

**$a_1(1260)$**  $I^G(J^{PC}) = 1^-(1^{++})$ 

See also our review under the  $a_1(1260)$  in PDG 06, Journal of Physics **G33** 1 (2006).

 **$a_1(1260)$  T-MATRIX POLE  $\sqrt{s}$** 

Note that  $\Gamma = -2 \operatorname{Im}(\sqrt{s})$ .

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>(1209<math>^{+13}_{-10}</math>)<math>-i(288^{+45}_{-12})</math> OUR ESTIMATE</b>			
$(1209 \pm 4^{+12}_{-9}) - i(288 \pm 6^{+45}_{-10})$	MIKHASENKO 18	RVUE	$\tau^- \rightarrow \pi^- \pi^+ \pi^- \nu_\tau$

 **$a_1(1260)$  MASS**

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1230 <math>\pm 40</math></b>	<b>OUR ESTIMATE</b>			
<b>1299 <math>\pm 12</math></b>	46M	<sup>1</sup> AGHASYAN	18B COMP	$190 \pi^- p \rightarrow \pi^- \pi^+ \pi^- p$
<b>• • • We do not use the following data for averages, fits, limits, etc. • • •</b>				
1195.05 $\pm 1.05 \pm 6.33$	894k	AAIJ	18AI LHCb	$D^0 \rightarrow K^\mp \pi^\pm \pi^\pm \pi^\mp$
1225 $\pm 9 \pm 20$	7k	<sup>2</sup> DARGENT	17 RVUE	$D^0 \rightarrow \pi^- \pi^+ \pi^- \pi^+$
1255 $\pm 6 \pm 7$	420k	<sup>3</sup> ALEKSEEV	10 COMP	$190 \pi^- Pb \rightarrow \pi^- \pi^- \pi^+ Pb'$
1243 $\pm 12 \pm 20$		<sup>4</sup> AUBERT	07AU BABR	$10.6 e^+ e^- \rightarrow \rho^0 \rho^\pm \pi^\mp \gamma$
1230–1270	6360	<sup>5</sup> LINK	07A FOCS	$D^0 \rightarrow \pi^- \pi^+ \pi^- \pi^+$
1203 $\pm 3$		<sup>6</sup> GOMEZ-DUM..04	RVUE	$\tau^+ \rightarrow \pi^+ \pi^+ \pi^- \nu_\tau$
1330 $\pm 24$	90k	SALVINI	04 OBLX	$\bar{p}p \rightarrow 2\pi^+ 2\pi^-$
1331 $\pm 10 \pm 3$	37k	<sup>7</sup> ASNER	00 CLE2	$10.6 e^+ e^- \rightarrow \tau^+ \tau^-, \tau^- \rightarrow \pi^- \pi^0 \pi^0 \nu_\tau$
1255 $\pm 7 \pm 6$	5904	<sup>8</sup> ABREU	98G DLPH	$e^+ e^-$
1207 $\pm 5 \pm 8$	5904	<sup>9</sup> ABREU	98G DLPH	$e^+ e^-$
1196 $\pm 4 \pm 5$	5904	<sup>10,11</sup> ABREU	98G DLPH	$e^+ e^-$
1240 $\pm 10$		BARBERIS	98B	$450 pp \rightarrow p_f \pi^+ \pi^- \pi^0 p_s$
1262 $\pm 9 \pm 7$		<sup>8,12</sup> ACKERSTAFF	97R OPAL	$E_{cm}^{ee} = 88–94, \tau \rightarrow 3\pi\nu$
1210 $\pm 7 \pm 2$		<sup>9,12</sup> ACKERSTAFF	97R OPAL	$E_{cm}^{ee} = 88–94, \tau \rightarrow 3\pi\nu$
1211 $\pm 7 \pm 50$		<sup>9</sup> ALBRECHT	93C ARG	$\tau^+ \rightarrow \pi^+ \pi^+ \pi^- \nu$
1121 $\pm 8$		<sup>13</sup> ANDO	92 SPEC	$8 \pi^- p \rightarrow \pi^+ \pi^- \pi^0 n$
1242 $\pm 37$		<sup>14</sup> IVANOV	91 RVUE	$\tau \rightarrow \pi^+ \pi^+ \pi^- \nu$
1260 $\pm 14$		<sup>15</sup> IVANOV	91 RVUE	$\tau \rightarrow \pi^+ \pi^+ \pi^- \nu$
1250 $\pm 9$		<sup>16</sup> IVANOV	91 RVUE	$\tau \rightarrow \pi^+ \pi^+ \pi^- \nu$
1208 $\pm 15$		ARMSTRONG	90 OMEG	$300.0 pp \rightarrow pp \pi^+ \pi^- \pi^0$

