	9.5k	<sup>1</sup> ANASTASI	16B	KLOE	$e^+e^- \rightarrow \pi^0e^+e^-$
1.01 $\pm$ 0.28 $\pm$ 0.29	52	<sup>2</sup> ACHASOV	02D	SND	$e^+e^- \rightarrow \pi^0e^+e^-$
1.22 $\pm$ 0.34 $\pm$ 0.21	46	<sup>3</sup> AKHMETSHIN 01C	CMD2		$e^+e^- \rightarrow \pi^0e^+e^-$
<b>• • • We do not use the following data for averages, fits, limits, etc. • • •</b>					
<12	90	DOLINSKY	88	ND	$e^+e^- \rightarrow \pi^0e^+e^-$

<sup>1</sup> Using  $B(\pi^0 \rightarrow \gamma\gamma)$  from the 2014 Edition of this Review (PDG 14).

<sup>2</sup> Using various branching ratios from the 2000 Edition of this Review (PDG 00).

<sup>3</sup> Using  $B(\pi^0 \rightarrow \gamma\gamma) = 0.98798 \pm 0.00032$ ,  $B(\phi \rightarrow \eta\gamma) = (1.297 \pm 0.033) \times 10^{-2}$ , and  $B(\eta \rightarrow \pi^+\pi^-\gamma) = (4.75 \pm 0.11) \times 10^{-2}$ .

 $\Gamma(\pi^0\eta\gamma)/\Gamma_{\text{total}}$   $\Gamma_{21}/\Gamma$ 

<u>VALUE (units <math>10^{-5}</math>)</u>	<u>CL%</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>7.27<math>\pm</math>0.30 OUR AVERAGE</b> Error includes scale factor of 1.5. See the ideogram below.					
7.06 $\pm$ 0.22	16.9k	<sup>1</sup> AMBROSINO	09F	KLOE	$1.02 e^+e^- \rightarrow \eta\pi^0\gamma$
8.51 $\pm$ 0.51 $\pm$ 0.57	607	<sup>2</sup> ALOISIO	02C	KLOE	$e^+e^- \rightarrow \eta\pi^0\gamma$
7.96 $\pm$ 0.60 $\pm$ 0.40	197	<sup>3</sup> ALOISIO	02C	KLOE	$e^+e^- \rightarrow \eta\pi^0\gamma$
8.8 $\pm$ 1.4 $\pm$ 0.9	36	<sup>4</sup> ACHASOV	00F	SND	$e^+e^- \rightarrow \eta\pi^0\gamma$
9.0 $\pm$ 2.4 $\pm$ 1.0	80	AKHMETSHIN 99C	CMD2		$e^+e^- \rightarrow \eta\pi^0\gamma$
<b>• • • We do not use the following data for averages, fits, limits, etc. • • •</b>					
7.01 $\pm$ 0.10 $\pm$ 0.20	13.3k	<sup>2,5</sup> AMBROSINO	09F	KLOE	$1.02 e^+e^- \rightarrow \eta\pi^0\gamma$
7.12 $\pm$ 0.13 $\pm$ 0.22	3.6k	<sup>3,6</sup> AMBROSINO	09F	KLOE	$1.02 e^+e^- \rightarrow \eta\pi^0\gamma$
8.3 $\pm$ 2.3 $\pm$ 1.2	20	ACHASOV	98B	SND	$e^+e^- \rightarrow 5\gamma$
<250	90	DOLINSKY	91	ND	$e^+e^- \rightarrow \pi^0\eta\gamma$

<sup>1</sup> Combined results of  $\eta \rightarrow \gamma\gamma$  and  $\eta \rightarrow \pi^+\pi^-\pi^0$  decay modes measurements.