1220	$\pm 15$	<sup>17</sup> ISGUR	89	RVUE	$\tau^+ \rightarrow \pi^+ \pi^+ \pi^- \nu$
1260	$\pm 25$	<sup>18</sup> BOWLER	88	RVUE	
1166	$\pm 18$	BAND	87	MAC	$\tau^+ \rightarrow \pi^+ \pi^+ \pi^- \nu$
1164	$\pm 41$	BAND	87	MAC	$\tau^+ \rightarrow \pi^+ \pi^0 \pi^0 \nu$
1250	$\pm 40$	<sup>17</sup> TORNQVIST	87	RVUE	
1046	$\pm 11$	ALBRECHT	86B	ARG	$\tau^+ \rightarrow \pi^+ \pi^+ \pi^- \nu$
1056	$\pm 20$	RUCKSTUHL	86	DLCO	$\tau^+ \rightarrow \pi^+ \pi^+ \pi^- \nu$
1194	$\pm 14$	SCHMIDKE	86	MRK2	$\tau^+ \rightarrow \pi^+ \pi^+ \pi^- \nu$
1255	$\pm 23$	BELLINI	85	SPEC	$40 \pi^- A \rightarrow \pi^- \pi^+ \pi^- A$
1240	$\pm 80$	<sup>19</sup> DANKOWY...	81	SPEC	$8.45 \pi^- p \rightarrow n 3\pi$
1280	$\pm 30$	<sup>19</sup> DAUM	81B	CNTR	$63,94 \pi^- p \rightarrow p 3\pi$
1041	$\pm 13$	<sup>20</sup> GAVILLET	77	HBC	$4.2 K^- p \rightarrow \Sigma 3\pi$

<sup>1</sup> Statistical error negligible.<sup>2</sup> Reanalysis of CLEO data using Breit-Wigner parameterization.<sup>3</sup> Superseded by AGHASYAN 2018B.<sup>4</sup> The  $\rho^\pm \pi^\mp$  state can be also due to the  $\pi(1300)$ .<sup>5</sup> Using the Breit-Wigner parameterization; strong correlation between mass and width.<sup>6</sup> Using the data of BARATE 98R.<sup>7</sup> From a fit to the  $3\pi$  mass spectrum including the  $K\bar{K}^*(892)$  threshold.<sup>8</sup> Uses the model of KUHN 90.<sup>9</sup> Uses the model of ISGUR 89.<sup>10</sup> Includes the effect of a possible  $a'_1$  state.<sup>11</sup> Uses the model of FEINDT 90.<sup>12</sup> Supersedes AKERS 95P.<sup>13</sup> Average and spread of values using 2 variants of the model of BOWLER 75.<sup>14</sup> Reanalysis of RUCKSTUHL 86.<sup>15</sup> Reanalysis of SCHMIDKE 86.<sup>16</sup> Reanalysis of ALBRECHT 86B.<sup>17</sup> From a combined reanalysis of ALBRECHT 86B, SCHMIDKE 86, and RUCKSTUHL 86.<sup>18</sup> From a combined reanalysis of ALBRECHT 86B and DAUM 81B.<sup>19</sup> Uses the model of BOWLER 75.<sup>20</sup> Produced in  $K^-$  backward scattering.

## **$a_1(1260)$ WIDTH**

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>250 to 600 OUR ESTIMATE</b>				
<b>380 <math>\pm 80</math></b>	46M	<sup>1</sup> AGHASYAN	18B COMP	$190 \pi^- p \rightarrow \pi^- \pi^+ \pi^- p$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
422.01 $\pm$ 2.10 $\pm$ 12.72 894k		AAIJ	18AI LHCb	$D^0 \rightarrow K^\mp \pi^\pm \pi^\pm \pi^\mp$
430 $\pm$ 24 $\pm$ 31		<sup>2</sup> DARGENT	17 RVUE	$D^0 \rightarrow \pi^- \pi^+ \pi^- \pi^+$
367 $\pm$ 9 $\pm$ 28 — 25 420k		<sup>3</sup> ALEKSEEV	10 COMP	$190 \pi^- Pb \rightarrow \pi^- \pi^- \pi^+ Pb'$
410 $\pm$ 31 $\pm$ 30		<sup>4</sup> AUBERT	07AU BABR	$10.6 e^+ e^- \rightarrow \rho^0 \rho^\pm \pi^\mp \gamma$
520–680	6360	<sup>5</sup> LINK	07A FOCS	$D^0 \rightarrow \pi^- \pi^+ \pi^- \pi^+$
480 $\pm$ 20		<sup>6</sup> GOMEZ-DUM..04	RVUE	$\tau^+ \rightarrow \pi^+ \pi^+ \pi^- \nu_\tau$
580 $\pm$ 41	90k	SALVINI	04 OBLX	$\bar{p}p \rightarrow 2\pi^+ 2\pi^-$
460 $\pm$ 85	205	<sup>7</sup> DRUTSKOY	02 BELL	$B \rightarrow D^{(*)} K^- K^{*0}$

814	$\pm 36$	$\pm 13$	37k	<sup>8</sup> ASNER	00	CLE2	$10.6 e^+ e^- \rightarrow \tau^+ \tau^-, \tau^- \rightarrow \pi^- \pi^0 \pi^0 \nu_\tau$
450	$\pm 50$		22k	<sup>9</sup> AKHMETSHIN 99E	CMD2	$1.05-1.38 e^+ e^- \rightarrow \pi^+ \pi^- \pi^0 \pi^0$	
570	$\pm 10$			<sup>10</sup> BONDAR	99	RVUE	$e^+ e^- \rightarrow 4\pi, \tau \rightarrow 3\pi \nu_\tau$
587	$\pm 27$	$\pm 21$	5904	<sup>11</sup> ABREU	98G	DLPH	$e^+ e^-$
478	$\pm 3$	$\pm 15$	5904	<sup>12</sup> ABREU	98G	DLPH	$e^+ e^-$
425	$\pm 14$	$\pm 8$	5904	<sup>13,14</sup> ABREU	98G	DLPH	$e^+ e^-$
400	$\pm 35$			BARBERIS	98B		$450 pp \rightarrow p_f \pi^+ \pi^- \pi^0 p_s$