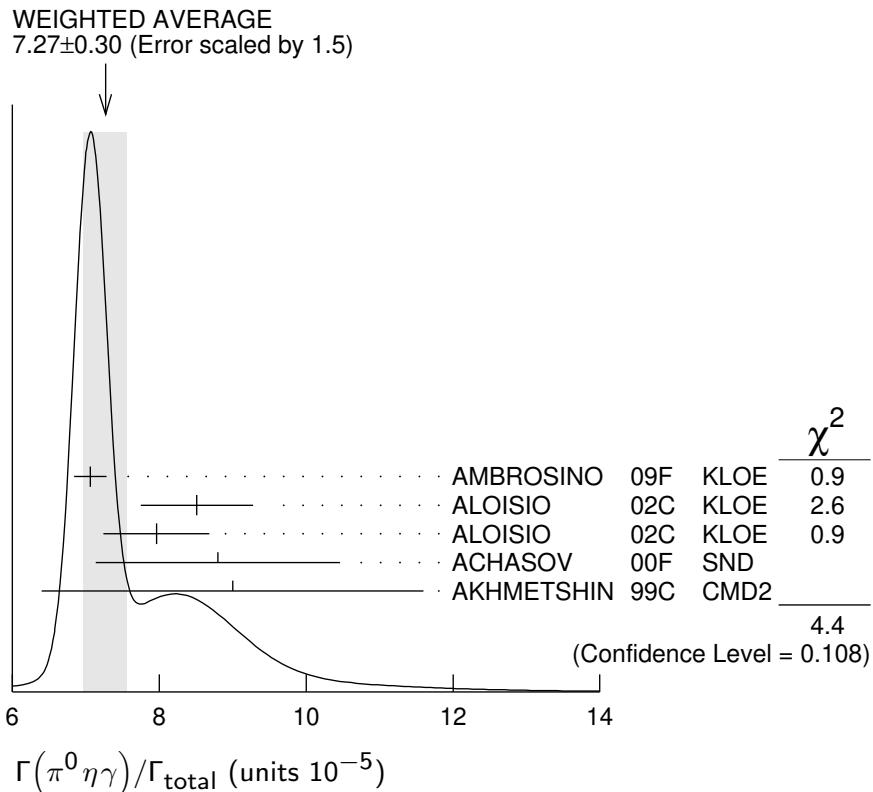
<sup>2</sup> From the decay mode  $\eta \rightarrow \gamma\gamma$ .

<sup>3</sup> From the decay mode  $\eta \rightarrow \pi^+\pi^-\pi^0$ .

<sup>4</sup> Supersedes ACHASOV 98B.

<sup>5</sup> Using  $B(\phi \rightarrow \eta\gamma) = (1.304 \pm 0.025)\%$ ,  $B(\eta \rightarrow 3\pi^0) = (32.56 \pm 0.23)\%$ , and  $B(\eta \rightarrow \gamma\gamma) = (39.31 \pm 0.20)\%$ .

<sup>6</sup> Using  $B(\phi \rightarrow \eta\gamma) = (1.304 \pm 0.025)\%$ ,  $B(\eta \rightarrow 3\pi^0) = (32.56 \pm 0.23)\%$ , and  $B(\eta \rightarrow \pi^+\pi^-\pi^0) = (22.73 \pm 0.28)\%$ .

 $\Gamma(a_0(980)\gamma)/\Gamma_{\text{total}}$  $\Gamma_{22}/\Gamma$ 

VALUE (units $10^{-5}$ )	CL%	EVTS
<b>7.6±0.6 OUR FIT</b>		
<b>7.6±0.6 OUR AVERAGE</b>		
7.4±0.7		
8.8±1.7	36	
<500	90	

DOCUMENT ID

TECN

COMMENT

<sup>1</sup> ALOISIO 02C KLOE  $e^+ e^- \rightarrow \eta \pi^0 \gamma$   
<sup>2</sup> ACHASOV 00F SND  $e^+ e^- \rightarrow \eta \pi^0 \gamma$

• • • We do not use the following data for averages, fits, limits, etc. • • •

<sup>3</sup> GOKALP 02 RVUE  $e^+ e^- \rightarrow \eta \pi^0 \gamma$   
DOLINSKY 91 ND  $e^+ e^- \rightarrow \pi^0 \eta \gamma$

<sup>1</sup> Using  $M_{a_0(980)}=984.8$  MeV and assuming  $a_0(980)\gamma$  dominance.

<sup>2</sup> Assuming  $a_0(980)\gamma$  dominance in the  $\eta \pi^0 \gamma$  final state.

<sup>3</sup> Using data of ACHASOV 00F.

 $\Gamma(f_0(980)\gamma)/\Gamma(a_0(980)\gamma)$  $\Gamma_{16}/\Gamma_{22}$ 

VALUE	CL%	EVTS
<b>6.1±0.6</b>		

DOCUMENT ID

TECN

COMMENT

<sup>1</sup> ALOISIO 02C KLOE  $e^+ e^- \rightarrow \eta \pi^0 \gamma$

<sup>1</sup> Using results of ALOISIO 02D and assuming that  $f_0(980)$  decays into  $\pi\pi$  only and  $a_0(980)$  into  $\eta\pi$  only.