621	$\pm 32$	$\pm 58$		<sup>11,15</sup> ACKERSTAFF	97R	OPAL	$E_{cm}^{ee} = 88-94, \tau \rightarrow 3\pi \nu$
457	$\pm 15$	$\pm 17$		<sup>12,15</sup> ACKERSTAFF	97R	OPAL	$E_{cm}^{ee} = 88-94, \tau \rightarrow 3\pi \nu$
446	$\pm 21$	$\pm 140$		<sup>12</sup> ALBRECHT	93C	ARG	$\tau^+ \rightarrow \pi^+ \pi^+ \pi^- \nu$
239	$\pm 11$			ANDO	92	SPEC	$8 \pi^- p \rightarrow \pi^+ \pi^- \pi^0 n$
266	$\pm 13$	$\pm 4$		<sup>16</sup> ANDO	92	SPEC	$8 \pi^- p \rightarrow \pi^+ \pi^- \pi^0 n$
465	$\pm 228$			<sup>17</sup> IVANOV	91	RVUE	$\tau \rightarrow \pi^+ \pi^+ \pi^- \nu$
298	$\pm 40$			<sup>18</sup> IVANOV	91	RVUE	$\tau \rightarrow \pi^+ \pi^+ \pi^- \nu$
488	$\pm 32$			<sup>19</sup> IVANOV	91	RVUE	$\tau \rightarrow \pi^+ \pi^+ \pi^- \nu$
430	$\pm 50$			ARMSTRONG	90	OMEG	$300.0 pp \rightarrow pp \pi^+ \pi^- \pi^0$
420	$\pm 40$			<sup>20</sup> ISGUR	89	RVUE	$\tau^+ \rightarrow \pi^+ \pi^+ \pi^- \nu$
396	$\pm 43$			<sup>21</sup> BOWLER	88	RVUE	
405	$\pm 75$	$\pm 25$		BAND	87	MAC	$\tau^+ \rightarrow \pi^+ \pi^+ \pi^- \nu$
419	$\pm 108$	$\pm 57$		BAND	87	MAC	$\tau^+ \rightarrow \pi^+ \pi^0 \pi^0 \nu$
521	$\pm 27$			ALBRECHT	86B	ARG	$\tau^+ \rightarrow \pi^+ \pi^+ \pi^- \nu$
476	$\pm 132$	$\pm 54$		RUCKSTUHL	86	DLCO	$\tau^+ \rightarrow \pi^+ \pi^+ \pi^- \nu$
462	$\pm 56$	$\pm 30$		SCHMIDKE	86	MRK2	$\tau^+ \rightarrow \pi^+ \pi^+ \pi^- \nu$
292	$\pm 40$			BELLINI	85	SPEC	$40 \pi^- A \rightarrow \pi^- \pi^+ \pi^- A$
380	$\pm 100$			<sup>22</sup> DANKOWY...	81	SPEC	$8.45 \pi^- p \rightarrow n 3\pi$
300	$\pm 50$			<sup>22</sup> DAUM	81B	CNTR	$63,94 \pi^- p \rightarrow p 3\pi$
230	$\pm 50$			<sup>23</sup> GAVILLET	77	HBC	$4.2 K^- p \rightarrow \Sigma 3\pi$

<sup>1</sup> Statistical error negligible.<sup>2</sup> Reanalysis of CLEO data using Breit-Wigner parameterization.<sup>3</sup> Superseded by AGHASYAN 2018B.<sup>4</sup> The  $\rho^\pm \pi^\mp$  state can be also due to the  $\pi(1300)$ .<sup>5</sup> Using the Breit-Wigner parameterization; strong correlation between mass and width.<sup>6</sup> Using the data of BARATE 98R.<sup>7</sup> From a fit of the  $K^- K^{*0}$  distribution assuming  $m_{a_1} = 1230$  MeV and purely resonant production of the  $K^- K^{*0}$  system.<sup>8</sup> From a fit to the  $3\pi$  mass spectrum including the  $K\bar{K}^*(892)$  threshold.<sup>9</sup> Using the  $a_1(1260)$  mass of 1230 MeV.<sup>10</sup> From AKHMETSHIN 99E and ASNER 00 data using the  $a_1(1260)$  mass of 1230 MeV.

- <sup>11</sup> Uses the model of KUHN 90.  
<sup>12</sup> Uses the model of ISGUR 89.  
<sup>13</sup> Includes the effect of a possible  $a'_1$  state.  
<sup>14</sup> Uses the model of FEINDT 90.  
<sup>15</sup> Supersedes AKERS 95P.  
<sup>16</sup> Average and spread of values using 2 variants of the model of BOWLER 75.  
<sup>17</sup> Reanalysis of RUCKSTUHL 86.  
<sup>18</sup> Reanalysis of SCHMIDKE 86.  
<sup>19</sup> Reanalysis of ALBRECHT 86B.  
<sup>20</sup> From a combined reanalysis of ALBRECHT 86B, SCHMIDKE 86, and RUCKSTUHL 86.  
<sup>21</sup> From a combined reanalysis of ALBRECHT 86B and DAUM 81B.  
<sup>22</sup> Uses the model of BOWLER 75.  
<sup>23</sup> Produced in  $K^-$  backward scattering.
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## $a_1(1260)$ DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $3\pi$	seen
$\Gamma_2$ $(\rho\pi)_{S\text{-wave}}, \rho \rightarrow \pi\pi$	seen
$\Gamma_3$ $(\rho\pi)_{D\text{-wave}}, \rho \rightarrow \pi\pi$	seen
$\Gamma_4$ $(\rho(1450)\pi)_{S\text{-wave}}, \rho \rightarrow \pi\pi$	seen
$\Gamma_5$ $(\rho(1450)\pi)_{D\text{-wave}}, \rho \rightarrow \pi\pi$	seen
$\Gamma_6$ $f_0(500)\pi, f_0 \rightarrow \pi\pi$	seen
$\Gamma_7$ $f_0(980)\pi, f_0 \rightarrow \pi\pi$	seen
$\Gamma_8$ $f_0(1370)\pi, f_0 \rightarrow \pi\pi$	seen
$\Gamma_9$ $f_2(1270)\pi, f_2 \rightarrow \pi\pi$	seen
$\Gamma_{10}$ $\pi^+\pi^-\pi^0$	seen
$\Gamma_{11}$ $\pi^0\pi^0\pi^0$	not seen
$\Gamma_{12}$ $KK\pi$	seen
$\Gamma_{13}$ $K^*(892)K$	seen
$\Gamma_{14}$ $\pi\gamma$	seen

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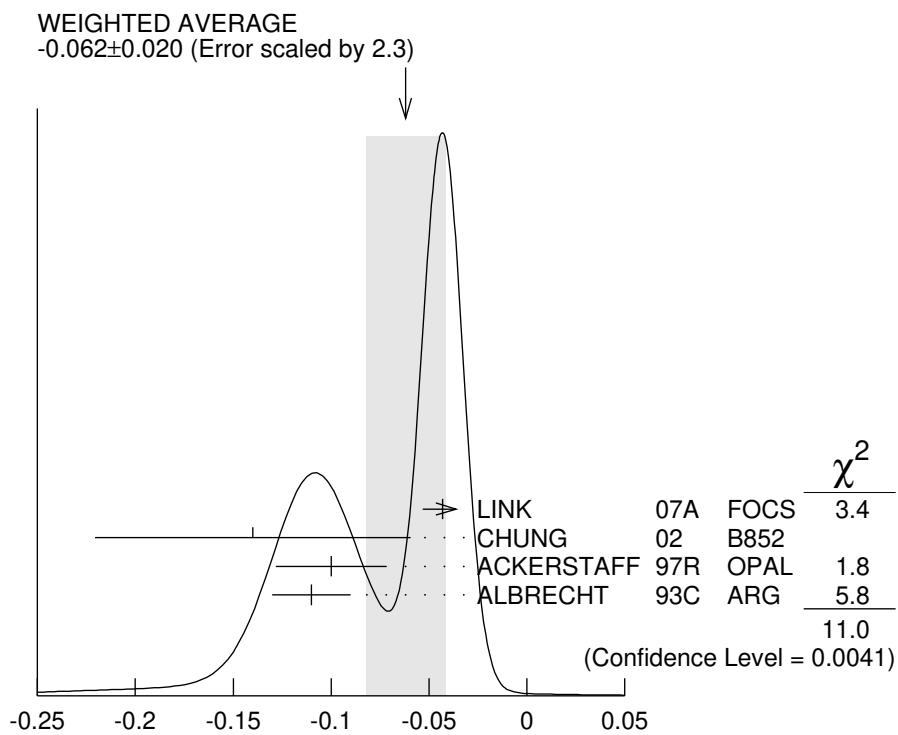
## $a_1(1260)$ PARTIAL WIDTHS

$\Gamma(\pi\gamma)$	$\Gamma_{14}$
VALUE (keV) <b>640±246</b>	DOCUMENT ID ZIELINSKI TECN SPEC COMMENT 200 $\pi^+Z \rightarrow Z3\pi$

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## $D\text{-wave}/S\text{-wave}$ AMPLITUDE RATIO IN DECAY OF $a_1(1260) \rightarrow \rho\pi$