 $\Gamma(K^0 \bar{K}^0 \gamma)/\Gamma_{\text{total}}$  $\Gamma_{23}/\Gamma$ 

VALUE	CL%	EVTS
<b>&lt;1.9 × 10<sup>-8</sup></b>	90	

DOCUMENT ID

TECN

COMMENT

AMBROSINO 09C KLOE  $e^+ e^- \rightarrow K_S^0 \bar{K}_S^0 \gamma$

$\Gamma(\eta'(958)\gamma)/\Gamma_{\text{total}}$				$\Gamma_{24}/\Gamma$
<u>VALUE (units <math>10^{-5}</math>)</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b><math>6.21 \pm 0.20</math> OUR FIT</b>				
<b><math>6.21 \pm 0.29</math> OUR AVERAGE</b>				
$6.21 \pm 0.27 \pm 0.11$	3407	<sup>1</sup> AMBROSINO 07A	KLOE	$1.02 e^+ e^- \rightarrow \pi^+ \pi^- 7\gamma$
$6.7 \begin{array}{l} +2.8 \\ -2.4 \end{array} \pm 0.8$	12	<sup>2</sup> AULCHENKO 03B	SND	$e^+ e^- \rightarrow \eta' \gamma$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$6.7 \begin{array}{l} +5.0 \\ -4.2 \end{array} \pm 1.5$	7	AULCHENKO 03B	SND	$e^+ e^- \rightarrow 7\gamma$
$6.10 \pm 0.61 \pm 0.43$	120	<sup>3</sup> ALOISIO 02E	KLOE	$1.02 e^+ e^- \rightarrow \pi^+ \pi^- 3\gamma$
$8.2 \begin{array}{l} +2.1 \\ -1.9 \end{array} \pm 1.1$	21	<sup>4</sup> AKHMETSHIN 00B	CMD2	$e^+ e^- \rightarrow \pi^+ \pi^- 3\gamma$
$4.9 \begin{array}{l} +2.2 \\ -1.8 \end{array} \pm 0.6$	9	<sup>5</sup> AKHMETSHIN 00F	CMD2	$e^+ e^- \rightarrow \pi^+ \pi^- \pi^+ \pi^- \geq 2\gamma$
$6.4 \pm 1.6$	30	<sup>6</sup> AKHMETSHIN 00F	CMD2	$e^+ e^- \rightarrow \eta'(958)\gamma$
$6.7 \begin{array}{l} +3.4 \\ -2.9 \end{array} \pm 1.0$	5	<sup>7</sup> AULCHENKO 99	SND	$e^+ e^- \rightarrow \pi^+ \pi^- 3\gamma$
$<11$	90	AULCHENKO 98	SND	$e^+ e^- \rightarrow 7\gamma$
$12 \begin{array}{l} +7 \\ -5 \end{array} \pm 2$	6	<sup>4</sup> AKHMETSHIN 97B	CMD2	$e^+ e^- \rightarrow \pi^+ \pi^- 3\gamma$
$<41$	90	DRUZHININ 87	ND	$e^+ e^- \rightarrow \gamma \eta \pi^+ \pi^-$

<sup>1</sup> AMBROSINO 07A reports  $[\Gamma(\phi(1020) \rightarrow \eta'(958)\gamma)/\Gamma_{\text{total}}] / [\mathcal{B}(\phi(1020) \rightarrow \eta\gamma)] = (4.77 \pm 0.09 \pm 0.19) \times 10^{-3}$  which we multiply by our best value  $\mathcal{B}(\phi(1020) \rightarrow \eta\gamma) = (1.301 \pm 0.024) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

<sup>2</sup> Averaging AULCHENKO 03B with AULCHENKO 99.

<sup>3</sup> Using  $\mathcal{B}(\phi \rightarrow \eta\gamma) = (1.297 \pm 0.033)\%$ .

<sup>4</sup> Using the value  $\mathcal{B}(\phi \rightarrow \eta\gamma) = (1.26 \pm 0.06) \times 10^{-2}$ .

<sup>5</sup> Using  $\mathcal{B}(\phi \rightarrow K_L^0 K_S^0) = (33.8 \pm 0.6)\%$ .

<sup>6</sup> Averaging AKHMETSHIN 00B with AKHMETSHIN 00F.

<sup>7</sup> Using the value  $\mathcal{B}(\eta' \rightarrow \eta\pi^+\pi^-) = (43.7 \pm 1.5) \times 10^{-2}$  and  $\mathcal{B}(\eta \rightarrow \gamma\gamma) = (39.25 \pm 0.31) \times 10^{-2}$ .

$\Gamma(\eta'(958)\gamma)/\Gamma(K_L^0 K_S^0)$				$\Gamma_{24}/\Gamma_2$
<u>VALUE (units <math>10^{-4}</math>)</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b><math>1.83 \pm 0.06</math> OUR FIT</b>				
<b><math>1.46 \begin{array}{l} +0.64 \\ -0.54 \end{array} \pm 0.18</math></b>	9	<sup>1</sup> AKHMETSHIN 00F	CMD2	$e^+ e^- \rightarrow \pi^+ \pi^- \pi^+ \pi^- \geq 2\gamma$

<sup>1</sup> Using various branching ratios of  $K_S^0$ ,  $K_L^0$ ,  $\eta$ ,  $\eta'$  from the 2000 edition (The European Physical Journal **C15** 1 (2000)) of this Review.

$\Gamma(\eta'(958)\gamma)/\Gamma(\eta\gamma)$				$\Gamma_{24}/\Gamma_5$
<u>VALUE (units <math>10^{-3}</math>)</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b><math>4.77 \pm 0.15</math> OUR FIT</b>				
<b><math>4.78 \pm 0.20</math> OUR AVERAGE</b>				
$4.77 \pm 0.09 \pm 0.19$	3407	AMBROSINO 07A	KLOE	$1.02 e^+ e^- \rightarrow \pi^+ \pi^- 7\gamma$
$4.70 \pm 0.47 \pm 0.31$	120	<sup>1</sup> ALOISIO 02E	KLOE	$1.02 e^+ e^- \rightarrow \pi^+ \pi^- 3\gamma$
$6.5 \begin{array}{l} +1.7 \\ -1.5 \end{array} \pm 0.8$	21	AKHMETSHIN 00B	CMD2	$e^+ e^- \rightarrow \pi^+ \pi^- 3\gamma$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$9.5^{+5.2}_{-4.0} \pm 1.4$       6      <sup>2</sup> AKHMETSHIN 97B CMD2  $e^+ e^- \rightarrow \pi^+ \pi^- 3\gamma$

<sup>1</sup> From the decay mode  $\eta' \rightarrow \eta \pi^+ \pi^-$ ,  $\eta \rightarrow \gamma \gamma$ .

<sup>2</sup> Superseded by AKHMETSHIN 00B.

### $\Gamma(\eta\pi^0\pi^0\gamma)/\Gamma_{\text{total}}$

VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT	$\Gamma_{25}/\Gamma$
<2	90	AULCHENKO 98	SND	$e^+ e^- \rightarrow 7\gamma$	

### $\Gamma(\mu^+\mu^-\gamma)/\Gamma_{\text{total}}$

VALUE (units $10^{-5}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	$\Gamma_{26}/\Gamma$
<b><math>1.43 \pm 0.45 \pm 0.14</math></b>	27188	<sup>1</sup> AKHMETSHIN 99B	CMD2	$e^+ e^- \rightarrow \mu^+ \mu^- \gamma$	

• • • We do not use the following data for averages, fits, limits, etc. • • •

$2.3 \pm 1.0$        $824 \pm 33$       <sup>2</sup> AKHMETSHIN 97C CMD2  $e^+ e^- \rightarrow \mu^+ \mu^- \gamma$

<sup>1</sup> For  $E_\gamma > 20$  MeV. Supersedes AKHMETSHIN 97C.

<sup>2</sup> For  $E_\gamma > 20$  MeV.