VALUE	DOCUMENT ID	TECN	COMMENT
<b>-0.062±0.020 OUR AVERAGE</b>	Error includes scale factor of 2.3. See the ideogram below.		
-0.043±0.009±0.005	LINK	07A FOCS	$D^0 \rightarrow \pi^-\pi^+\pi^-\pi^+$
-0.14 ± 0.04 ± 0.07	<sup>1</sup> CHUNG	02 B852	$18.3 \pi^- p \rightarrow \pi^+\pi^-\pi^- p$
-0.10 ± 0.02 ± 0.02	2,3 ACKERSTAFF	97R OPAL	$E_{cm}^{ee} = 88\text{--}94, \tau \rightarrow 3\pi\nu$
-0.11 ± 0.02	2 ALBRECHT	93C ARG	$\tau^+ \rightarrow \pi^+\pi^+\pi^-\nu$

<sup>1</sup> Deck-type background not subtracted.<sup>2</sup> Uses the model of ISGUR 89.<sup>3</sup> Supersedes AKERS 95P.*D-wave/S-wave AMPLITUDE RATIO IN DECAY OF  $a_1(1260) \rightarrow \rho\pi$*  **$a_1(1260)$  BRANCHING RATIOS** **$\Gamma((\rho\pi)_{S\text{-wave}}, \rho \rightarrow \pi\pi)/\Gamma_{\text{total}}$**  **$\Gamma_2/\Gamma$** 

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>• • • We do not use the following data for averages, fits, limits, etc. • • •</b>				
60.19	37k	<sup>1</sup> ASNER	00	CLE2 $e^+e^- \rightarrow \tau^+\tau^-$ , $\tau^- \rightarrow \pi^-\pi^0\pi^0\nu_\tau$

<sup>1</sup> From a fit to the Dalitz plot. **$\Gamma((\rho\pi)_{D\text{-wave}}, \rho \rightarrow \pi\pi)/\Gamma_{\text{total}}$**  **$\Gamma_3/\Gamma$** 

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>• • • We do not use the following data for averages, fits, limits, etc. • • •</b>				
$1.30 \pm 0.60 \pm 0.22$	37k	<sup>1</sup> ASNER	00	CLE2 $e^+e^- \rightarrow \tau^+\tau^-$ , $\tau^- \rightarrow \pi^-\pi^0\pi^0\nu_\tau$

<sup>1</sup> From a fit to the Dalitz plot. **$\Gamma((\rho(1450)\pi)_{S\text{-wave}}, \rho \rightarrow \pi\pi)/\Gamma_{\text{total}}$**  **$\Gamma_4/\Gamma$** 

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>• • • We do not use the following data for averages, fits, limits, etc. • • •</b>				
$0.56 \pm 0.84 \pm 0.32$	37k	<sup>1,2</sup> ASNER	00	CLE2 $e^+e^- \rightarrow \tau^+\tau^-$ , $\tau^- \rightarrow \pi^-\pi^0\pi^0\nu_\tau$

<sup>1</sup> From a fit to the Dalitz plot.<sup>2</sup> Assuming for  $\rho(1450)$  mass and width of 1370 and 386 MeV respectively.

$\Gamma((\rho(1450)\pi)_D\text{-wave}, \rho \rightarrow \pi\pi)/\Gamma_{\text{total}}$   $\Gamma_5/\Gamma$

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>• • •</b> We do not use the following data for averages, fits, limits, etc. <b>• • •</b>				
$2.04 \pm 1.20 \pm 0.28$	37k	1,2 ASNER	00	CLE2 $10.6 e^+ e^- \rightarrow \tau^+ \tau^-, \tau^- \rightarrow \pi^- \pi^0 \pi^0 \nu_\tau$

<sup>1</sup> From a fit to the Dalitz plot.<sup>2</sup> Assuming for  $\rho(1450)$  mass and width of 1370 and 386 MeV respectively.