### $\Gamma(\rho\gamma\gamma)/\Gamma_{\text{total}}$

VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT	$\Gamma_{27}/\Gamma$
<1.2	90	AULCHENKO 08	CMD2	$\phi \rightarrow \pi^+ \pi^- \gamma\gamma$	

• • • We do not use the following data for averages, fits, limits, etc. • • •

<5      90      AKHMETSHIN 98      CMD2  $e^+ e^- \rightarrow \pi^+ \pi^- \gamma\gamma$

### $\Gamma(\eta\pi^+\pi^-)/\Gamma_{\text{total}}$

VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT	$\Gamma_{28}/\Gamma$
< 1.8	90	AKHMETSHIN 00E	CMD2	$e^+ e^- \rightarrow \pi^+ \pi^- \pi^+ \pi^- \pi^0$	

• • • We do not use the following data for averages, fits, limits, etc. • • •

< 6.1      90      AULCHENKO 08      CMD2  $\phi \rightarrow \eta\pi^+\pi^-$

<30      90      AKHMETSHIN 98      CMD2  $e^+ e^- \rightarrow \pi^+ \pi^- \gamma\gamma$

### $\Gamma(\eta\mu^+\mu^-)/\Gamma_{\text{total}}$

VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	$\Gamma_{29}/\Gamma$
<9.4	90	AKHMETSHIN 01	CMD2	$e^+ e^- \rightarrow \eta e^+ e^-$	

### $\Gamma(\eta U \rightarrow \eta e^+ e^-)/\Gamma_{\text{total}}$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT	$\Gamma_{30}/\Gamma$
$<1 \times 10^{-6}$	90	<sup>1</sup> BABUSCI 13B	KLOE	$1.02 e^+ e^- \rightarrow \eta e^+ e^-$	

<sup>1</sup> For a narrow vector  $U$  with mass between 5 and 470 MeV, from the combined analysis of  $\eta \rightarrow \pi^+ \pi^- \pi^0$  and  $\eta \rightarrow \pi^0 \pi^0 \pi^0$  from ARCHILLI 12. Measured 90% CL limits as a function of  $m_U$  range from  $2.2 \times 10^{-8}$  to  $10^{-6}$ .

### $\Gamma(\text{invisible})/\Gamma(K^+K^-)$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT	$\Gamma_{31}/\Gamma_1$
$<3.4 \times 10^{-4}$	90	ABLIKIM 18S	BES3	$J/\psi \rightarrow \phi \eta \rightarrow \phi \pi^+ \pi^- \pi^0$	

**Lepton Family number (LF) violating modes**

$\Gamma(e^\pm \mu^\mp)/\Gamma_{\text{total}}$	$\Gamma_{32}/\Gamma$			
<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
$<2 \times 10^{-6}$	90	ACHASOV	10A SND	$e^+ e^- \rightarrow e^\pm \mu^\mp$

 **$\pi^+ \pi^- \pi^0 / \rho \pi$  AMPLITUDE RATIO  $a_1$  IN DECAY OF  $\phi \rightarrow \pi^+ \pi^- \pi^0$** 

NIECKNIG 12 describes final-state interactions between the three pions in a dispersive framework using data on the  $\pi\pi$   $P$ -wave scattering phase shift.

<u>VALUE (units <math>10^{-2}</math>)</u>	<u>CL%</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>9.1±1.2 OUR AVERAGE</b>					
10.1±4.4±1.7		80k	<sup>1</sup> AKHMETSHIN 06	CMD2	$1.017-1.021 e^+ e^- \rightarrow \pi^+ \pi^- \pi^0$
9.0±1.1±0.6		1.98M	<sup>2,3</sup> ALOISIO 03	KLOE	$1.02 e^+ e^- \rightarrow \pi^+ \pi^- \pi^0$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$-6 < a_1 < 6$		500k	<sup>3</sup> ACHASOV 02	SND	$e^+ e^- \rightarrow \pi^+ \pi^- \pi^0$
$-16 < a_1 < 11$	90	9.8k	<sup>1,4</sup> AKHMETSHIN 98	CMD2	$e^+ e^- \rightarrow \pi^+ \pi^- \gamma\gamma$

<sup>1</sup> Dalitz plot analysis taking into account interference between the contact and  $\rho\pi$  amplitudes.

<sup>2</sup> From a fit without limitations on charged and neutral  $\rho$  masses and widths.

<sup>3</sup> Recalculated by us to match the notations of AKHMETSHIN 98.

<sup>4</sup> Assuming zero phase for the contact term.

**PARAMETER  $\beta$  IN  $\phi \rightarrow Pe^+ e^-$  DECAYS**

In the one-pole approximation the electromagnetic transition form factor for  $\phi \rightarrow Pe^+ e^-$  ( $P = \pi, \eta$ ) is given as a function of the  $e^+ e^-$  invariant mass squared,  $q^2$ , by the expression:

$$|F(q^2)|^2 = (1 - q^2/\Lambda^2)^{-2},$$

where vector meson dominance predicts parameter  $\Lambda \approx 0.770$  GeV ( $\Lambda^{-2} \approx 1.687$  GeV $^{-2}$ ). The slope of this form factor,  $\beta = dF/dq^2(q^2=0)$ , equals  $\Lambda^{-2}$  in this approximation.

The measurements below obtain  $\beta$  in the one-pole approximation.

**PARAMETER  $\beta$  IN  $\phi \rightarrow \pi^0 e^+ e^-$  DECAY**

<u>VALUE (GeV<math>^{-2}</math>)</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>2.02±0.11</b>	9.5k	<sup>1</sup> ANASTASI 16B	KLOE	$1.02 e^+ e^- \rightarrow \pi^0 e^+ e^-$

<sup>1</sup> The error combines statistical and systematic uncertainties.

**PARAMETER  $\beta$  IN  $\phi \rightarrow \eta e^+ e^-$  DECAY**

<u>VALUE (GeV<math>^{-2}</math>)</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>1.29±0.13 OUR AVERAGE</b>				

$1.28 \pm 0.10$	$+0.09$	30k	BABUSCI 15	KLOE	$1.02 e^+ e^- \rightarrow \eta e^+ e^-$
3.8	$\pm 1.8$	213	<sup>1</sup> ACHASOV 01B	SND	$1.02 e^+ e^- \rightarrow \eta e^+ e^-$

<sup>1</sup> The uncertainty is statistical only. The systematic one is negligible, in comparison.