$\Gamma(f_0(500)\pi, f_0 \rightarrow \pi\pi)/\Gamma_{\text{total}}$   $\Gamma_6/\Gamma$

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>• • •</b> We do not use the following data for averages, fits, limits, etc. <b>• • •</b>				
seen		CHUNG	02	B852 $18.3 \pi^- p \rightarrow \pi^+ \pi^- \pi^- p$
$18.76 \pm 4.29 \pm 1.48$	37k	1,2 ASNER	00	CLE2 $10.6 e^+ e^- \rightarrow \tau^+ \tau^-, \tau^- \rightarrow \pi^- \pi^0 \pi^0 \nu_\tau$

<sup>1</sup> From a fit to the Dalitz plot.<sup>2</sup> Assuming for  $f_0(500)$  ( $\sigma$ ) mass and width of 860 and 880 MeV respectively.

$\Gamma(f_0(500)\pi, f_0 \rightarrow \pi\pi)/[\Gamma((\rho\pi)_S\text{-wave}, \rho \rightarrow \pi\pi) + \Gamma((\rho\pi)_S\text{-wave}, \rho \rightarrow \pi\pi)]$   $\Gamma_6/(\Gamma_2+\Gamma_2)$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>• • •</b> We do not use the following data for averages, fits, limits, etc. <b>• • •</b>				
$0.06 \pm 0.05$	90k	SALVINI	04	OBLX $\bar{p}p \rightarrow 2\pi^+ 2\pi^-$
$\sim 0.3$	28k	AKHMETSHIN	99E	CMD2 $1.05-1.38 e^+ e^- \rightarrow \pi^+ \pi^- \pi^+ \pi^-$
$0.003 \pm 0.003$		1 LONGACRE	82	RVUE

<sup>1</sup> Uses multichannel Aitchison-Bowler model (BOWLER 75). Uses data from GAVILET 77, DAUM 80, and DANKOWYCH 81.

$\Gamma(f_0(980)\pi, f_0 \rightarrow \pi\pi)/\Gamma_{\text{total}}$   $\Gamma_7/\Gamma$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
seen		1 ALEXEEV	21	COMP $\pi^- p \rightarrow \pi^- \pi^+ \pi^- p$
<b>• • •</b> We do not use the following data for averages, fits, limits, etc. <b>• • •</b>				
not seen	37k	ASNER	00	CLE2 $10.6 e^+ e^- \rightarrow \tau^+ \tau^-, \tau^- \rightarrow \pi^- \pi^0 \pi^0 \nu_\tau$

<sup>1</sup> The  $a_1(1260)^- \rightarrow f_0(980)\pi^-$  decay mode via the Triangle Singularity mechanism from MIKHASENKO 15 and ACETI 16 explains the  $a_1(1420)^-$  signal observed by ADOLPH 15C.

$\Gamma(f_0(1370)\pi, f_0 \rightarrow \pi\pi)/\Gamma_{\text{total}}$   $\Gamma_8/\Gamma$

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>• • •</b> We do not use the following data for averages, fits, limits, etc. <b>• • •</b>				
$7.40 \pm 2.71 \pm 1.26$	37k	1,2 ASNER	00	CLE2 $10.6 e^+ e^- \rightarrow \tau^+ \tau^-, \tau^- \rightarrow \pi^- \pi^0 \pi^0 \nu_\tau$

<sup>1</sup> From a fit to the Dalitz plot.<sup>2</sup> Assuming for  $f_0(1370)$  mass and width of 1186 and 350 MeV respectively.

$\Gamma(f_2(1270)\pi, f_2 \rightarrow \pi\pi)/\Gamma_{\text{total}}$   $\Gamma_9/\Gamma$

<u>VALUE</u> (units $10^{-2}$ )	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>• • •</b> We do not use the following data for averages, fits, limits, etc. <b>• • •</b>				
$1.19 \pm 0.49 \pm 0.17$	37k	<sup>1,2</sup> ASNER	00	CLE2 $e^+ e^- \rightarrow \tau^+ \tau^-, \tau^- \rightarrow \pi^- \pi^0 \pi^0 \nu_\tau$

<sup>1</sup> From a fit to the Dalitz plot.