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AKHMETSHIN	00B	PL B473 337	R.R. Akhmetshin <i>et al.</i>	(Novosibirsk CMD-2 Collab.)
AKHMETSHIN	00E	PL B491 81	R.R. Akhmetshin <i>et al.</i>	(Novosibirsk CMD-2 Collab.)
AKHMETSHIN	00F	PL B494 26	R.R. Akhmetshin <i>et al.</i>	(Novosibirsk CMD-2 Collab.)
AULCHENKO	00A	JETP 90 927	V.M. Aulchenko <i>et al.</i>	(Novosibirsk SND Collab.)
		Translated from ZETF 117 1067.		
BRAMON	00	PL B486 406	A. Bramon <i>et al.</i>	
PDG	00	EPJ C15 1	D.E. Groom <i>et al.</i>	(PDG Collab.)
ACHASOV	99	PL B449 122	M.N. Achasov <i>et al.</i>	
ACHASOV	99C	PL B456 304	M.N. Achasov <i>et al.</i>	
AKHMETSHIN	99B	PL B462 371	R.R. Akhmetshin <i>et al.</i>	(Novosibirsk CMD-2 Collab.)
AKHMETSHIN	99C	PL B462 380	R.R. Akhmetshin <i>et al.</i>	(Novosibirsk CMD-2 Collab.)
AKHMETSHIN	99D	PL B466 385	R.R. Akhmetshin <i>et al.</i>	
Also		PL B508 217 (errat.)	R.R. Akhmetshin <i>et al.</i>	(Novosibirsk CMD-2 Collab.)
AKHMETSHIN	99F	PL B460 242	R.R. Akhmetshin <i>et al.</i>	(Novosibirsk CMD-2 Collab.)
AULCHENKO	99	JETP 69 97	V.M. Aulchenko <i>et al.</i>	
		Translated from ZETFP 69 87.		
ACHASOV	98B	PL B438 441	M.N. Achasov <i>et al.</i>	(Novosibirsk SND Collab.)
ACHASOV	98F	JETPL 68 573	M.N. Achasov <i>et al.</i>	(Novosibirsk SND Collab.)
ACHASOV	98I	PL B440 442	M.N. Achasov <i>et al.</i>	
AKHMETSHIN	98	PL B434 426	R.R. Akhmetshin <i>et al.</i>	(CMD-2 Collab.)
AULCHENKO	98	PL B436 199	V.M. Aulchenko <i>et al.</i>	(Novosibirsk SND Collab.)
BARBERIS	98	PL B432 436	D. Barberis <i>et al.</i>	(Omega Expt.)
AKHMETSHIN	97B	PL B415 445	R.R. Akhmetshin <i>et al.</i>	(NOVO, BOST, PIT+)
AKHMETSHIN	97C	PL B415 452	R.R. Akhmetshin <i>et al.</i>	(Novosibirsk CMD-2 Collab.)
BENAYOUN	96	ZPHY C72 221	M. Benayoun <i>et al.</i>	(IPNP, NOVO)
AKHMETSHIN	95	PL B364 199	R.R. Akhmetshin <i>et al.</i>	(Novosibirsk CMD-2 Collab.)
DOLINSKY	91	PRPL 202 99	S.I. Dolinsky <i>et al.</i>	(NOVO)
KUHN	90	ZPHY C48 445	J.H. Kuhn <i>et al.</i>	(MPIM)
ACHASOV	89	NP B315 465	N.N. Achasov, V.N. Ivanchenko	
DOLINSKY	89	ZPHY C42 511	S.I. Dolinsky <i>et al.</i>	(NOVO)
BARKOV	88	SJNP 47 248	L.M. Barkov <i>et al.</i>	(NOVO)
		Translated from YAF 47 393.		
DOLINSKY	88	SJNP 48 277	S.I. Dolinsky <i>et al.</i>	(NOVO)
		Translated from YAF 48 442.		
DRUZHININ	87	ZPHY C37 1	V.P. Druzhinin <i>et al.</i>	(NOVO)
ARMSTRONG	86	PL 166B 245	T.A. Armstrong <i>et al.</i>	(ATHU, BARI, BIRM+)
ATKINSON	86	ZPHY C30 521	M. Atkinson <i>et al.</i>	(BONN, CERN, GLAS+)
BEBEK	86	PRL 56 1893	C. Bebek <i>et al.</i>	(CLEO Collab.)
DAVENPORT	86	PR D33 2519	T.F. Davenport (TUFTS, ARIZ, FNAL, FSU, NDAM+)	
DIJKSTRA	86	ZPHY C31 375	H. Dijkstra <i>et al.</i>	(ANIK, BRIS, CERN+)
FRAME	86	NP B276 667	D. Frame <i>et al.</i>	(GLAS)
GOLUBEV	86	SJNP 44 409	V.B. Golubev <i>et al.</i>	(NOVO)
		Translated from YAF 44 633.		
ALBRECHT	85D	PL 153B 343	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
GOLUBEV	85	SJNP 41 756	V.B. Golubev <i>et al.</i>	(NOVO)
		Translated from YAF 41 1183.		
DRUZHININ	84	PL 144B 136	V.P. Druzhinin <i>et al.</i>	(NOVO)
ARMSTRONG	83B	NP B224 193	T.A. Armstrong <i>et al.</i>	(BARI, BIRM, CERN+)
BARATE	83	PL 121B 449	R. Barate <i>et al.</i>	(SACL, LOIC, SHMP, IND)
KURDADZE	83C	JETPL 38 366	L.M. Kurdadze <i>et al.</i>	(NOVO)
		Translated from ZETFP 38 306.		
ARENTON	82	PR D25 2241	M.W. Arenton <i>et al.</i>	(ANL, ILL)
PELLINEN	82	PS 25 599	A. Pellinen, M. Roos	(HELS)
DAUM	81	PL 100B 439	C. Daum <i>et al.</i>	(AMST, BRIS, CERN, CRAC+)
IVANOV	81	PL 107B 297	P.M. Ivanov <i>et al.</i>	(NOVO)
Also		Private Comm.	S.I. Eidelman	(NOVO)
VASSERMAN	81	PL 99B 62	I.B. Vasserman <i>et al.</i>	(NOVO)
Also		SJNP 35 240	L.M. Kurdadze <i>et al.</i>	(NOVO)
		Translated from YAF 35 352.		
CORDIER	80	NP B172 13	A. Cordier <i>et al.</i>	(LALO)
CORDIER	79	PL 81B 389	A. Cordier <i>et al.</i>	(LALO)
BUKIN	78B	SJNP 27 521	A.D. Bokin <i>et al.</i>	(NOVO)
		Translated from YAF 27 985.		
BUKIN	78C	SJNP 27 516	A.D. Bokin <i>et al.</i>	(NOVO)
		Translated from YAF 27 976.		