<sup>2</sup> Assuming for  $f_2(1270)$  mass and width of 1275 and 185 MeV respectively.

$\Gamma(\pi^+ \pi^- \pi^0)/\Gamma_{\text{total}}$   $\Gamma_{10}/\Gamma$

<u>VALUE</u>	<u>DOCUMENT ID</u>	<u>COMMENT</u>
<b>seen</b>	BARBERIS 98B	$450 pp \rightarrow p_f \pi^+ \pi^- \pi^0 p_s$

$\Gamma(\pi^0 \pi^0 \pi^0)/\Gamma(\pi^+ \pi^- \pi^0)$   $\Gamma_{11}/\Gamma_{10}$

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>COMMENT</u>
<b>• • •</b> We do not use the following data for averages, fits, limits, etc. <b>• • •</b>			

<0.008 90 <sup>1</sup> BARBERIS 01  $450 pp \rightarrow p_f 3\pi^0 p_s$

<sup>1</sup> Inconsistent with observations of  $\sigma\pi$ ,  $f_0(1370)\pi$ , and  $f_2(1270)\pi$  decay modes.

$\Gamma(K^*(892)K)/\Gamma_{\text{total}}$   $\Gamma_{13}/\Gamma$

<u>VALUE</u> (units $10^{-2}$ )	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>• • •</b> We do not use the following data for averages, fits, limits, etc. <b>• • •</b>				
2.2 ± 0.5	2255	<sup>1</sup> COAN 04	CLEO	$\tau^- \rightarrow K^- \pi^- K^+ \nu_\tau$
8 to 15	205	<sup>2</sup> DRUTSKOY 02	BELL	$B \rightarrow D^{(*)} K^- K^{*0}$
$3.3 \pm 0.5 \pm 0.1$	37k	<sup>3</sup> ASNER 00	CLE2	$10.6 e^+ e^- \rightarrow \tau^+ \tau^-, \tau^- \rightarrow \pi^- \pi^0 \pi^0 \nu_\tau$
2.6 ± 0.3		<sup>4</sup> BARATE 99R	ALEP	$\tau \rightarrow K \bar{K} \pi \nu_\tau$

<sup>1</sup> Using structure functions from KUHN 92 and DECKER 93A and  $B(\tau^- \rightarrow K^- \pi^- K^+ \nu_\tau) = (0.155 \pm 0.006 \pm 0.009)\%$  from BRIERE 03.

<sup>2</sup> From a comparison to ALAM 94 assuming purely resonant production of the  $K^- K^{*0}$  system.

<sup>3</sup> From a fit to the  $3\pi$  mass spectrum including the  $K \bar{K}^*(892)$  threshold.

<sup>4</sup> Assuming  $a_1(1260)$  dominance and taking  $B(\tau \rightarrow a_1(1260) \nu_\tau)$  from BUSKULIC 96.

## $a_1(1260)$ REFERENCES

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ANDO	92	PL B291 496	A. Ando <i>et al.</i>	(KEK, KYOT, NIRS, SAGA+)
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IVANOV	91	ZPHY C49 563	Y.P. Ivanov, A.A. Osipov, M.K. Volkov	(JINR)
ARMSTRONG	90	ZPHY C48 213	T.A. Armstrong, M. Benayoun, W. Beusch	(WA76 Coll.)
FEINDT	90	ZPHY C48 681	M. Feindt	(HAMB)
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BAND	87	PL B198 297	H.R. Band <i>et al.</i>	(MAC Collab.)
TORNQVIST	87	ZPHY C36 695	N.A. Tornqvist	(HELS)
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DAUM	81B	NP B182 269	C. Daum <i>et al.</i>	(AMST, CERN, CRAC, MPIM+)
DAUM	80	PL 89B 281	C. Daum <i>et al.</i>	(AMST, CERN, CRAC, MPIM+) JP
GAVILLET	77	PL 69B 119	P. Gavillet <i>et al.</i>	(AMST, CERN, NIJM+) JP
BOWLER	75	NP B97 227	M.G. Bowler <i>et al.</i>	(OXFTP, DARE)

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