COOPER	78B	NP B146 1	A.M. Cooper <i>et al.</i>	(TATA, CERN, CDEF+)
LOSTY	78	NP B133 38	M.J. Losty <i>et al.</i>	(CERN, AMST, NIJM+)
AKERLOF	77	PRL 39 861	C.W. Akerlof <i>et al.</i>	(FNAL, MICH, PURD)
ANDREWS	77	PRL 38 198	D.E. Andrews <i>et al.</i>	(ROCH)
BALDI	77	PL 68B 381	R. Baldi <i>et al.</i>	(GEVA)
CERRADA	77B	NP B126 241	M. Cerrada <i>et al.</i>	(AMST, CERN, NIJM+)
COHEN	77	PRL 38 269	D. Cohen <i>et al.</i>	(ANL)
LAVEN	77	NP B127 43	H. Laven <i>et al.</i>	(AACH3, BERL, CERN, LOIC+)
LYONS	77	NP B125 207	L. Lyons, A.M. Cooper, A.G. Clark	(OXF)
COSME	76	PL 63B 352	G. Cosme <i>et al.</i>	(ORSAY)
KALBFLEISCH	76	PR D13 22	G.R. Kalbfleisch, R.C. Strand, J.W. Chapman	(BNL+)
PARROUR	76	PL 63B 357	G. Parrou <i>et al.</i>	(ORSAY)
PARROUR	76B	PL 63B 362	G. Parrou <i>et al.</i>	(ORSAY)
KALBFLEISCH	75	PR D11 987	G.R. Kalbfleisch, R.C. Strand, J.W. Chapman	(BNL+)
AYRES	74	PRL 32 1463	D.S. Ayres <i>et al.</i>	(ANL)
BESCH	74	NP B70 257	H.J. Besch <i>et al.</i>	(BONN)
COSME	74	PL 48B 155	G. Cosme <i>et al.</i>	(ORSAY)
COSME	74B	PL 48B 159	G. Cosme <i>et al.</i>	(ORSAY)
DEGROOT	74	NP B74 77	A.J. de Groot <i>et al.</i>	(AMST, NIJM)
AUGUSTIN	73	PRL 30 462	J.E. Augustin <i>et al.</i>	(ORSAY)
BALLAM	73	PR D7 3150	J. Ballam <i>et al.</i>	(SLAC, LBL)
BINNIE	73B	PR D8 2789	D.M. Binnie <i>et al.</i>	(LOIC, SHMP)
AGUILAR-...	72B	PR D6 29	M. Aguilar-Benitez <i>et al.</i>	(BNL)
ALVENSLEB...	72	PRL 28 66	H. Alvensleben <i>et al.</i>	(MIT, DESY)
BORENSTEIN	72	PR D5 1559	S.R. Borenstein <i>et al.</i>	(BNL, MICH)
COLLEY	72	NP B50 1	D.C. Colley <i>et al.</i>	(BIRM, GLAS)
BALAKIN	71	PL 34B 328	V.E. Balakin <i>et al.</i>	(NOVO)
CHATELUS	71	Thesis LAL 1247	Y. Chatelus	(STRB)
Also		PL 32B 416	J.C. Bizot <i>et al.</i>	(ORSAY)
HAYES	71	PR D4 899	S. Hayes <i>et al.</i>	(CORN)
STOTTLE...	71	Thesis ORO 2504 170	A.R. Stottlemeyer	(UMD)
BIZOT	70	PL 32B 416	J.C. Bizot <i>et al.</i>	(ORSAY)
Also		Liverpool Sym. 69	J.P. Perez-y-Jorba	
EARLES	70	PRL 25 1312	D.R. Earles <i>et al.</i>	(NEAS)
LINDSEY	66	PR 147 913	J.S. Lindsey, G. Smith	(LRL)
LONDON	66	PR 143 1034	G.W. London <i>et al.</i>	(BNL, SYRA) IGJPC
BADIER	65B	PL 17 337	J. Badier <i>et al.</i>	(EPOL, SACL, AMST)
LINDSEY	65	PRL 15 221	J.S. Lindsey, G.A. Smith	(LRL)
LINDSEY	65	data included in LINDSEY 66.		
SCHLEIN	63	PRL 10 368	P.E. Schlein <i>et al.</i>	(UCLA) IGJP